

# Searching for millicharged particles at the LHC

Steven Lowette

Vrije Universiteit Brussel – IIHE

TIFR DHEP seminar  
20 May 2021



# Towards discovery at LHC

## Run 1: 2010-2012

- 7-8 TeV, 25/fb
- long-standing targets: Higgs boson,  $B_s \rightarrow \mu\mu$ , QGP, ...
- many limits on SUSY, ED, 4<sup>th</sup> gen, etc

# Towards discovery at LHC

## Run 1: 2010-2012

- 7-8 TeV, 25/fb
- long-standing targets: Higgs boson,  $B_s \rightarrow \mu\mu$ , QGP, ...
- many limits on SUSY, ED, 4<sup>th</sup> gen, etc

## Run 2: 2015-2018

- 13 TeV, 140/fb
- hard / rare processes:  $t\bar{t} \rightarrow H/H \rightarrow b\bar{b}, \mu\mu$ , B anomalies,  $\cancel{P}$  in charm, ... [ $\gamma\gamma(750)$ ...]
- many more limits on BSM physics

# Towards discovery at LHC

## Run 1: 2010-2012

- 7-8 TeV, 25/fb
- long-standing targets: Higgs boson,  $B_s \rightarrow \mu\mu$ , QGP, ...
- many limits on SUSY, ED, 4<sup>th</sup> gen, etc

## Run 2: 2015-2018

- 13 TeV, 140/fb
- hard / rare processes:  $t\bar{t} \rightarrow H/H \rightarrow b\bar{b}, \mu\mu$ , B anomalies,  $\phi$  in charm, ... [ $\gamma\gamma(750)$ ...]
- many more limits on BSM physics

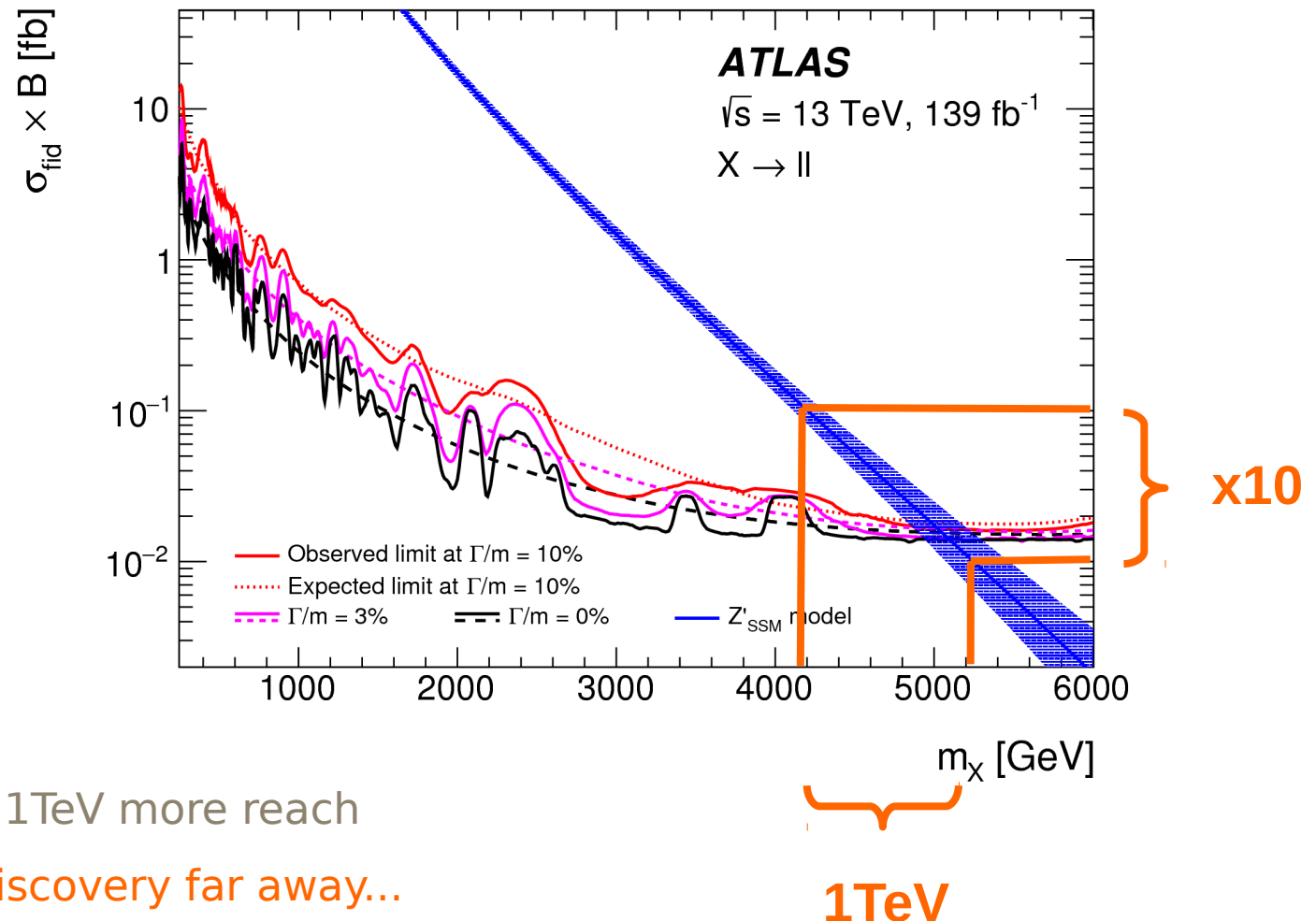
## Run 3: 2022-2024

- 13-14 TeV, 300/fb
- aim for lower mass, lower cross section, difficult final states (eg. LL)
- BSM sensitivity? more lumi needed!

## Much more lumi - **but how much?**

- eg. new physics bump hunt on Drell-Yan background

arXiv:1903.06248 [hep-ex]



- factor 10 in lumi  $\rightarrow$  1TeV more reach
  - no hint yet  $\rightarrow$  **discovery far away...**

# So what else?

## The LHC is unique

- only player at the energy frontier
  - since a decade, more to come
- only player at the intensity frontier
  - at the EW scale
- whatever LHC is sensitive to should be done now or “never”
  - maximize return on investment
  - small investment can make big difference
- what else can we do with the LHC?
  - how can new physics still be hidden?

Page 7

Search or Article

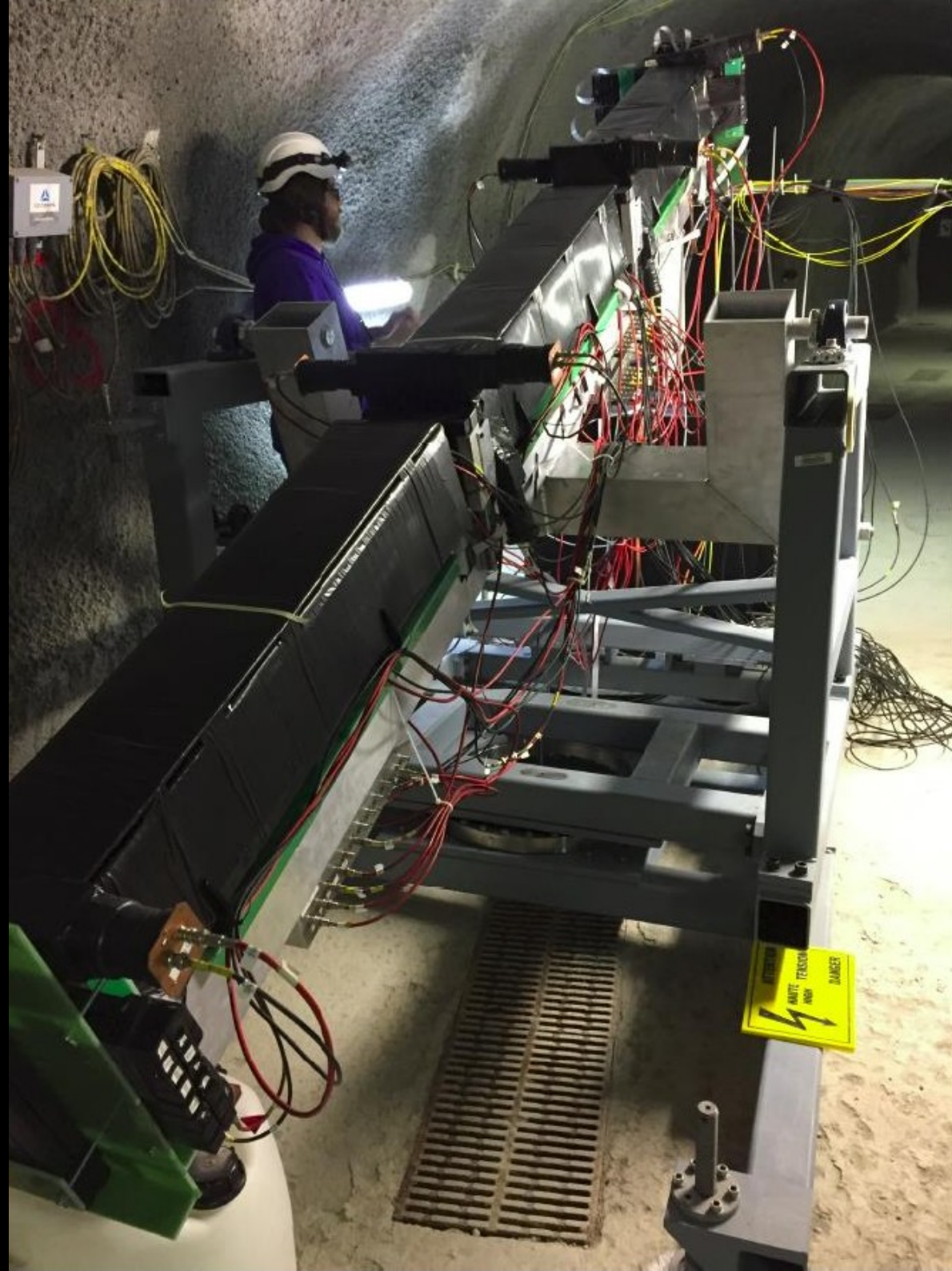
[\(Help | Advanced search\)](#)

Page 8



# milliQan

**New experiment  
to search for  
millicharged  
particles  
at the LHC**



## Electric charge quantization

- all our elementary particles have an electric charge of  $N \cdot e/3$

Q U A R K S	<div><b>UP</b> mass 2,3 MeV/c<sup>2</sup> charge <math>\frac{2}{3}</math> spin <math>\frac{1}{2}</math> </div>	<div><b>CHARM</b> 1,275 GeV/c<sup>2</sup> <math>\frac{2}{3}</math> <math>\frac{1}{2}</math> </div>	<div><b>TOP</b> 173,07 GeV/c<sup>2</sup> <math>\frac{2}{3}</math> <math>\frac{1}{2}</math> </div>	<div><b>GLUON</b> 0 0 1 </div>	<div><b>HIGGS BOSON</b> 126 GeV/c<sup>2</sup> 0 0 </div>	
	<div><b>DOWN</b> 4,8 MeV/c<sup>2</sup> <math>-\frac{1}{3}</math> <math>\frac{1}{2}</math> </div>	<div><b>STRANGE</b> 95 MeV/c<sup>2</sup> <math>-\frac{1}{3}</math> <math>\frac{1}{2}</math> </div>	<div><b>BOTTOM</b> 4,18 GeV/c<sup>2</sup> <math>-\frac{1}{3}</math> <math>\frac{1}{2}</math> </div>	<div><b>PHOTON</b> 0 0 1 </div>	G A U G E B O S O N S	
				<div><b>Z BOSON</b> 91,2 GeV/c<sup>2</sup> 0 1 </div>		
				<div><b>W BOSON</b> 80,4 GeV/c<sup>2</sup> <math>\pm 1</math> 1 </div>		
L E P T O N S	<div><b>ELECTRON</b> 0,511 MeV/c<sup>2</sup> -1 <math>\frac{1}{2}</math> </div>	<div><b>MUON</b> 105,7 MeV/c<sup>2</sup> -1 <math>\frac{1}{2}</math> </div>	<div><b>TAU</b> 1,777 GeV/c<sup>2</sup> -1 <math>\frac{1}{2}</math> </div>			
	<div><b>ELECTRON NEUTRINO</b> &lt;2,2 eV/c<sup>2</sup> 0 <math>\frac{1}{2}</math> </div>	<div><b>MUON NEUTRINO</b> &lt;0,17 MeV/c<sup>2</sup> 0 <math>\frac{1}{2}</math> </div>	<div><b>TAU NEUTRINO</b> &lt;15,5 MeV/c<sup>2</sup> 0 <math>\frac{1}{2}</math> </div>			

- but... we don't know why electric charge appears quantized
  - explanations so far involve magnetic monopoles (never found) or SU(5) grand unified theories (ruled out)
  - it is theoretically entirely consistent to add new particles with small electric charge

## So how do millicharges then work?

- millicharged particles may actually arise rather naturally in extensions of the Standard Model
- suppose we add a  $U(1)'$  massless boson to the SM, a **dark photon**

$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{4} B'_{\mu\nu} B'^{\mu\nu} - \frac{\kappa}{2} B'_{\mu\nu} B^{\mu\nu}$$

kinetic mixing  
[Holdom '86]



- the kinetic mixing term can be generated through new heavy particles that couple both to hypercharge and to new  $U(1)'$



The diagram shows a fermion loop (circle with an arrow) with a photon (gamma) and a U(1)' gauge boson (A') attached. The fermion is labeled psi.

$$\sim \frac{e g_D}{16\pi^2} \log \frac{m_\psi}{M_*}$$

- generates **coupling  $10^{-3}$**  for  $m_\psi \sim \text{EW scale}$
- let's now add a **new fermion only charged under  $U(1)'$**

$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{4} B'_{\mu\nu} B^{\mu\nu'} - \frac{\kappa}{2} B'_{\mu\nu} B^{\mu\nu} + i\bar{\psi}(\not{\partial} + ie'B' + iM_{\text{mCP}})\psi$$

- and redefine the field

$$B' \rightarrow B' + \kappa B$$

- mixing term disappears and new fermion gets hypercharge
- after EWSB **new fermion has arbitrary electric charge**:  $Q = \kappa e' \cos \theta_W$

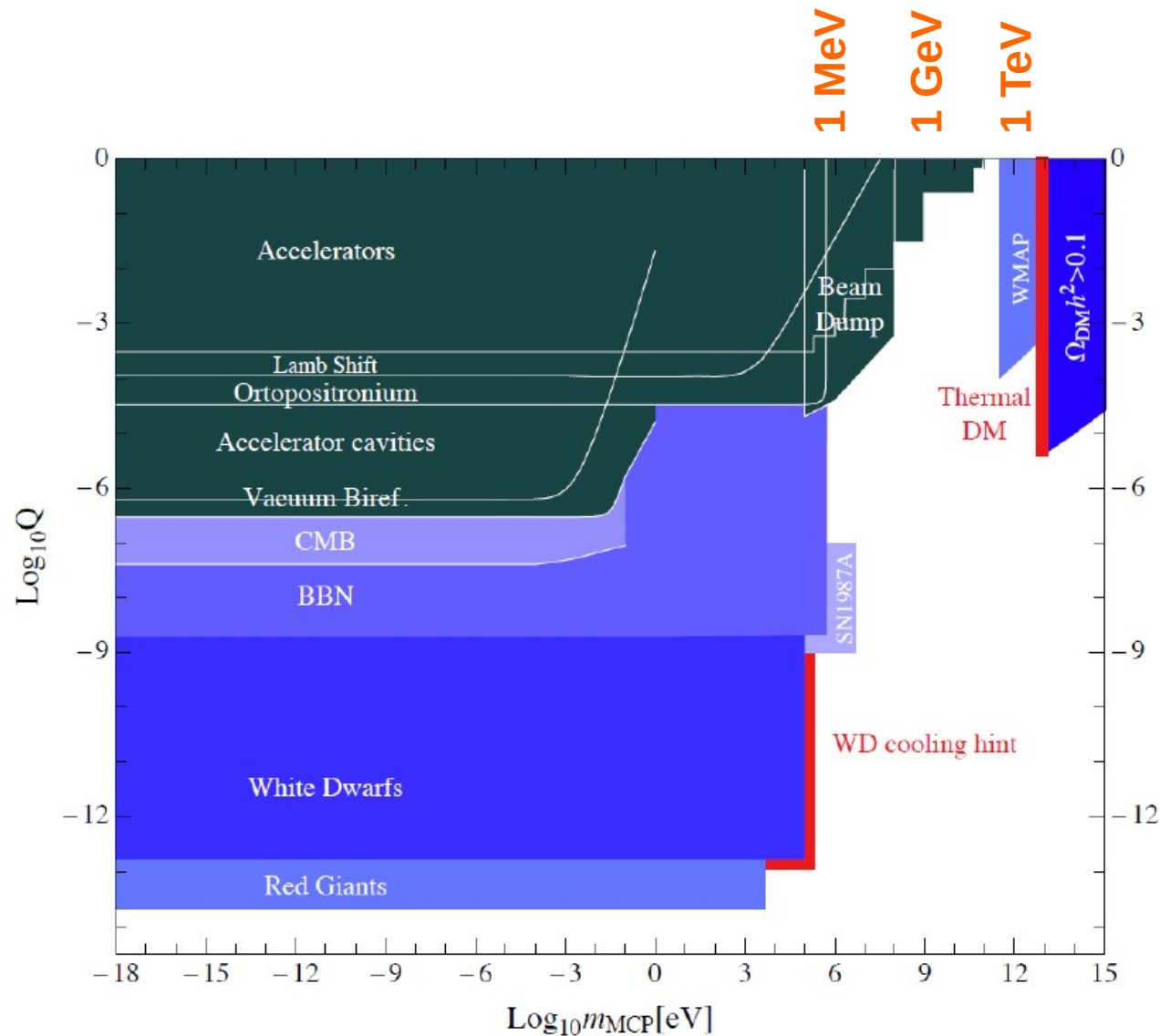
## Theoretical

- kinetic mixing (a.k.a. vector portal) is 1 of 4 possible portals to **dark sector**
  - others go via Higgs, neutrinos, or axions
- dark sector can contain **dark matter**!
- extra U(1) groups rather generic feature of GUTs and other theories
- very economical BSM extension, easy to embed

## Experimental

- **opportunity**: very weak limits above  $\sim 1\text{GeV}$ ; unique role LHC
- EDGES 21cm result can be explained with subdominant millicharged DM (arXiv:1803.03091)

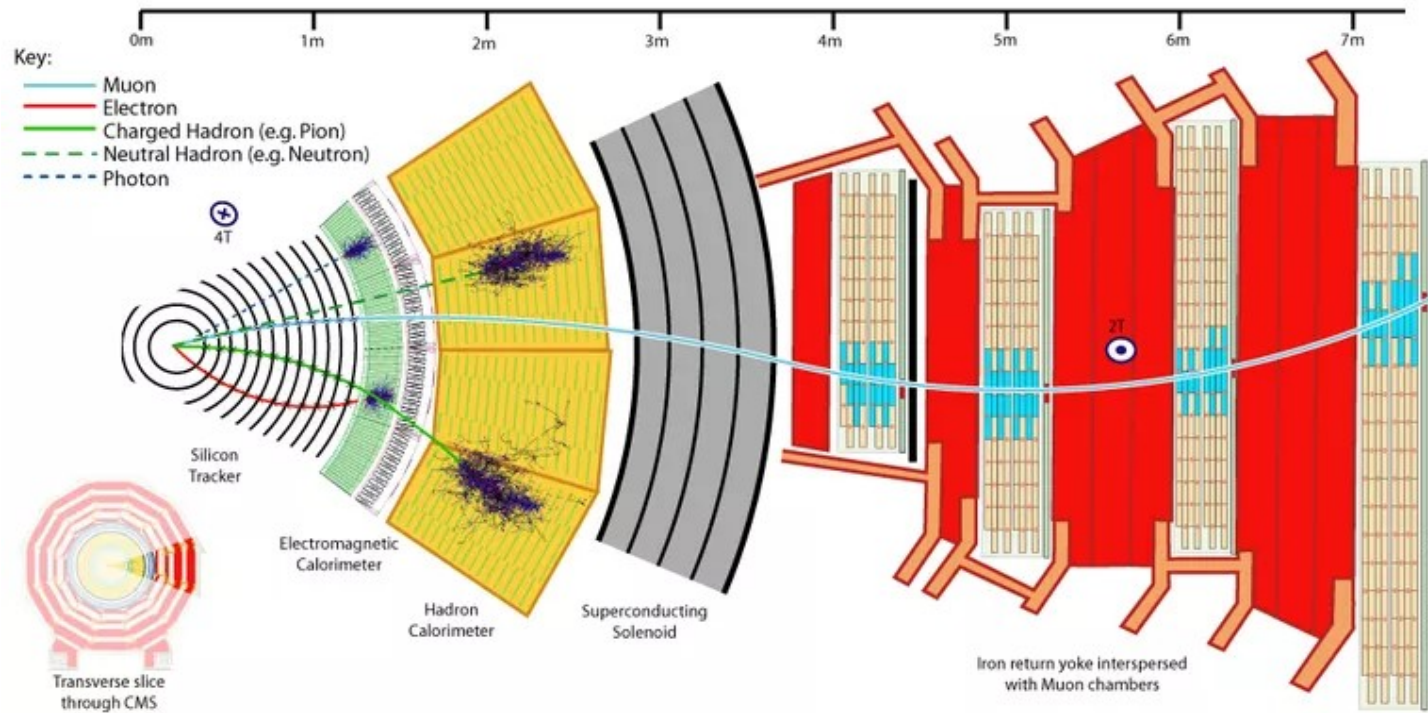
# Existing bounds





# How to detect millicharges?

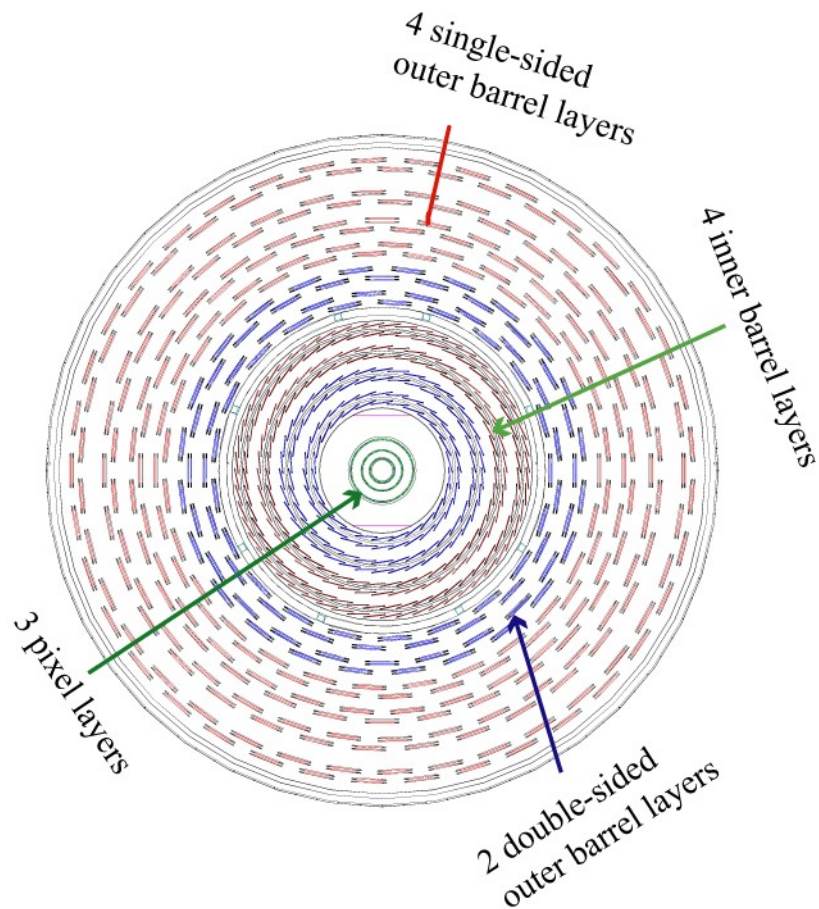
Just use CMS!?



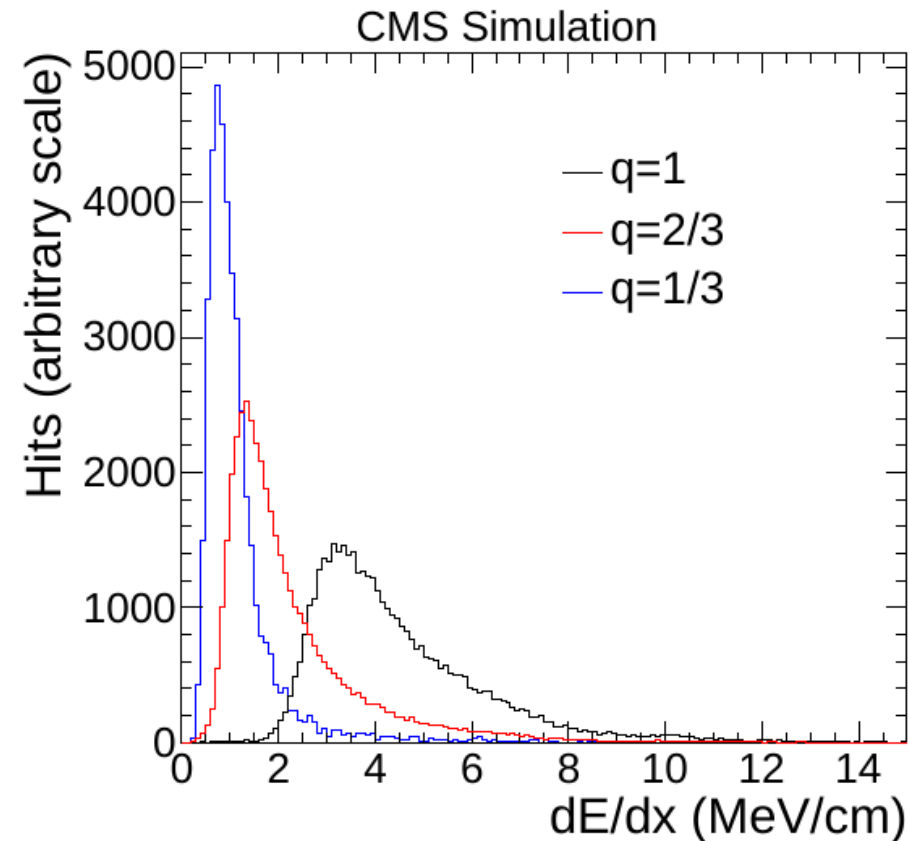
- at lower charge
  - tracks become straighter → momentum  $\sim 1/Q$
  - energy loss becomes smaller →  $dE/dx \sim Q^2$

## Search for tracks with many low-dE/dx hits

- 1 track brings many independent measurements of ionization loss



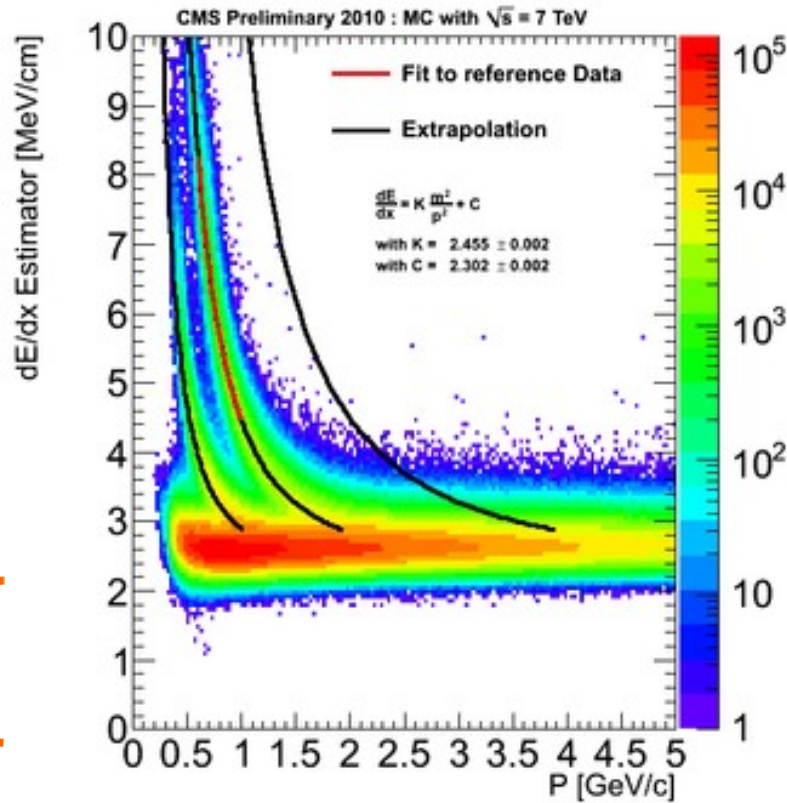
CMS-PAS-EXO-11-074



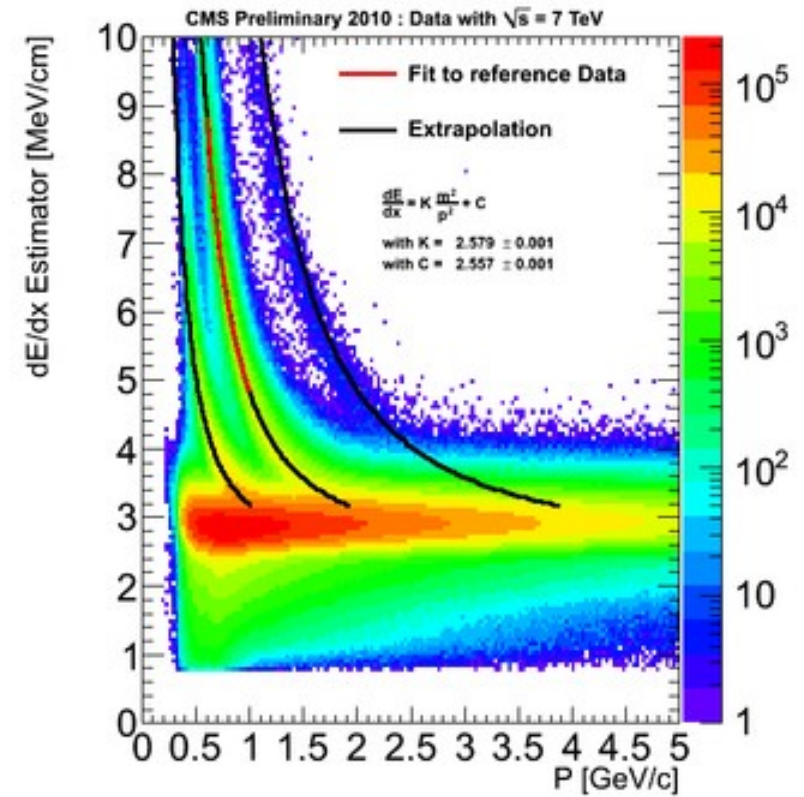


## dE/dx performance in CMS tracker

area of  
interest:  
difficult!



simulation

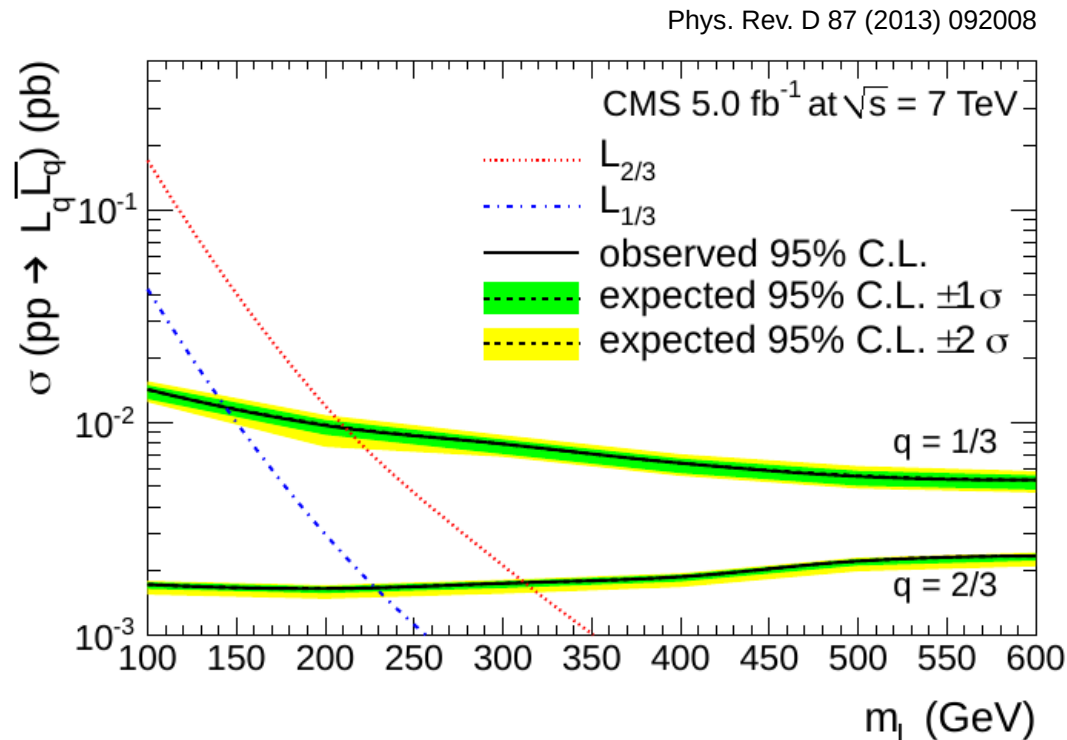


data

- also very sensitive to detector aging

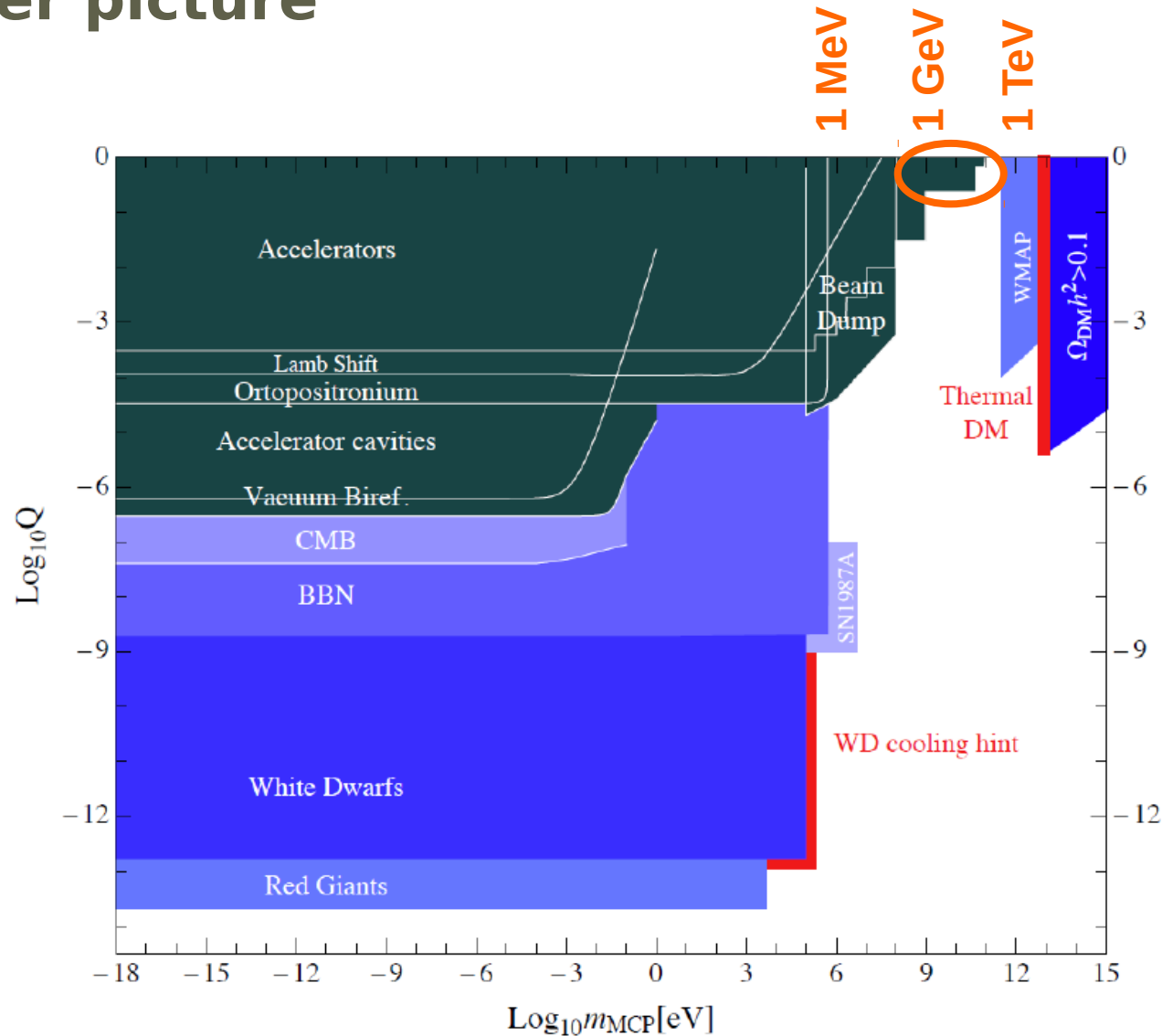
## Current analysis reach

- predict background fully from data
- can suppress all backgrounds above  $\sim 6$  hits with low  $dE/dx$



- LHC Run2 update almost ready!

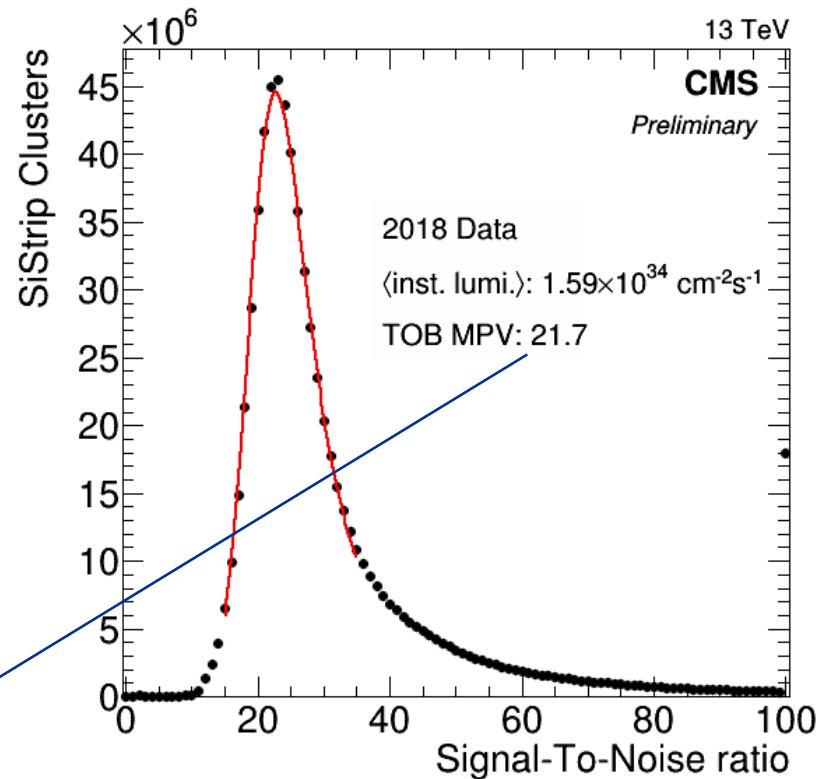
## The bigger picture



# Going lower

## What about lower charges?

- fundamental limitation of the CMS/ATLAS detectors



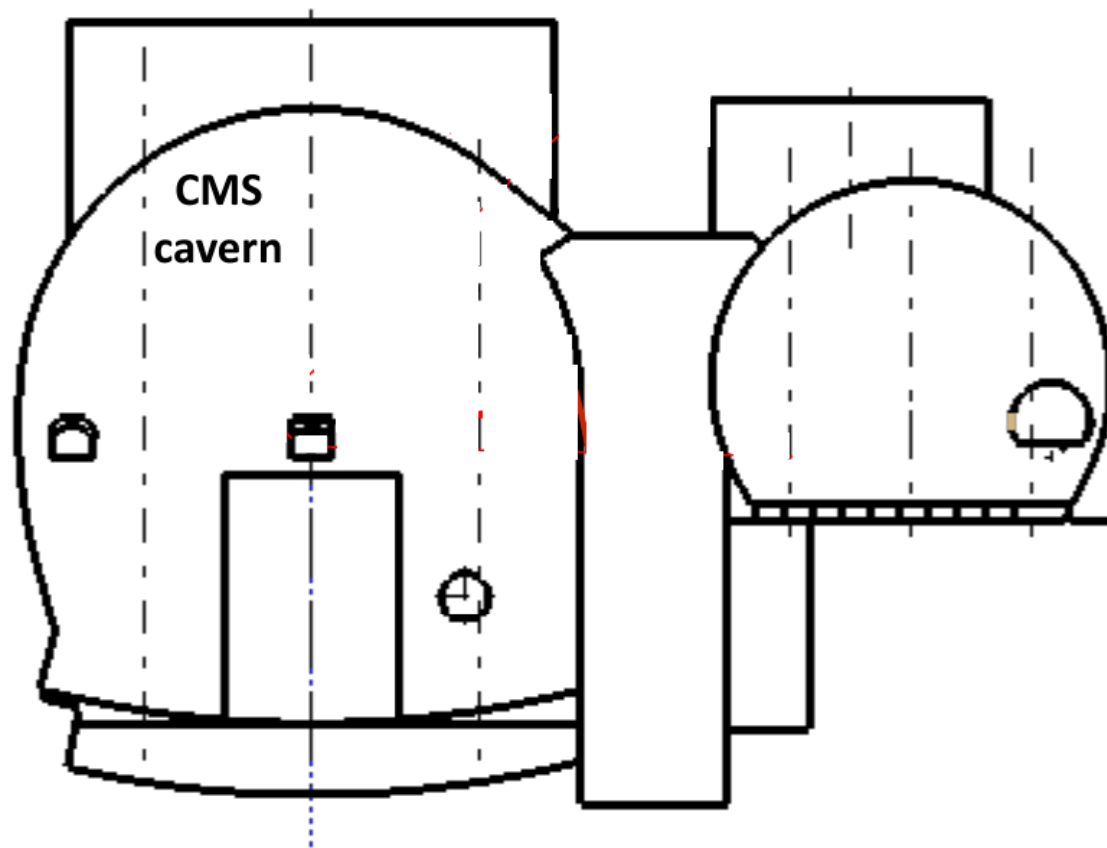
CMS DP-2018/052

- $1/Q^2 \sim 21.7$ 
  - cluster charge from particle with  $Q = 0.2$  has MPV  $\sim$  noise level

## Facing the $Q^2$ suppression

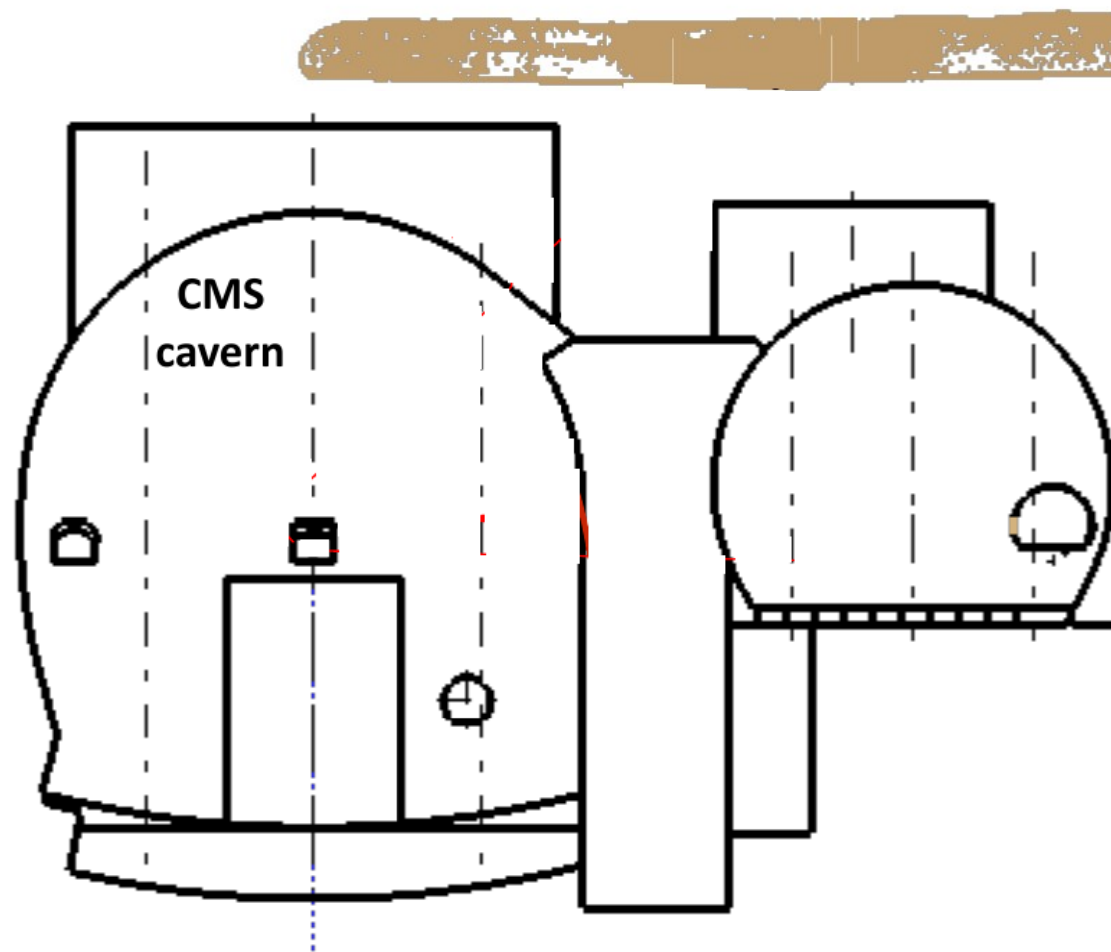
- need much more sensitive detection technique
  - with charge down to  $10^{-3}$ ,  $dE/dx$  suppressed by  $10^{-6}$
  - counting of single photons in “large” scintillator volume
- need to go to a low-background area
  - out of the CMS cavern, to suppress radiation backgrounds
  - still shielded from cosmic muons by  $\sim 100\text{m}$  overburden
- stay relatively close to the interaction point
  - minimize  $r^2$  suppression of flux

# milliQan location

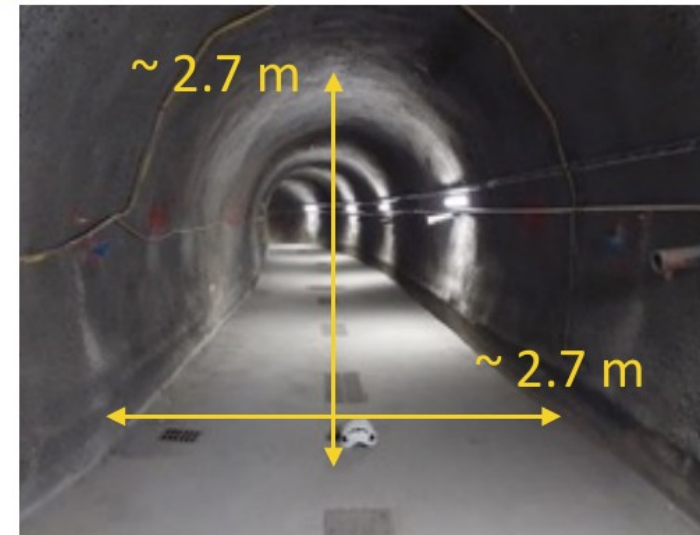


100m  
underground

# milliQan location

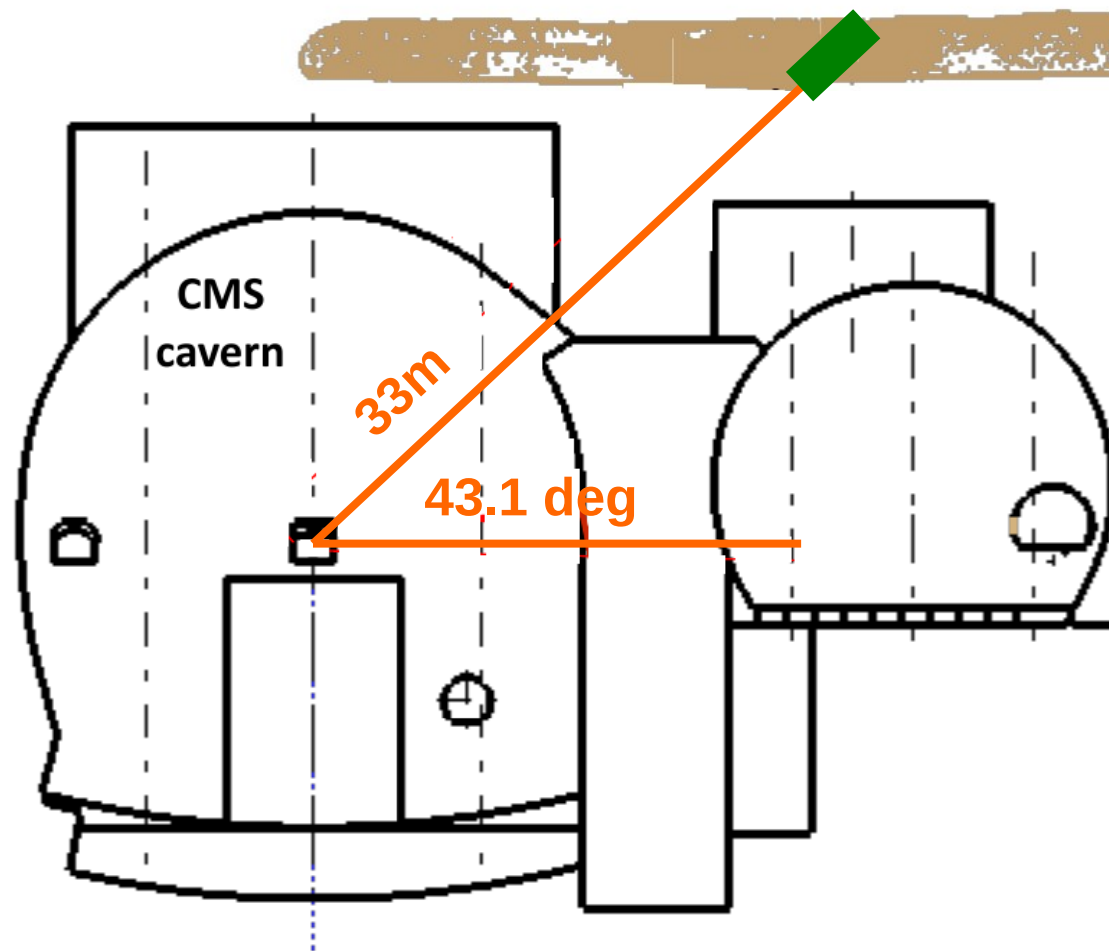


Drainage gallery

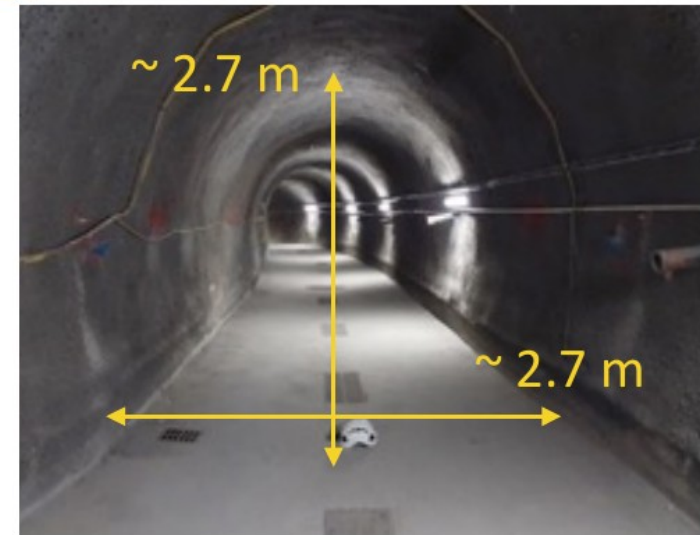


100m  
underground

# milliQan location



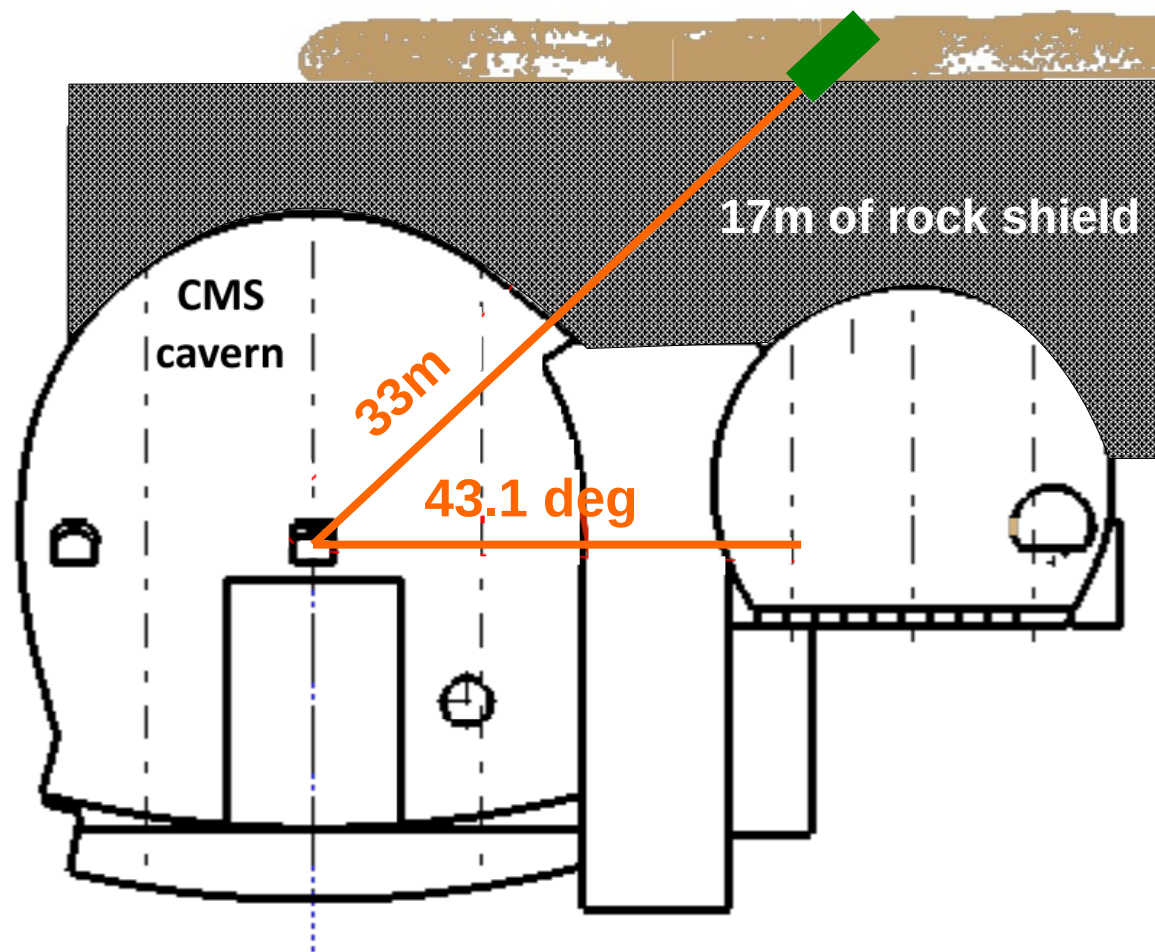
Drainage gallery



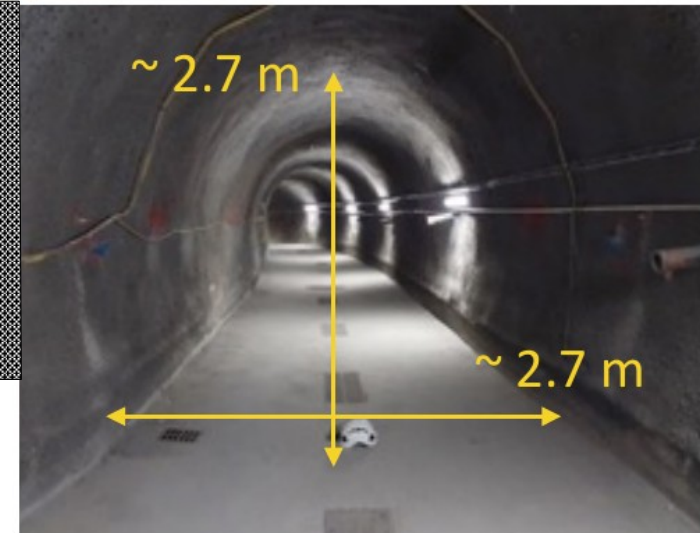
100m  
underground



# milliQan location

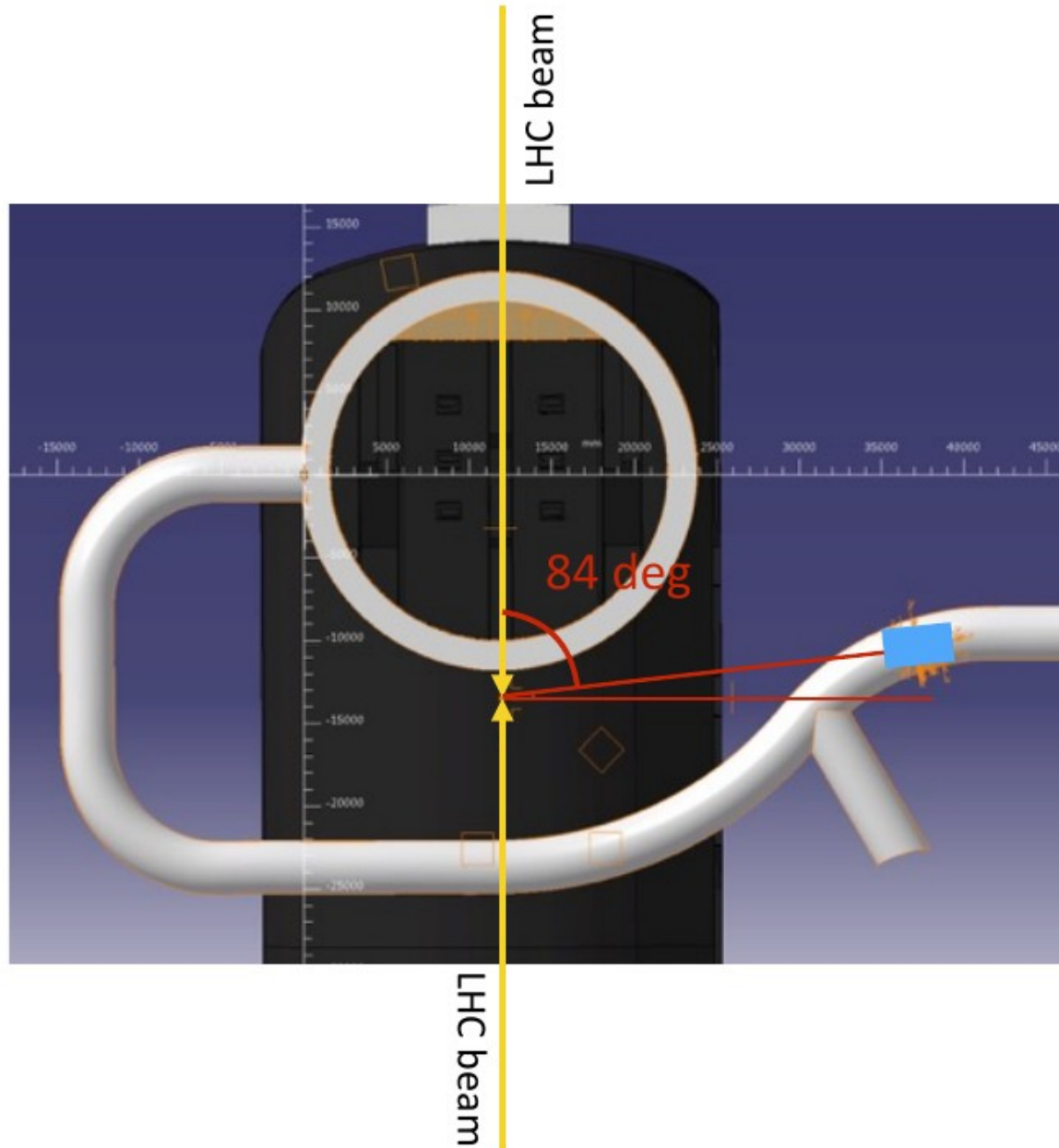


Drainage gallery



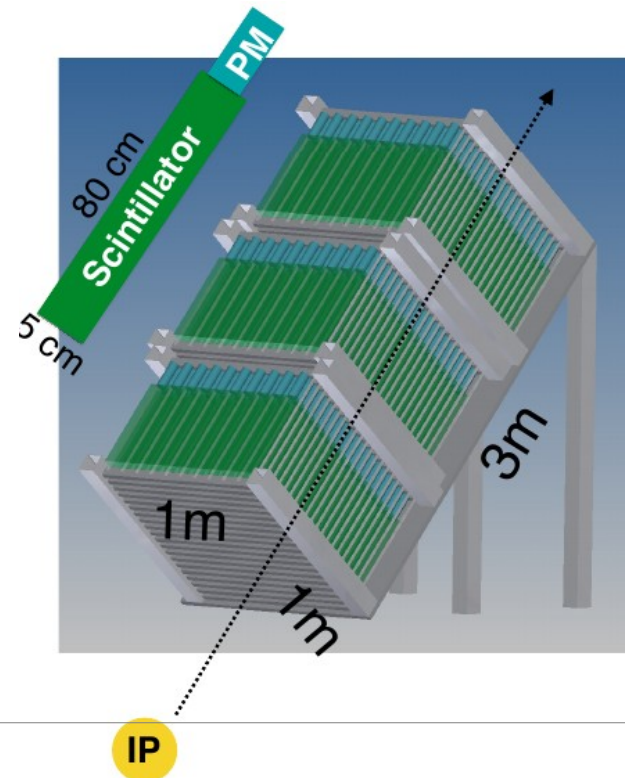
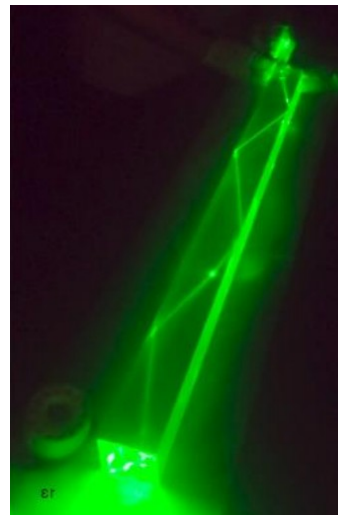
100m  
underground

# milliQan location



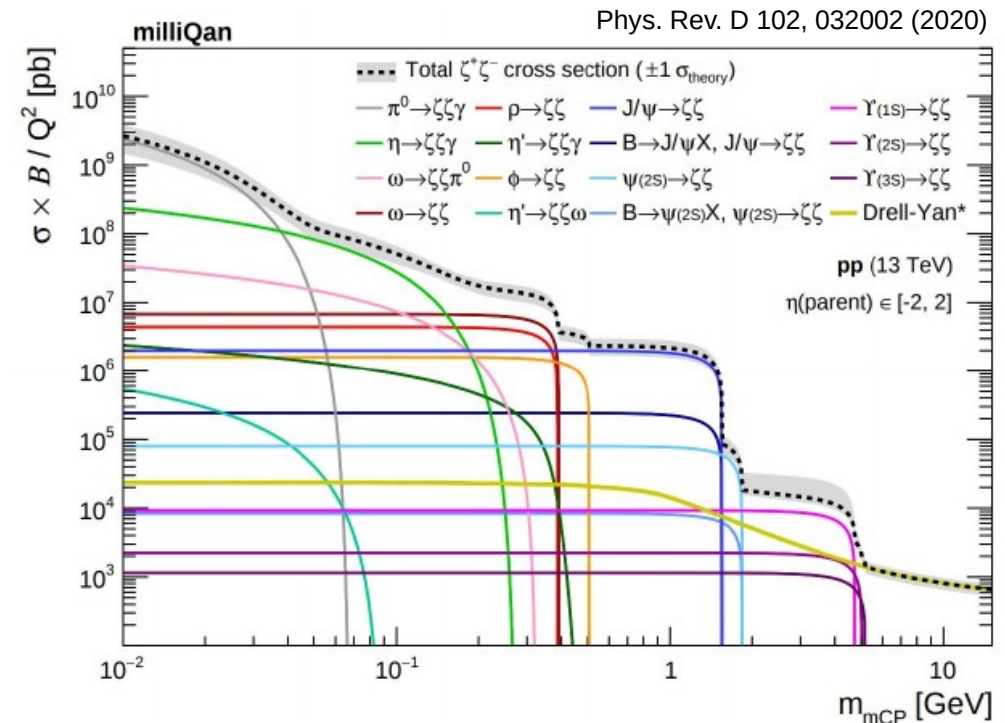
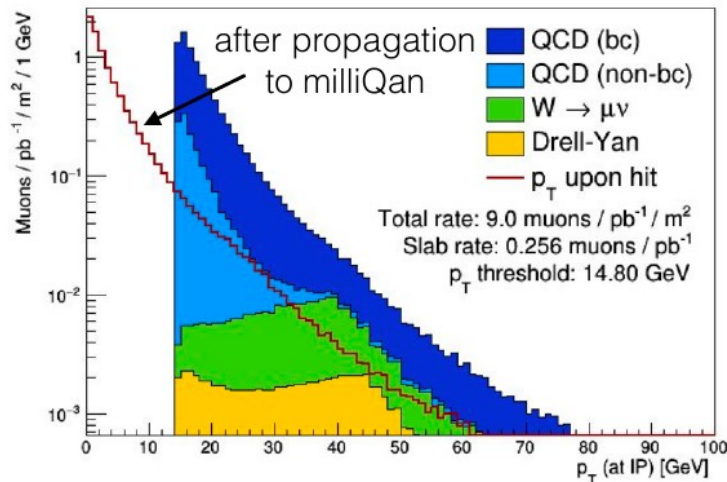
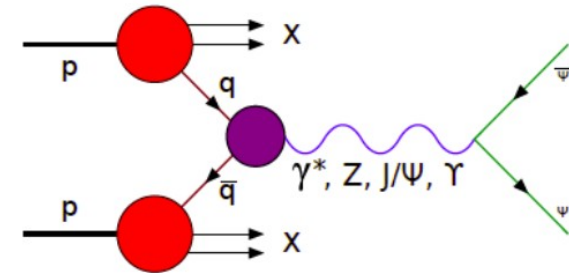
# milliQan detector principle

- concept: [arXiv:1410.6816](https://arxiv.org/abs/1410.6816); LOI: [arXiv:1607.04669](https://arxiv.org/abs/1607.04669)
- basic element is  $5 \times 5 \times 80$  cm<sup>3</sup> plastic scintillator
- attached to photomultiplier tube
- $1 \times 1 \times 3$  m<sup>3</sup> in 3 length-layers
- search coincidence of few photons in consecutive scintillators pointing to IP



## Production and transport simulation

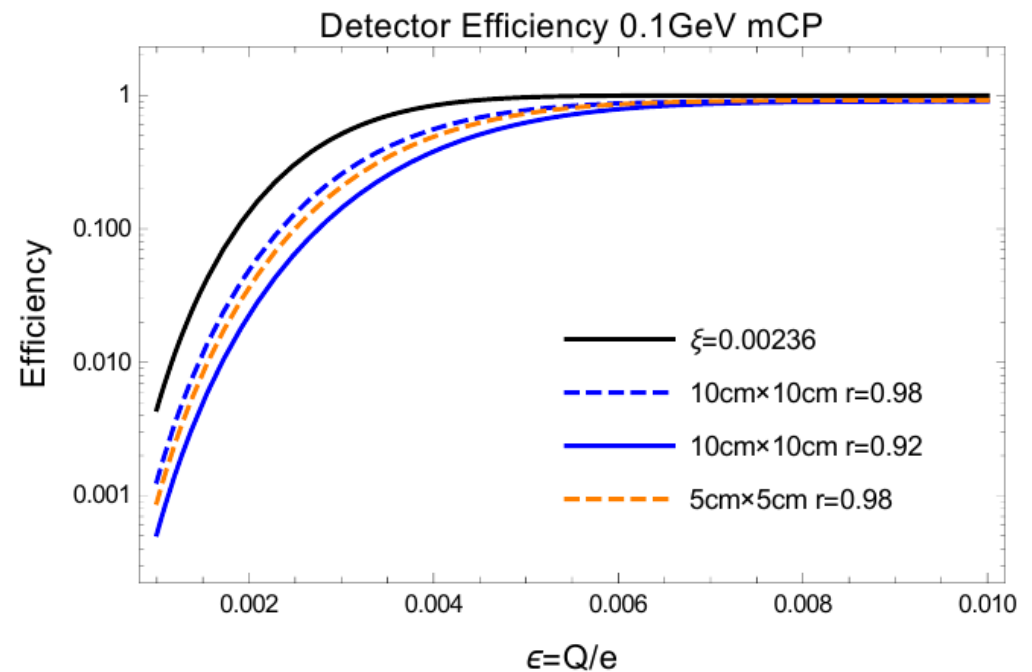
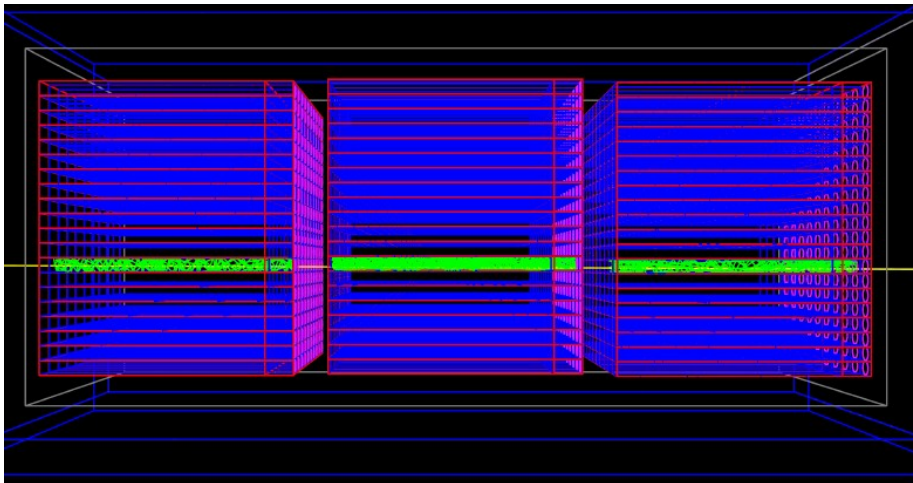
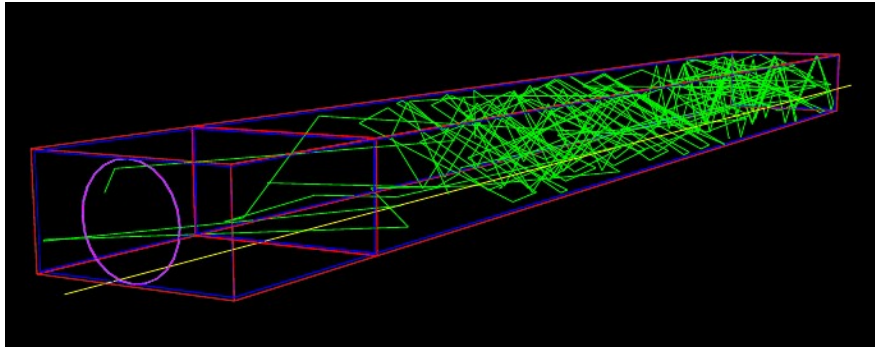
- any process that can make  $e^+e^-$  can make millicharged particles
  - low masses dominated by QCD production of  $\pi^0$ ,  $\eta$ ,  $\rho$ ,  $\omega$ ,  $\phi$ , then  $J/\psi$  and  $\Upsilon$
  - above 5GeV it's all  $Z/\gamma^*$
- propagate through CMS material and 17m of rock
  - with multiple scattering and CMS magnetic field





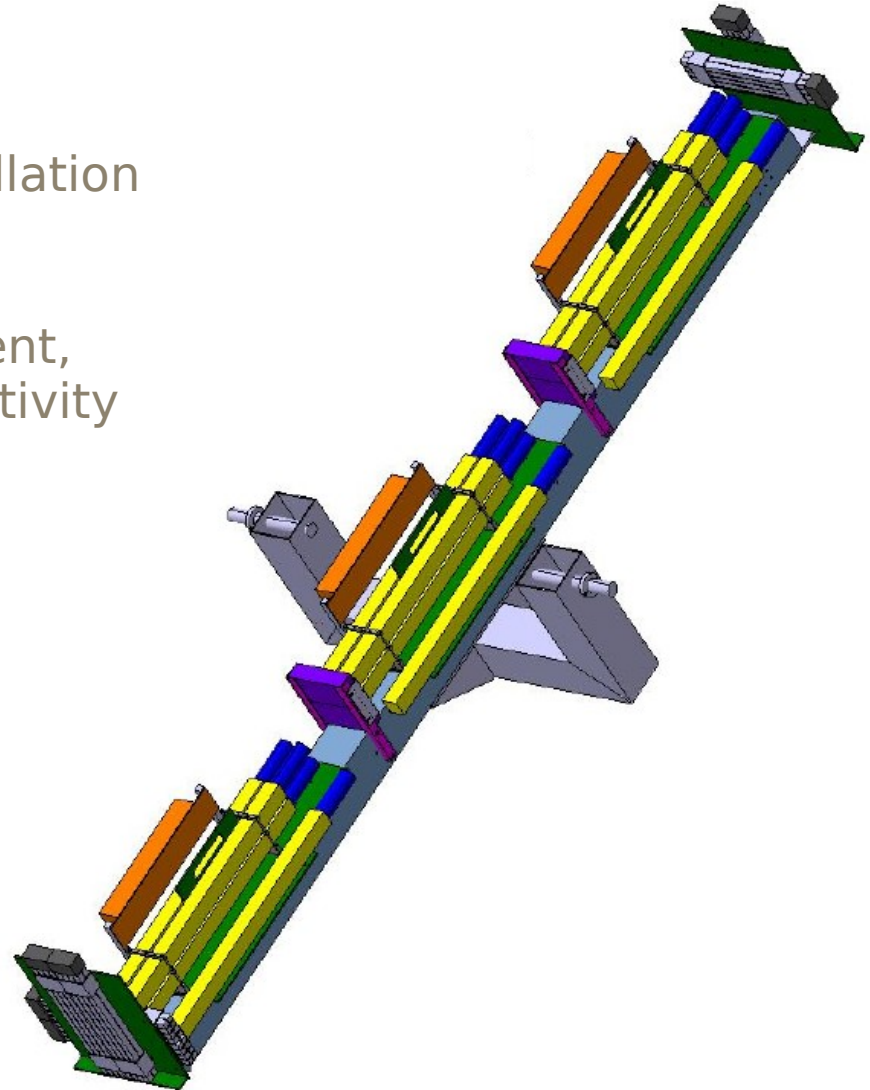
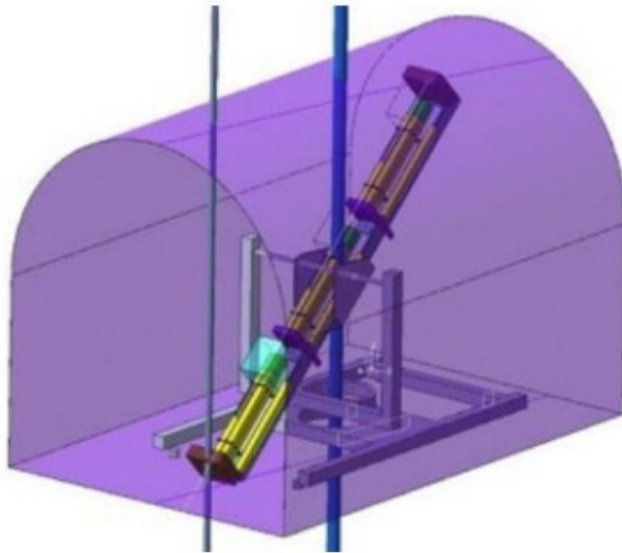
## Full Geant4 detector simulation

- models reflectivity, the light attenuation length, and the shape of the scintillator
- PMT parameters as input

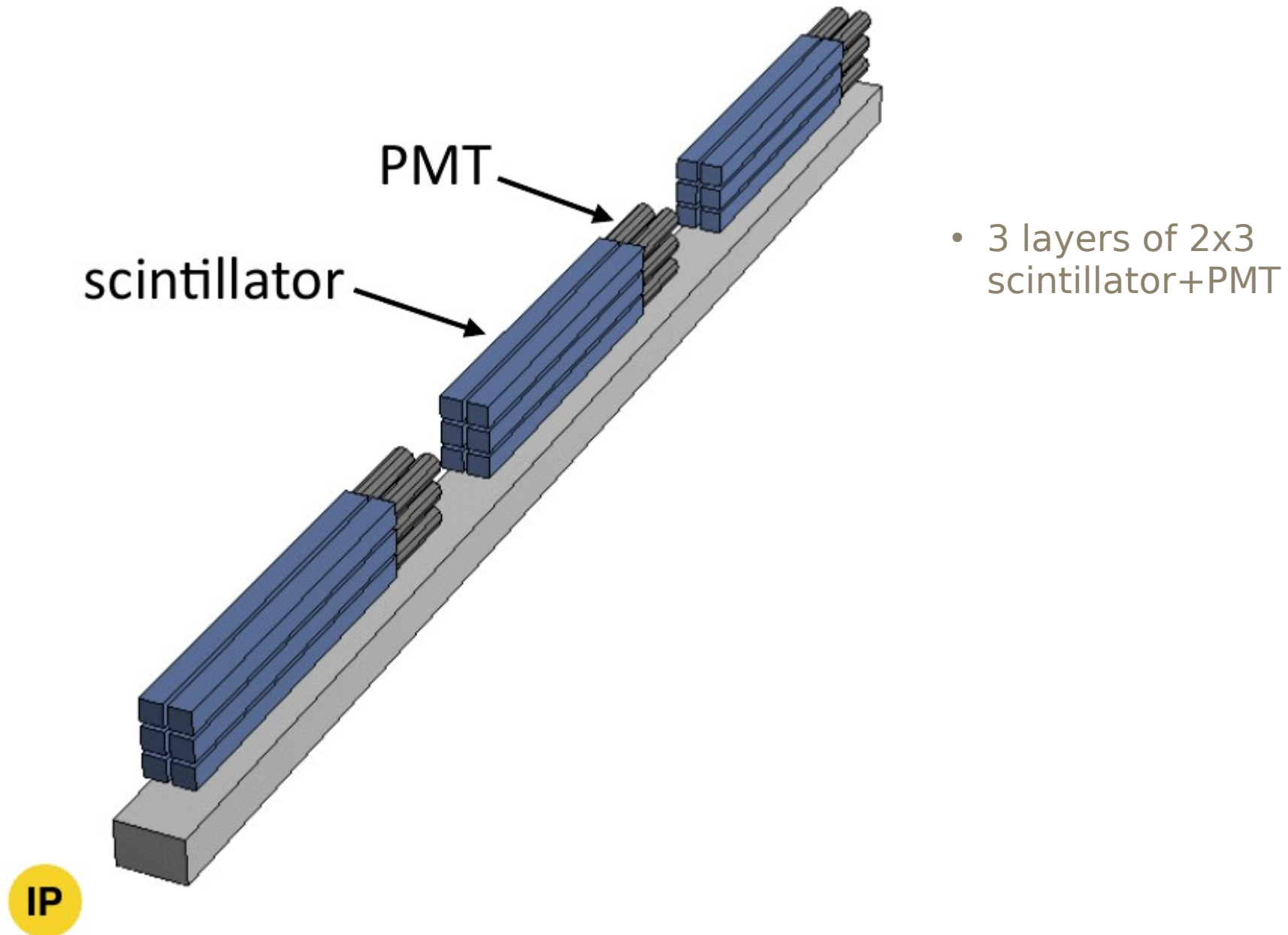


## 1% prototype test

- exercise detector assembly and installation
- establish remote operation
- measure backgrounds, check alignment, perform calibrations, determine sensitivity
- input for full detector design

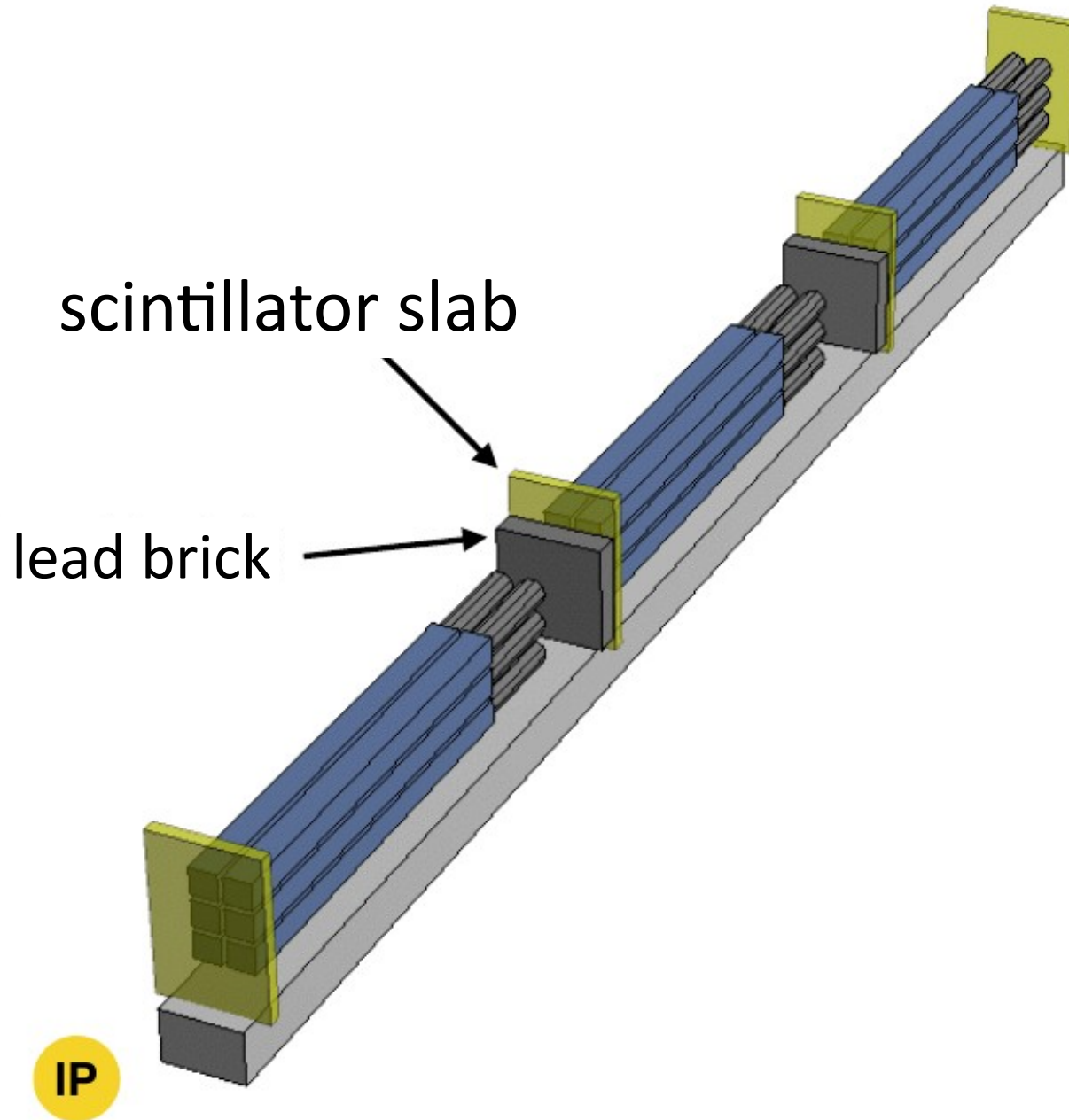


# Demonstrator



IP

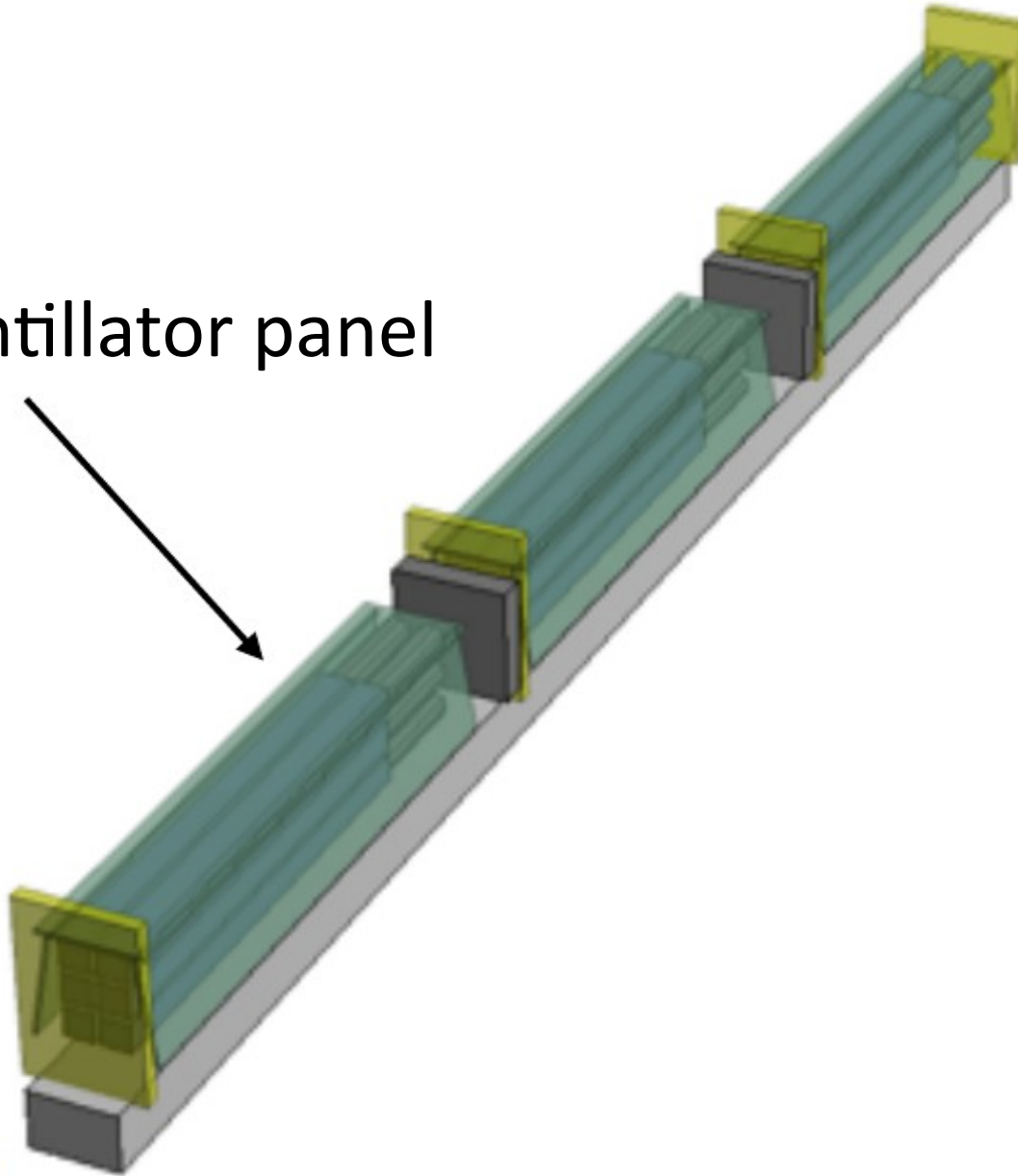
# Demonstrator



- 3 layers of 2x3 scintillator+PMT
- scintillator slabs and lead bricks
  - tag thru-going particles, get time info, shield radiation



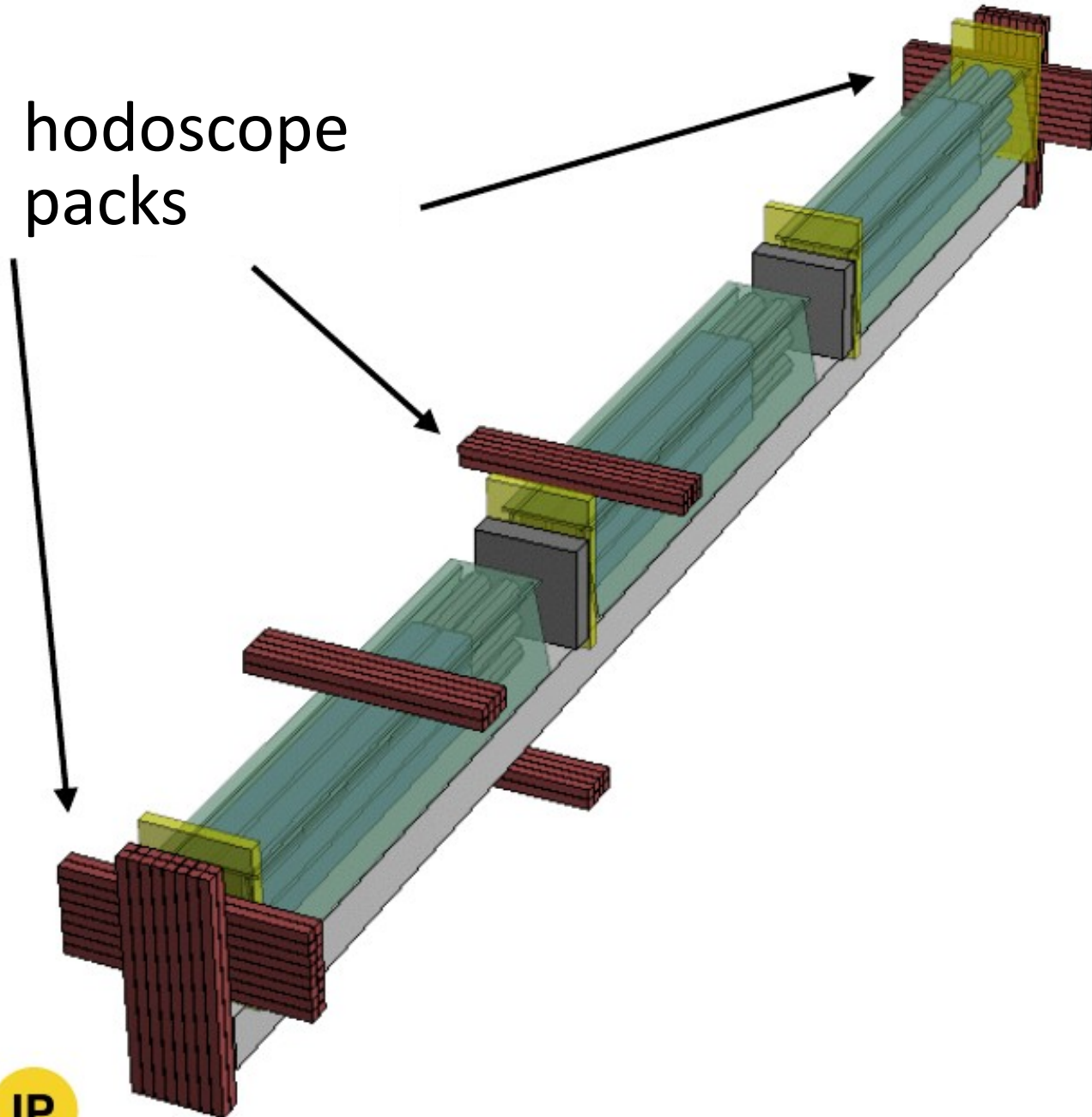
scintillator panel



- 3 layers of 2x3 scintillator+PMT
- scintillator slabs and lead bricks
  - tag thru-going particles, get time info, shield radiation
- scintillator panels to cover top and sides
  - tag/reject cosmic muons

# Demonstrator

hodoscope  
packs



- 3 layers of 2x3 scintillator+PMT
- scintillator slabs and lead bricks
  - tag thru-going particles, get time info, shield radiation
- scintillator panels to cover top and sides
  - tag/reject cosmic muons
- hodoscope packs
  - tracking of beam/cosmic muons

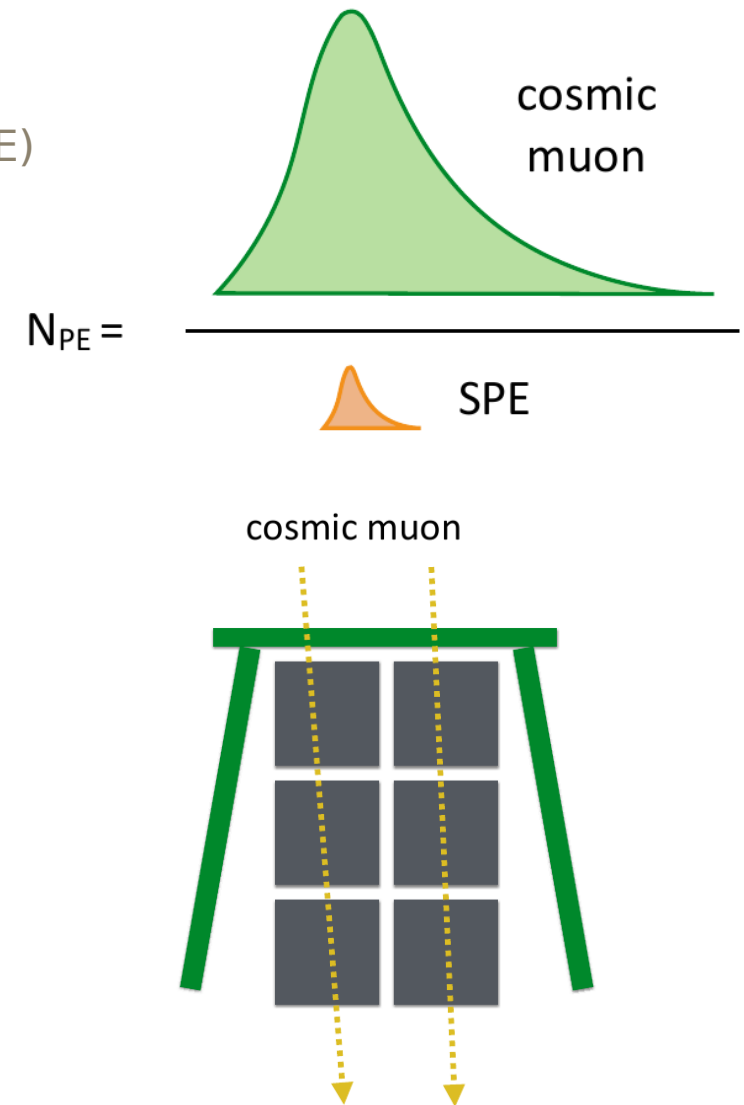
IP



- took data since April '18
  - 2000h, 37/fb
  - and lots beam-off data
- 3 layers of 2x3 scintillator+PMT
- scintillator slabs and lead bricks
  - tag thru-going particles, get time info, shield radiation
- scintillator panels to cover top and sides
  - tag/reject cosmic muons
- hodoscope packs
  - tracking of beam/cosmic muons

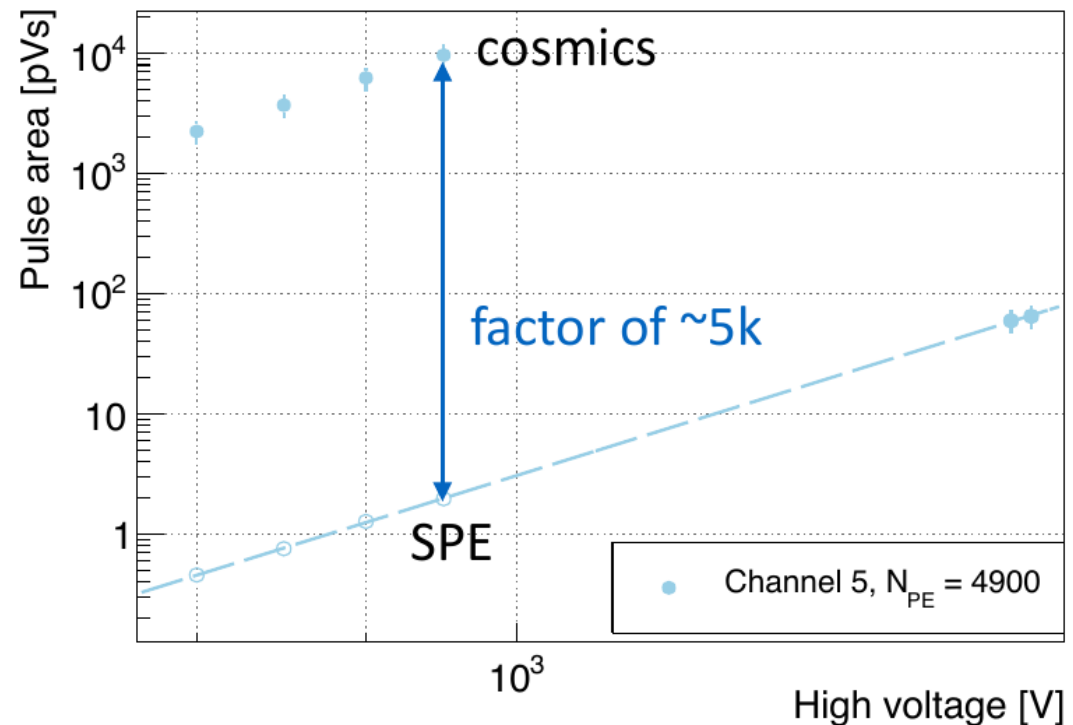
## In situ charge calibration

- calculate  $N_{PE}$  for cosmic muons ( $Q = 1e$ )
  - $N_{PE} = \text{Pulse area (cosmic muon)} / \text{Pulse area (SPE)}$
- extrapolate it to fractional charges by  $Q^2$
- this tells us how small a charge milliQan can detect
- cosmic muons taken from vertical path
- Single PhotoElectron (SPE) from afterpulses
  - validated with LED on bench



## In situ charge calibration

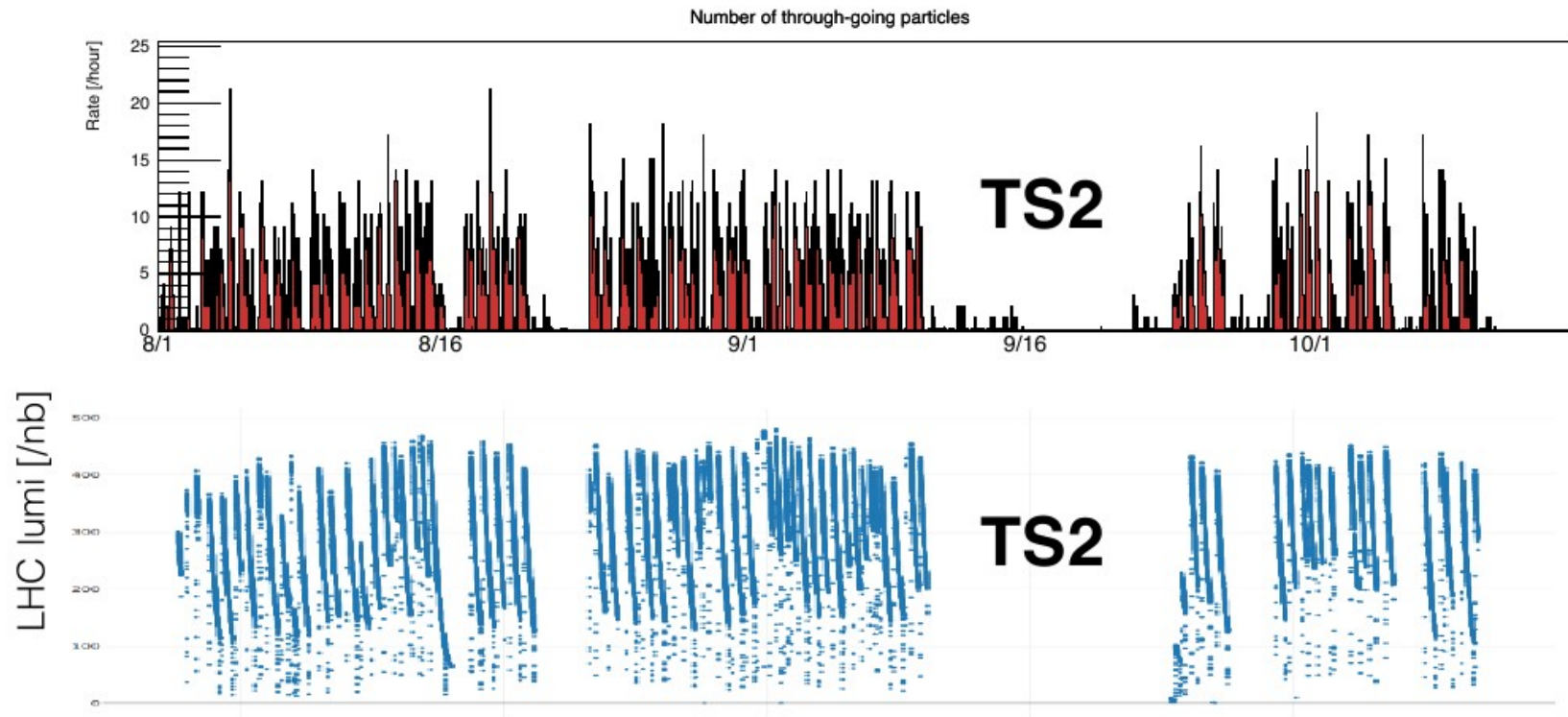
- pulse area as function of HV for a PMT
- $N_{PE}$  for  $Q = 1e$  is  $\sim 5k$
- flight distance of cosmic muons arriving perpendicular to scintillator is 5cm
  - for through-going muons the flight distance is 80cm
  - $N_{PE}$  for through-going muons is  $5k \times 80/5 = 80k$
- $N_{PE} \sim Q^2$ 
  - $N_{PE} = 1$  for  $Q \sim 0.003 e$
- consistent with full Geant4 simulation results





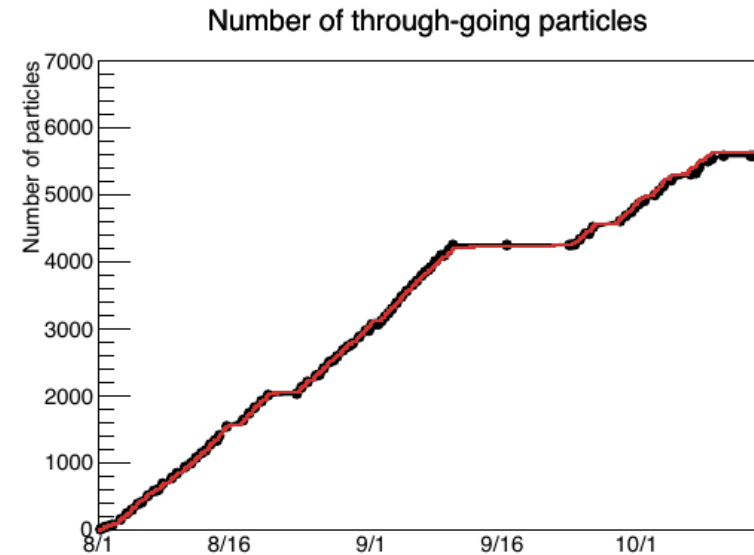
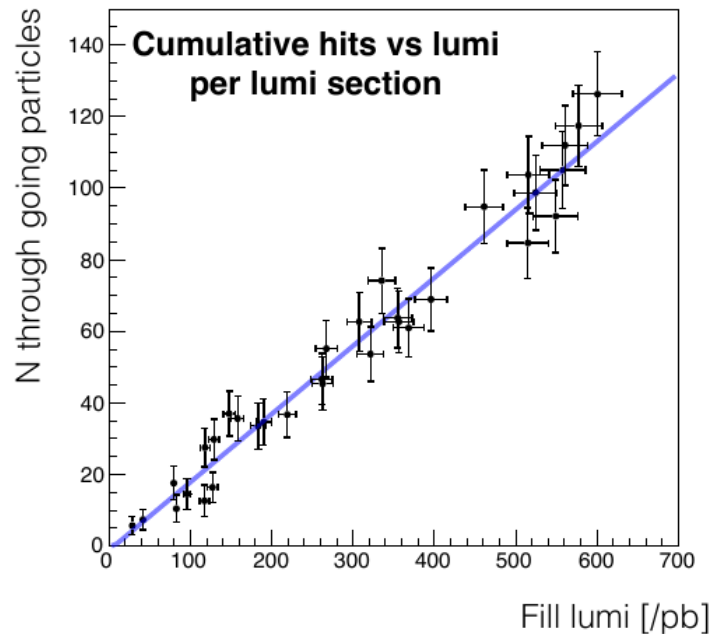
## Alignment

- check alignment with LHC beam
- plot rate of events with muon hit in all 4 slabs



- agrees well with LHC fill / lumi data

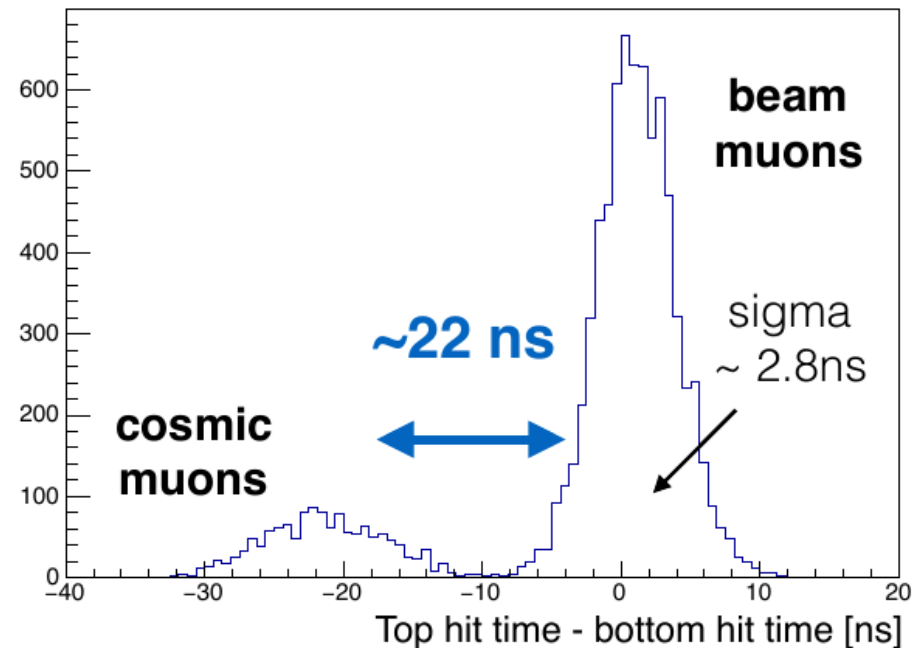
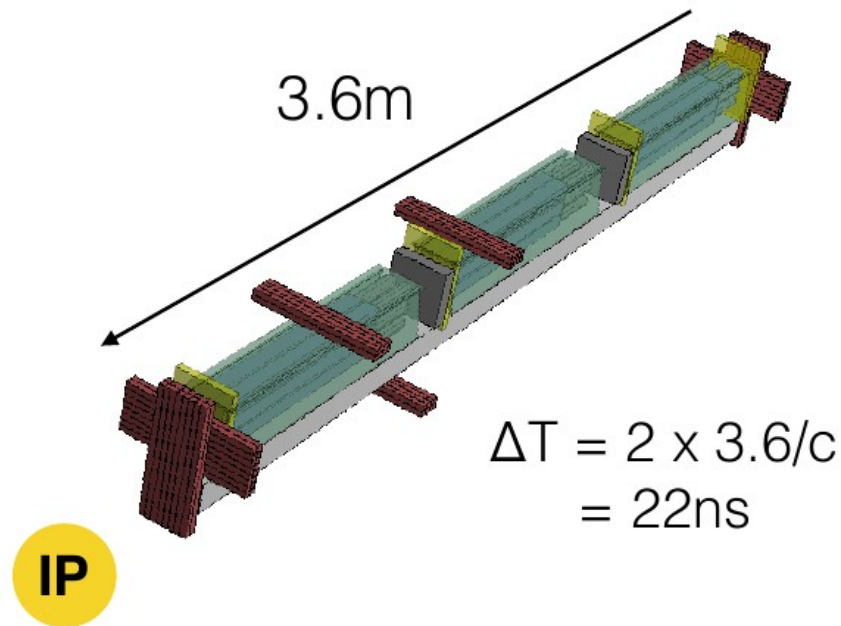
## Alignment



- measured rate:  $0.20 \pm 0.01 / \text{pb}^{-1}$       predicted rate:  $0.25 \pm 0.08 / \text{pb}^{-1}$
- **very good match data – simulation!**
  - uncertainties from B-hadron cross section and amount of material to cross
- in principle precision from survey is sufficient, no need for angular scan

## Timing

- need good timing resolution
  - mCP resolution limited by length of scintillator  $\sim 2\text{ns}$
- when timed-in use time-coincidence to suppress backgrounds
  - eg. cosmics

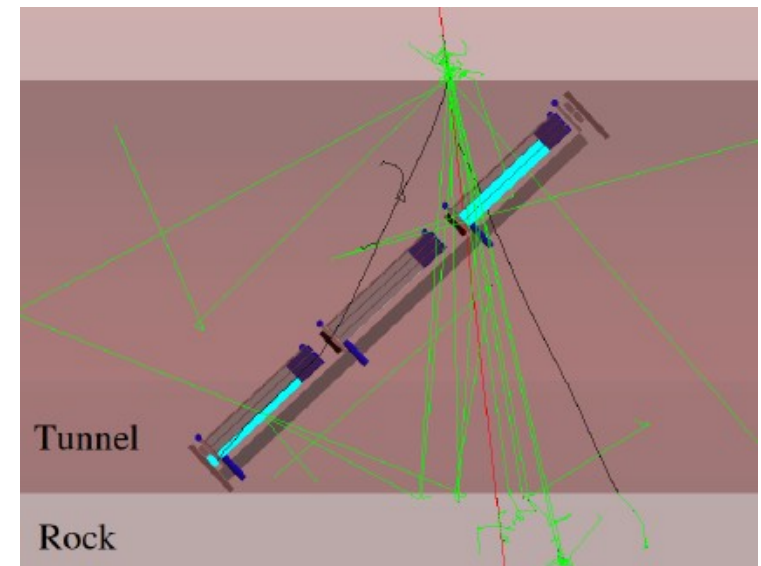
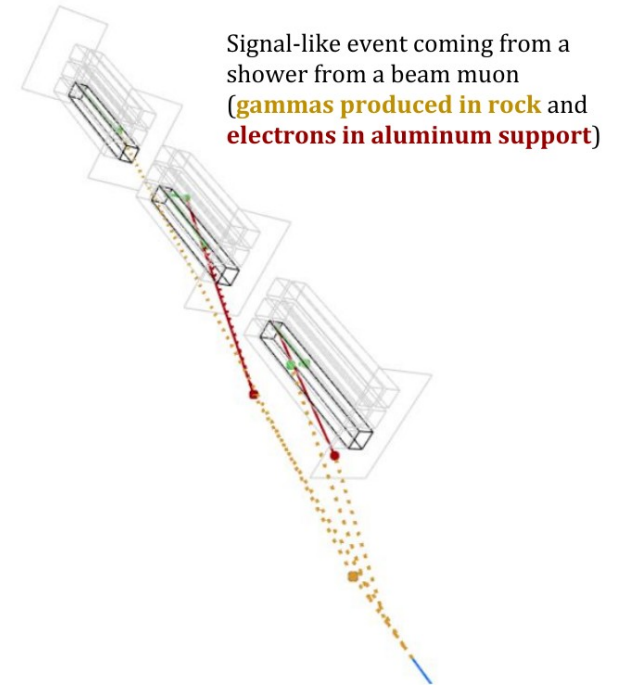




# Demonstrator results

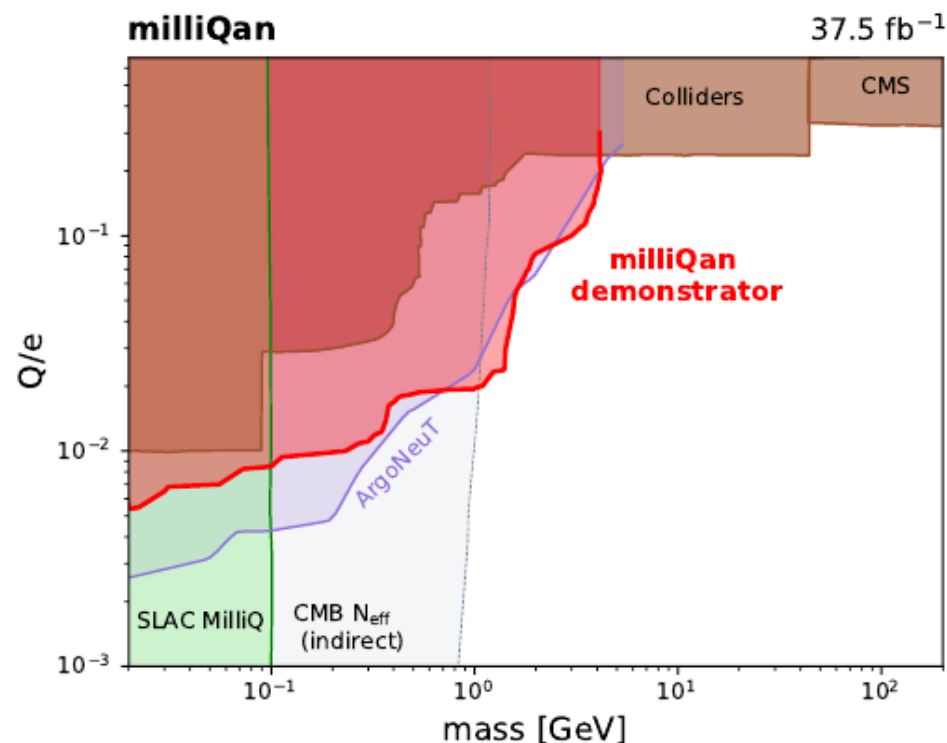
## Backgrounds

- sources of 3 in-time hits
  - PMT dark rate
  - afterpulses
  - radiation
  - showers from beam/cosmic muons
- important lesson from demonstrator
  - background from PMT dark rate subdominant
- further background suppression can be achieved
  - extra shielding
  - tagging of external sources
  - going to 4-layer design



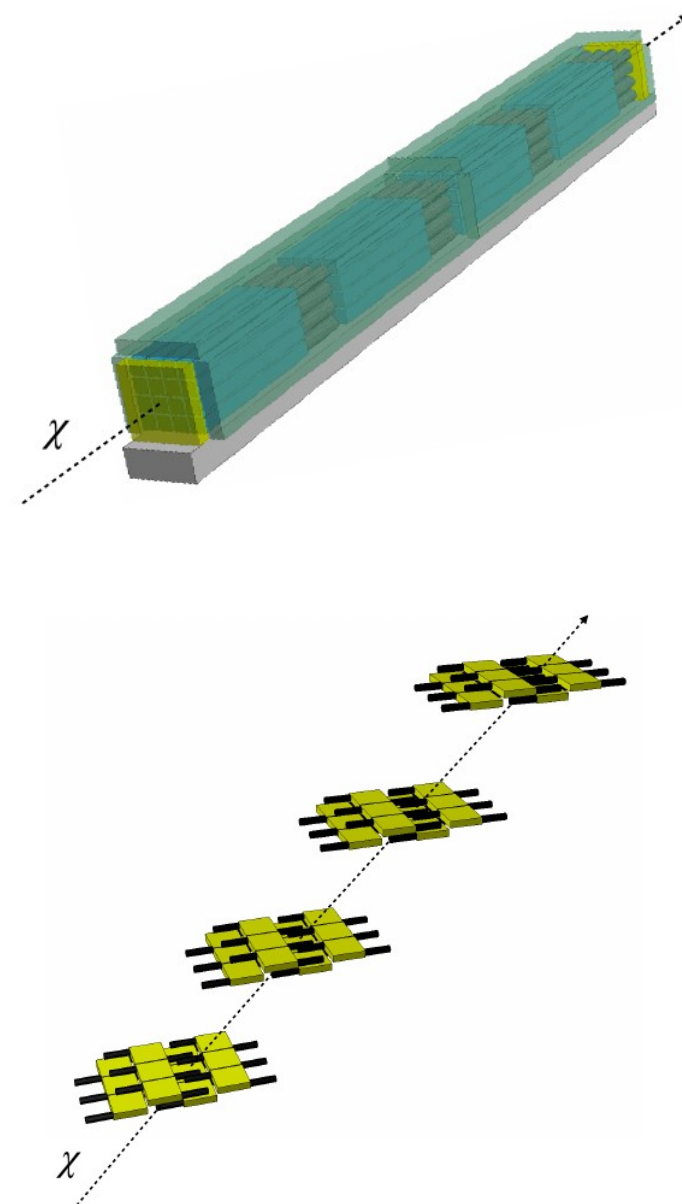
# Demonstrator sensitivity

- analysis with demonstrator data published:  
Phys.Rev.D 102, 032002 (2020)  
arXiv:2005.06518 [hep-ex]
- new sensitivity already with demonstrator data!
  - new particles with masses between 20 and 4700 MeV are excluded for charges between 0.006 and 0.3 e



# Towards Run-3 Detector

- collaboration is gearing up towards a detector for LHC Run-3
- use the knowledge from the demonstrator to build a **4x4x4 bar detector**
  - 5x5x60 cm<sup>3</sup> scintillator bars to fit 4 layers
  - thicker 5cm scintillator panels for active veto
- additionally, add a **4-layer slab detector**
  - bar detector good for low charge, but at high mass limitation is angular acceptance
  - 4 layers of 40x60x5 cm<sup>3</sup> scintillator slabs
- preparing construction this Fall
- installation and commissioning at CERN in time for LHC Run-3 operations in Spring 2022



# Towards Run-3 Detector

## Building bars: “easy”!



3D printed  
PMT casing

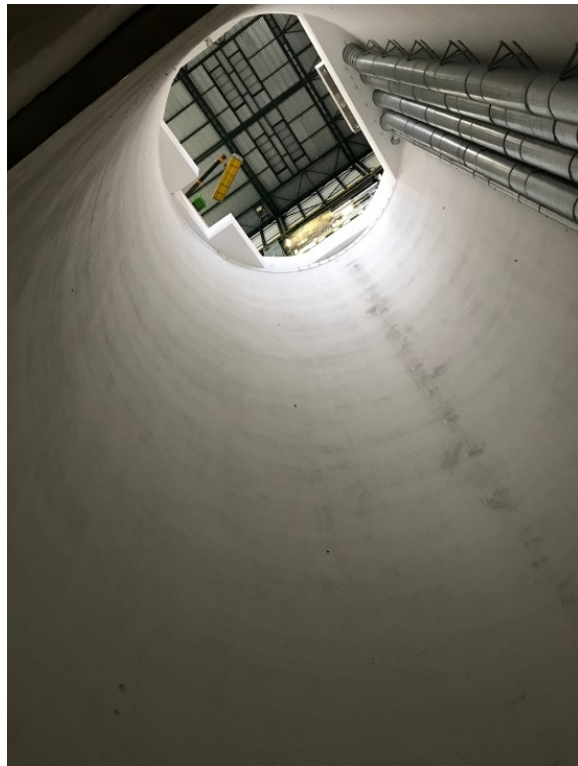
Bars wrapped in layers of reflective and light blocking materials (including tyvek, tinfoil, electrical tape)





# Towards Run-3 Detector

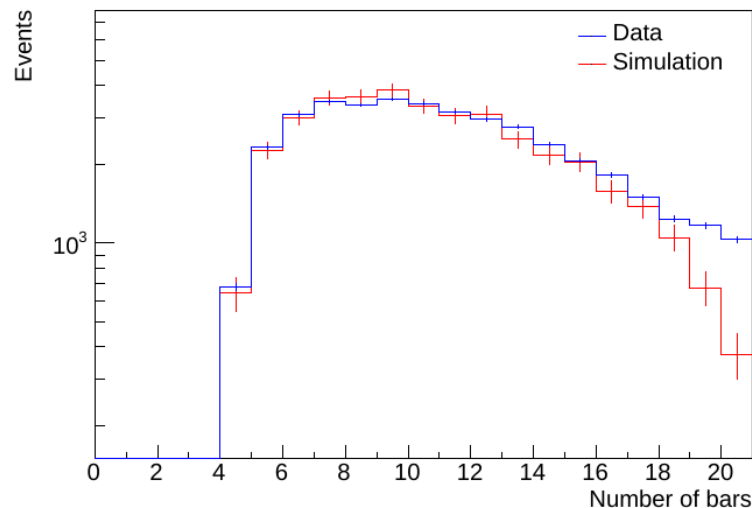
## Installation



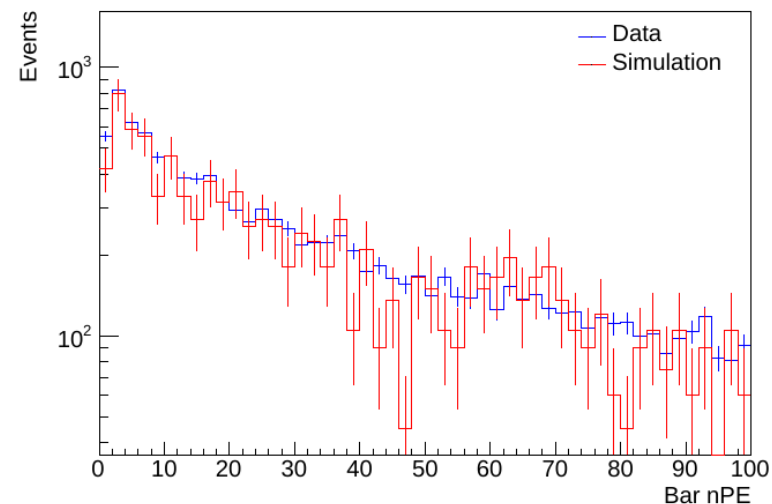
## Detector simulation

- cosmic shower simulation refined using demonstrator data
- examples of well predicted quantities: (see arXiv:2104.07151)

number of bars with detected pulse from cosmic muons



number of photoelectrons in pulses from cosmic muons

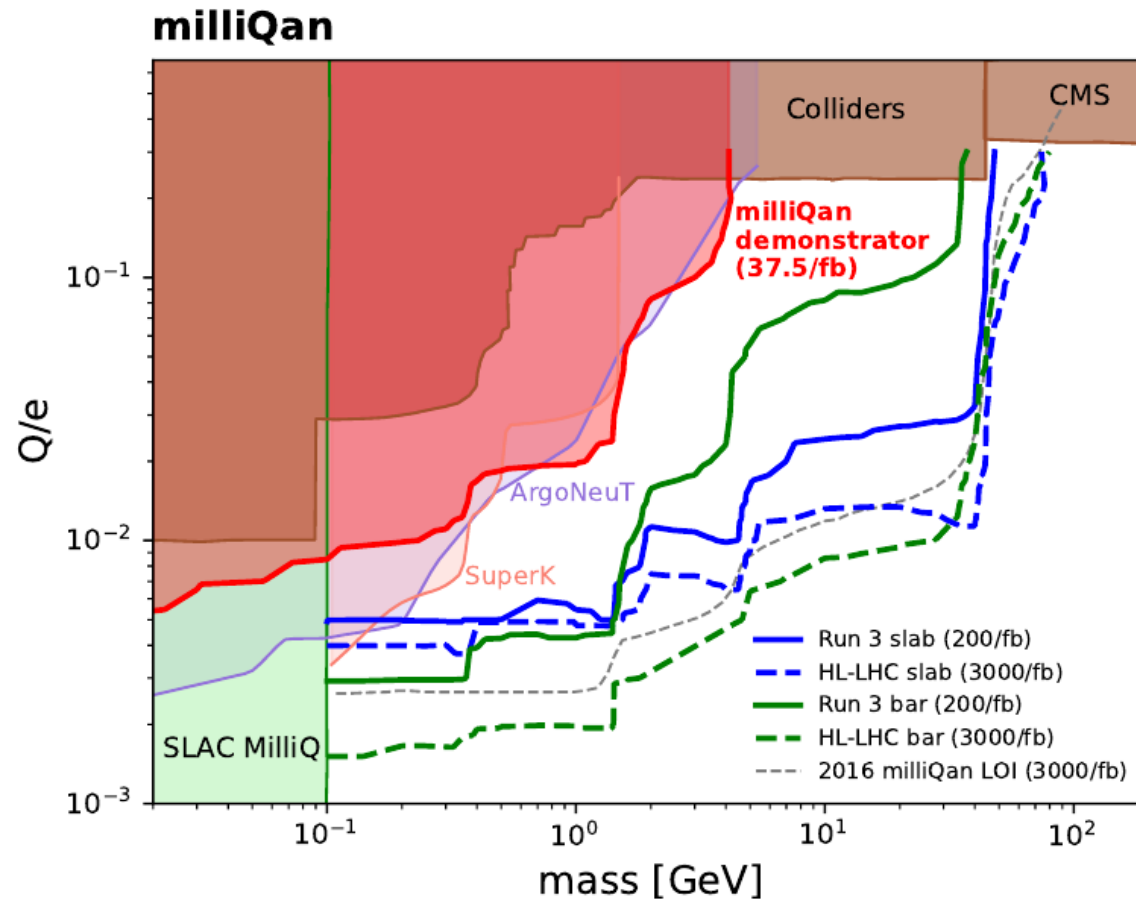


- main expected background very well predicted in simulation  
→ optimize design for best sensitivity

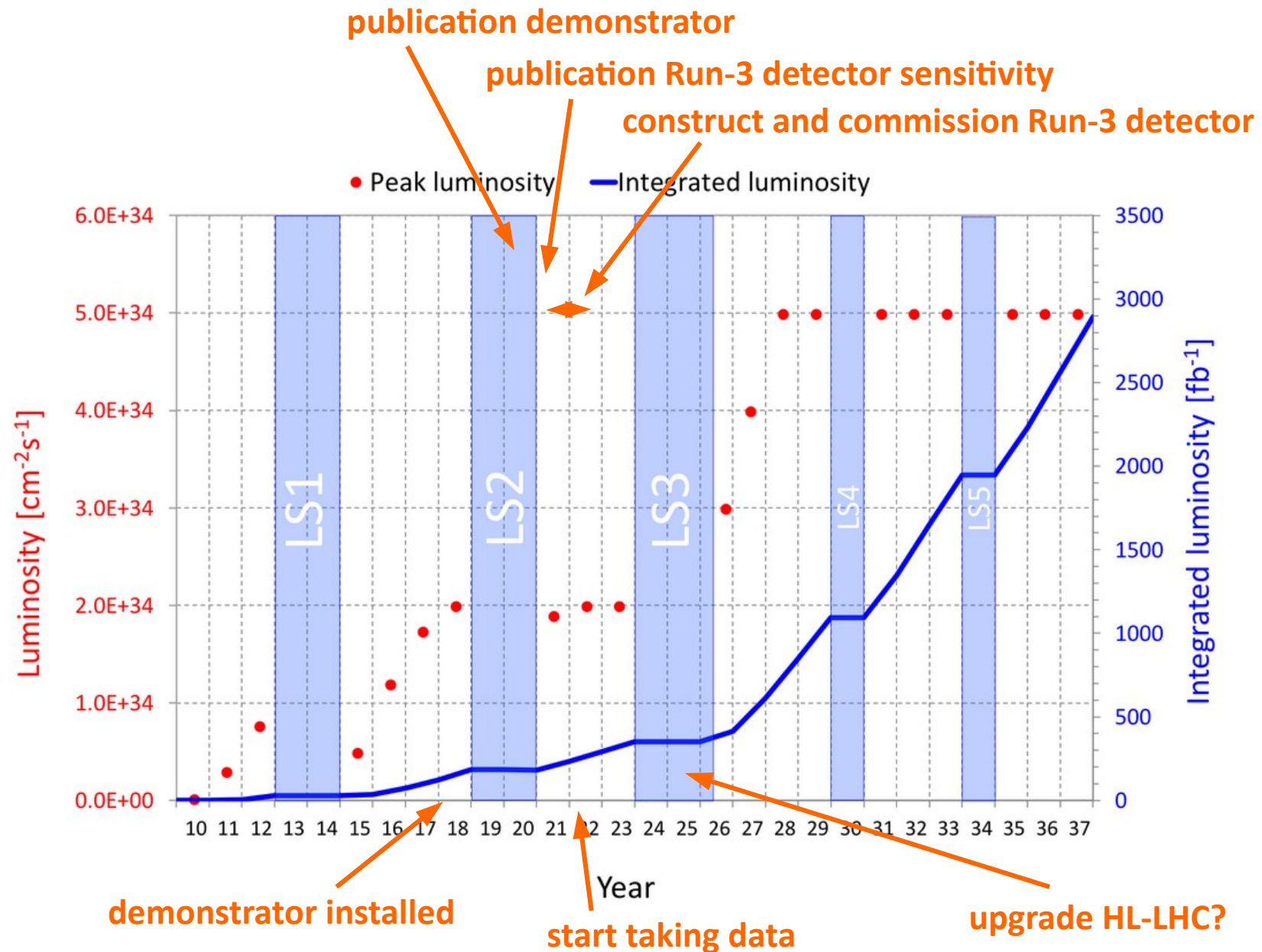


## Sensitivity projection

- expect to rapidly enter new discovery territory

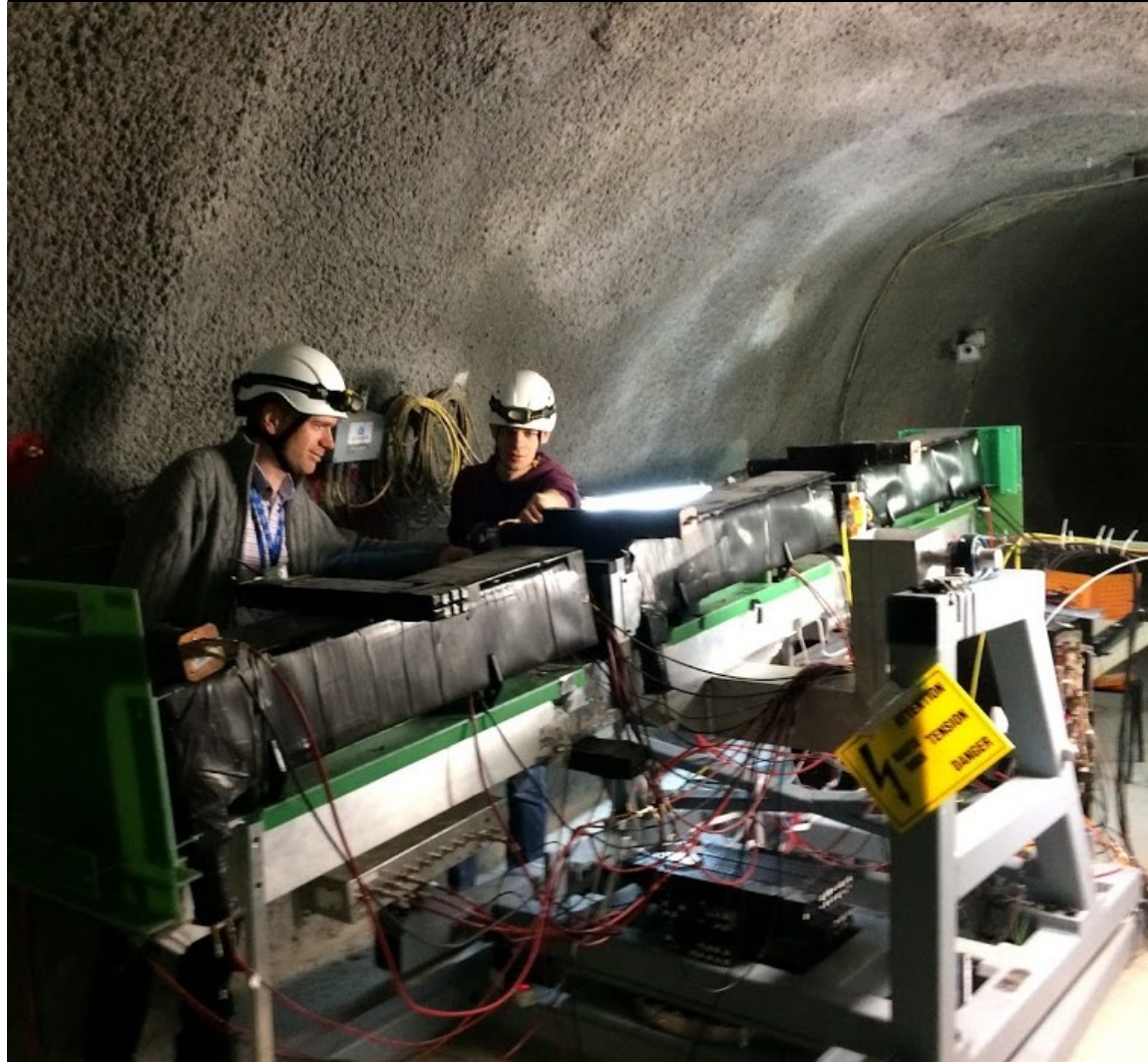


# milliQan timeline



# Conclusions

- the milliQan detector provides unique sensitivity to millicharged particles
  - uncovered phase space  $0.1 < m < 100 \text{ GeV}$ ,  $Q < .3e$
- 1% demonstrator successfully validates feasibility
  - successful construction and operation; first sensitivity
  - many lessons learned from commissioning and in-situ background studies
- Run-3 detector is underway
  - 4x4x4 bar detector complemented with 4-layer slab detector
  - large new sensitivity projected



# The Collaboration



C. Hill, B. Francis,  
M. Carrigan, L. Lavezzo, B. Manley



D. Stuart, C. Campagnari,  
M. Citron, B. Marsh,  
B. Odegard,  
R. Schmitz, F. Setti,  
R. Heller



D. Miller,  
M. Swiatlowski



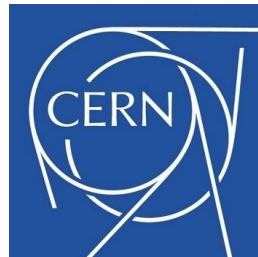
S. Lowette



A. Haas, M. Ghimire



Y-D. Tsai



A. Ball, A. De Roeck,  
M. Gastal, R. Loos,  
H. Shakeshaft



M. Ezzeldine,  
H. Zaraket



F. Golf



J. Brooke,  
J. Goldstein