

Forbush decreases observed with GRAPES-3

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Outline

- (Quick intro to) Coronal Mass Ejections (CMEs) from the Sun



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- (Quick intro to) Forbush Decreases (FDs)

The logo of IISER Pune is a large, stylized, light red graphic. It features a central grey circle, possibly representing the sun, enclosed within a square frame that has a wavy, flame-like shape extending from its top-left corner. Below the graphic, the text "IISER PUNE" is written in a bold, grey, sans-serif font.

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- (Quick intro to) Coronal Mass Ejections (CMEs) from the Sun
- (Quick intro to) Forbush Decreases (FDs)
- Main Q we try to answer: What are Forbush decreases (predominantly) due to: (near-Earth manifestations of) CMEs or their associated shocks?

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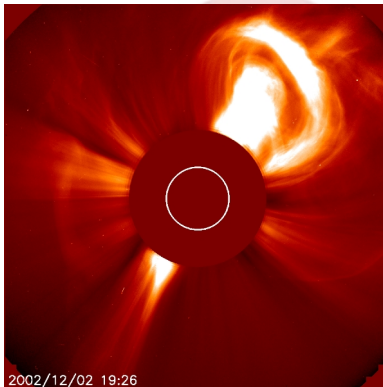
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- (Quick intro to) Forbush Decreases (FDs)
- Main Q we try to answer: What are Forbush decreases (predominantly) due to: (near-Earth manifestations of) CMEs or their associated shocks?
- This work is a followup to Subramanian et al 2009, and is part of Arun Babu's PhD thesis

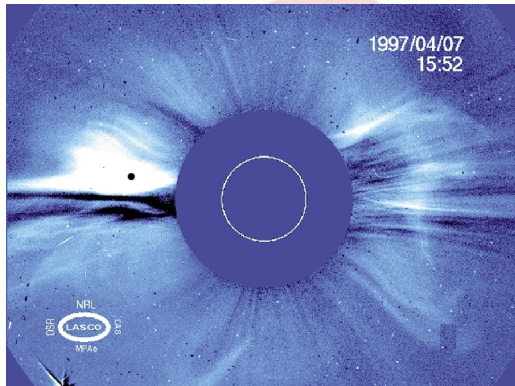
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CME



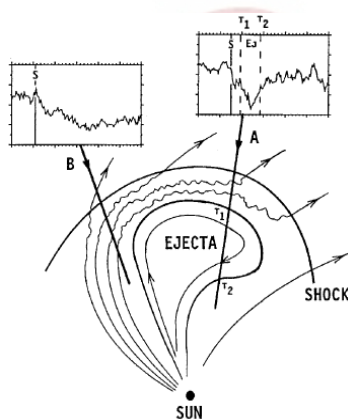
Plane-of-sky CME (from LASCO/SOHO)

CME



Halo (Earth-directed) CME (from LASCO/SOHO)

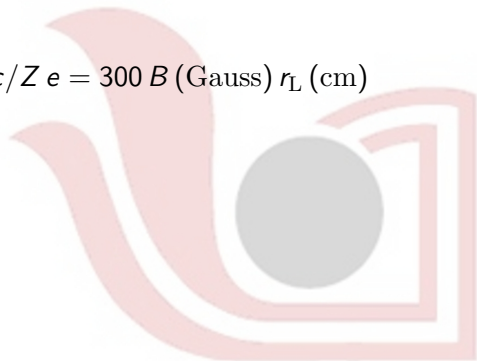
Forbush Decrease



Forbush decreases (from Cane 2000)

Cutoff Rigidity

- Rigidity R_g (volts) $\equiv P c / Z e = 300 B$ (Gauss) r_L (cm)



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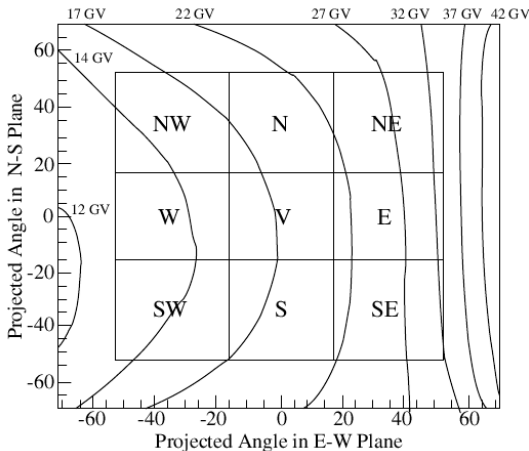
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Cutoff Rigidity

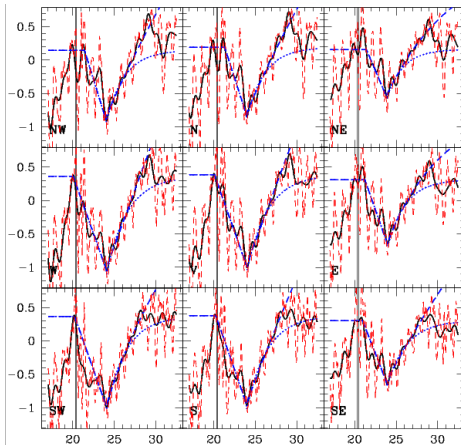
- Rigidity R_g (volts) $\equiv P c / Z e = 300 B$ (Gauss) r_L (cm)
- Protons below the cutoff rigidity don't make it to the top of the atmosphere (to produce a shower); they are deflected by the geomagnetic field back into space.
- The cutoff rigidity is very dependent on the B field geometry;
→ 0 for a nearly vertical field,
→ ∞ for nearly horizontal field;
i.e., its dependent on the viewing direction (different for East, West, North, South)

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9 directional bins make GRAPES-3 a multi-rigidity instrument

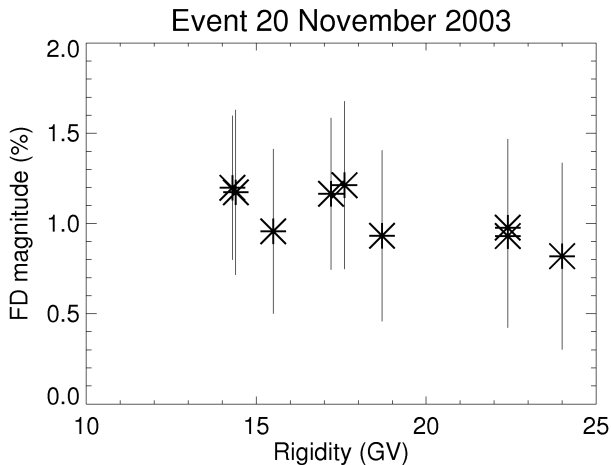


FD on 20 Nov 2003



Forbush decrease on Nov 20 2003 observed with GRAPES-3

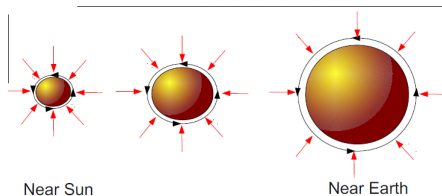
FD on 20 Nov 2003



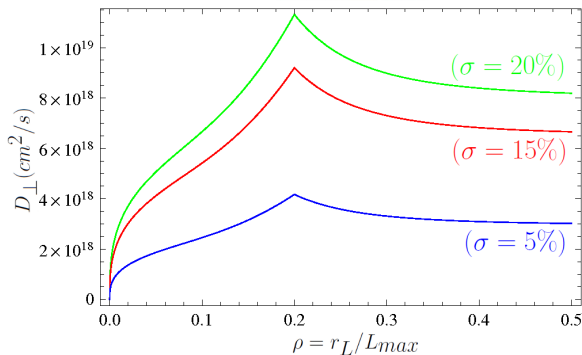
Forbush decrease on Nov 20 2003 observed with GRAPES-3



The CME-only model (following Subramanian et al 2009)



High energy galactic (*not of solar origin*) cosmic rays progressively diffuse into the expanding, propagating CME bubble, across the B fields bounding it. At the earth, the density contrast between the CME interior & outside (in high energy CR protons) is manifested as the Forbush decrease. The diffusion coefficient D_{\perp} depends upon the CR proton rigidity Rg and the turbulence level σ^2 near the CME. D_{\perp} is taken from analytical fits to Monte Carlo simulations of charged particle diffusion in turbulent B fields (Candia & Roulet 2004).

$D_{\perp}(Rg, \sigma^2)$: Candia & Roulet (2004)

$\rho \equiv r_L/L_{max}$ (similar to proton rigidity). $L_{max} \equiv$ largest lengthscale in the problem. r_L is proton gyroradius. Turbulence level near CME $\sigma^2 \equiv \langle B_{turb}^2/B_0^2 \rangle$

The CME-only model: details

Flux of protons entering CME is

$$F (\text{cm}^{-2} \text{s}^{-1}) = D_{\perp}(\rho, \sigma^2) \frac{\partial N_a}{\partial r}$$

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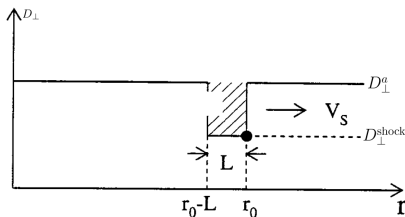
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- Turbulence level near CME $\sigma^2 \equiv \langle B_{\text{turb}}^2 / B_0^2 \rangle$.

The shock-only model (following Wibberenz et al 1998)



The Forbush decrease is due to the shock - a propagating, diffusive barrier. The B field enhancement at the shock acts as an “umbrella” against galactic cosmic rays.

$$FD = \frac{V_s L}{D_{\perp}^a} \left(\frac{D_{\perp}^a}{D_{\perp}^{\text{shock}}} - 1 \right)$$

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- D_{\perp} is given by Candia & Roulet's (2004) formalism, as before.

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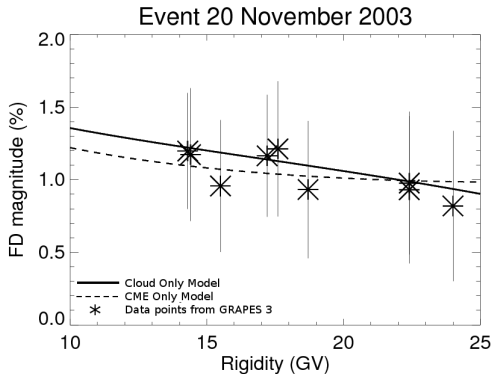
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- Shock "thickness" L , velocity V_s , B fields inside and outside the shock estimated from in-situ spacecraft obs (OMNI database)

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- Turbulence level σ_{shock}^2 taken to be twice $\sigma_{\text{ambient}}^2$

Results - I

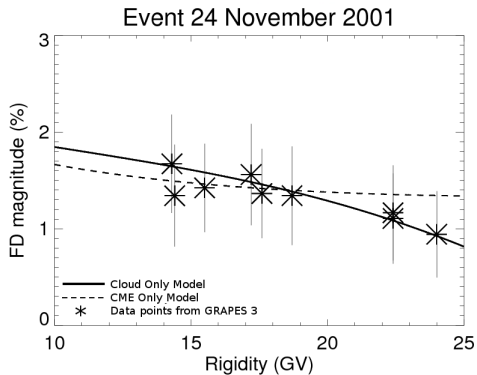


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For CME-only model, $\sigma = 5\%$,
 while for shock-only model $\sigma_{\text{ambient}} = 25\%$.

Typical quiet sun turbulence level σ (at 15 – 50 R_{\odot}) $\approx 6\text{--}15\%$ (Spangler 2002).

Results - II



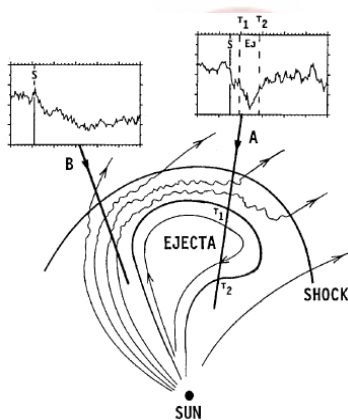
For CME-only model, $\sigma = 8\%$,
 while for shock-only model $\sigma_{\text{ambient}} = 100\%$!
 Typical quiet sun turbulence level σ (at $15 - 50 R_{\odot}$) $\approx 6-15\%$ (Spangler 2002).

Conclusions

We have fitted CME-only and shock-only models to **multi-rigidity data** of FDs for the first time. We use a couple of representative, well observed events, and use as many observed parameters as possible. Our main fitting parameter is the turbulence level σ . On this basis, we find that the low density (CR) bubble/CME is the major contributor to the observed FD, as opposed to the shock/umbrella driven by it.

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