

# Why INO ?

V. M. Datar

Tata Institute of Fundamental Research

Email: [vivek.datar@tifr.res.in](mailto:vivek.datar@tifr.res.in)

*Vigyan Samagam, Nehru Science Centre, Mumbai (8<sup>th</sup> May 2019)*

## Prelude...

- Indian scientists working at Kolar Gold Fields discovered atmospheric neutrinos while measuring cosmic muons at various depths (1965)
- Subsequently a proton decay experiment was set up at -2.3km and -2km depths but no p-decay events found. ~ 1990 KGF expt closed down.
- Indian scientists started an initiative in 2000 to re-initiate neutrino research. Resulted in INO project, lab under a mountain (sanctioned by Gol in Jan 2015)
- **Studying neutrinos takes us beyond the Standard Model of particle physics!**

# The INO Collaboration

Collaborating Institutions:

- AMU
- BHU
- DU
- HPU
- IGCAR
- IITG
- IMSc
- JU
- MU
- PRL
- SINP
- SU
- UoH
- BARC
- CU
- HNBGU
- HRI
- IITB
- IITM
- IOP
- KU
- NBU
- PU
- SMIT
- TIFR
- VECC

• INO Collaborating Institutions

+IISER (Mohali), American College ,  
Tezpur Univ, CKU (Gulbarga)

~28 institutions (national labs, Universities,  
IITs) participating

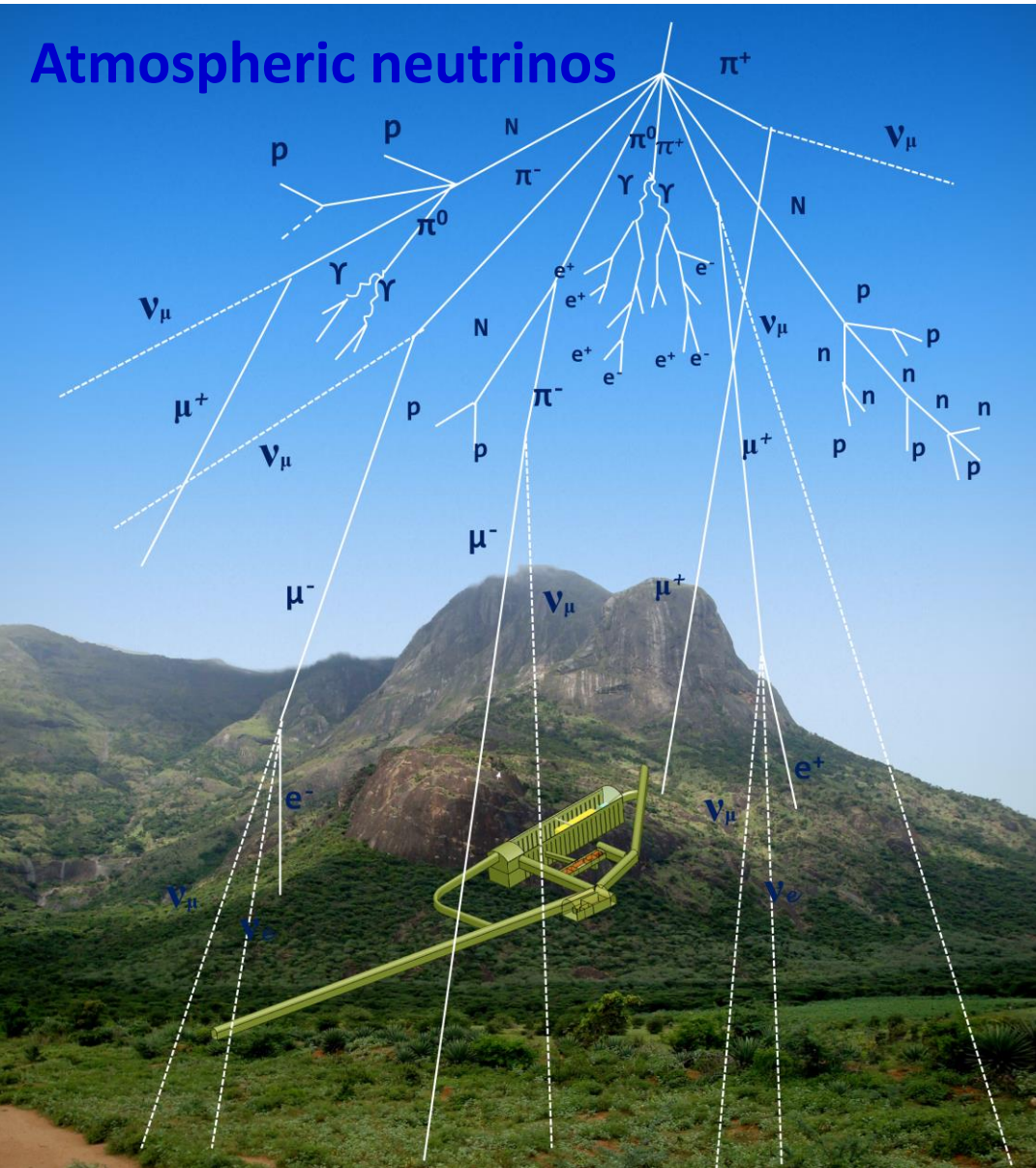


Participants of the INO Collaboration meeting at Madurai Kamaraj University (22-23 March 2018)

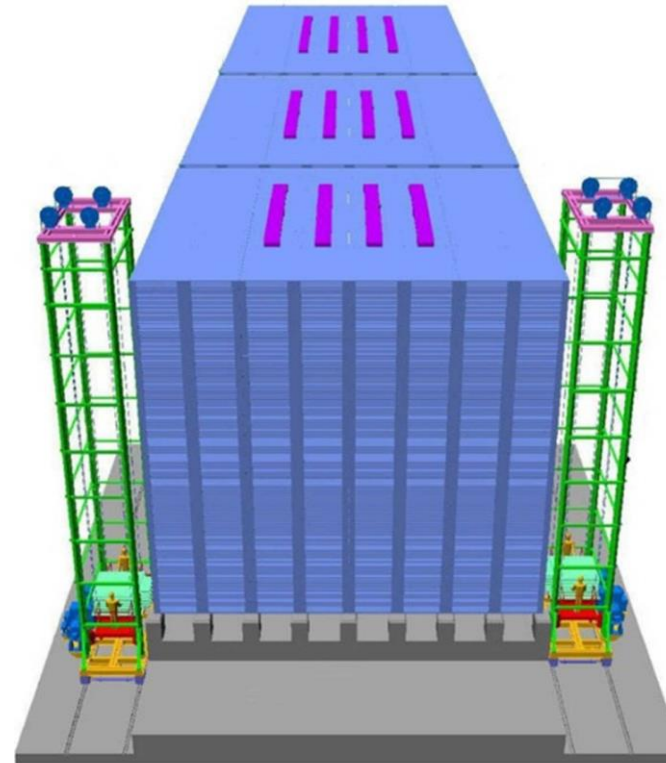
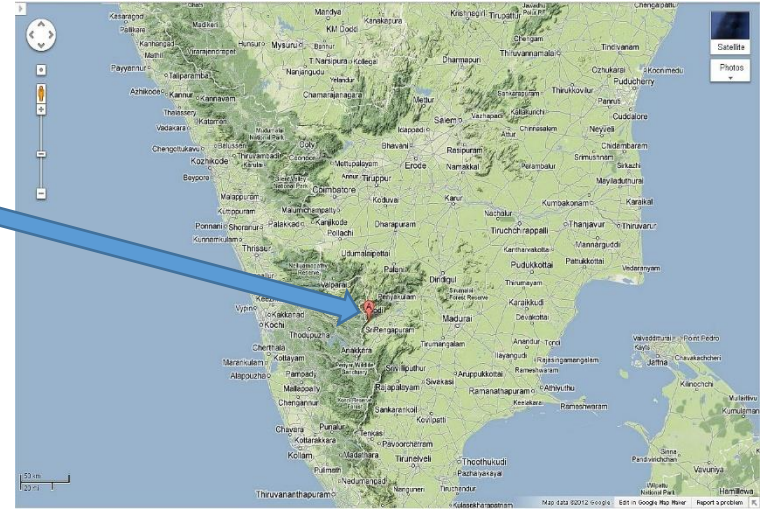


# INO at Pottipuram (Theni)

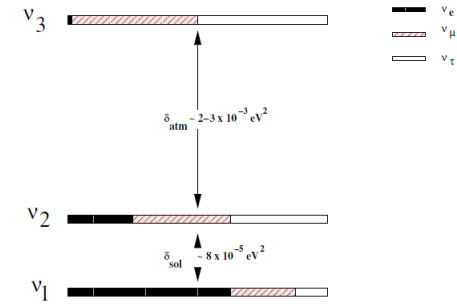
## Atmospheric neutrinos



INO site



## Mass ordering of $\nu$



Is  $m_1, m_2 < m_3$  (NH)  
or  $m_3 < m_1, m_2$  (IH) ?

# Experiments planned at INO

- **Atmospheric neutrinos @ ICAL** (NH/IH, neutrino mixing parameters, DM annih., NSI), anomalous KGF events, Magnetic Monopole search....
- **Neutrinoless Double Beta Decay in  $^{124}\text{Sn}$**  using cryogenic bolometric detector – TINTIN (TIFR led collab.)
- **Dark Matter** search using a cryogenic scintillator - DINO (SINP led collab.)
- **Low energy accelerator for nuclear reaction cross sections ~ Gamow energy** of astrophysical interest (IUC-DAEF + Univ., IIT groups): 3 MV Tandetron accelerator + Recoil Mass Separator + gas jet target

## Why make a huge magnet?

- **Neutrinos** need matter (iron) to interact and produce **charged particles** ( $\mu^\pm$  with charge of  $\pm 1$ )
  - These bend in magnetic field in opposite directions
- ⇒ **identification of  $\nu_\mu$ /anti- $\nu_\mu$  event**

## Why underground?

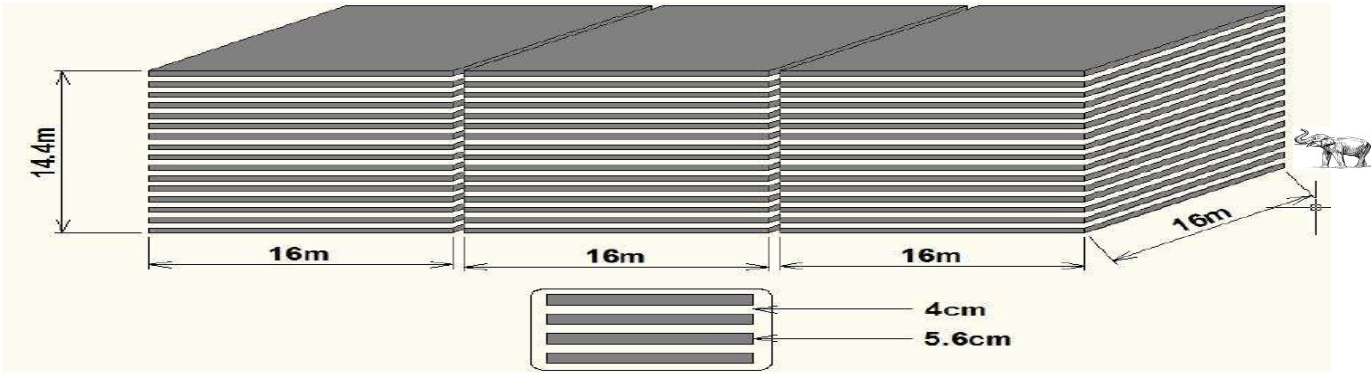
***Cosmic muons*** most important background.

**How deep ?** 1 km  $\Rightarrow$   $\sim 10^6$  reduction  $\Rightarrow$

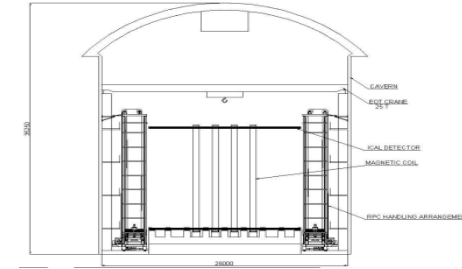
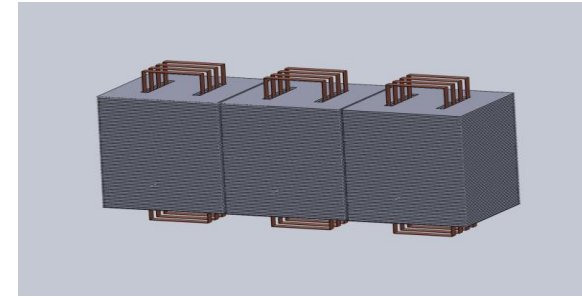
**Mines/tunnels**

**Logistics better at tunnel**

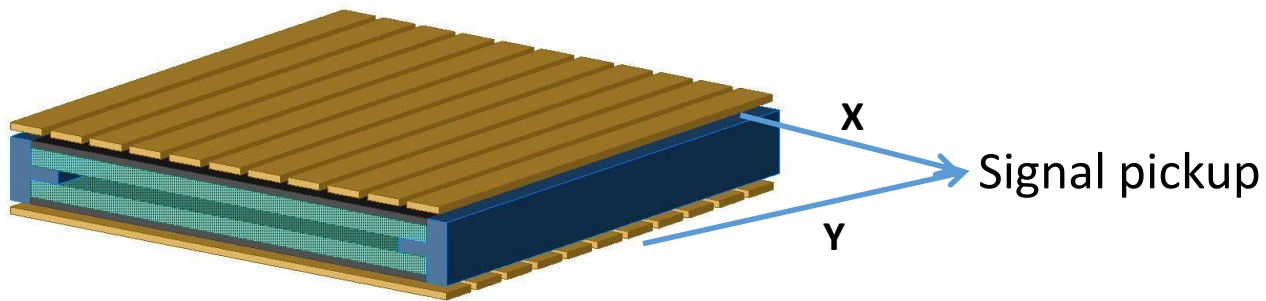
# Schematic of Iron Calorimeter at INO



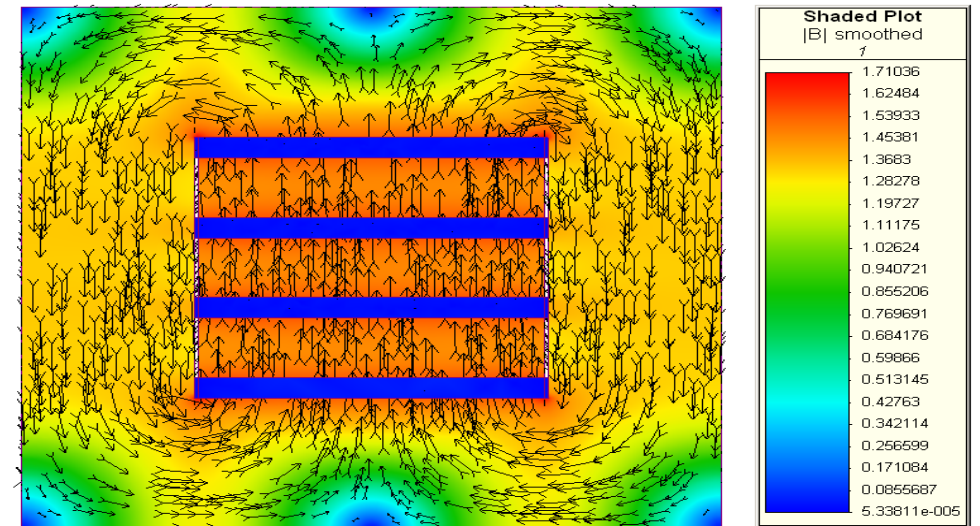
51 kton **world's largest electromagnet**



17 kton module has 150 layers Fe+RPC  
B-field > 1 Tesla (90%)



Glass RPC for detecting charged particles  
~**30,000 RPCs** required, ~ 3.8 M channels

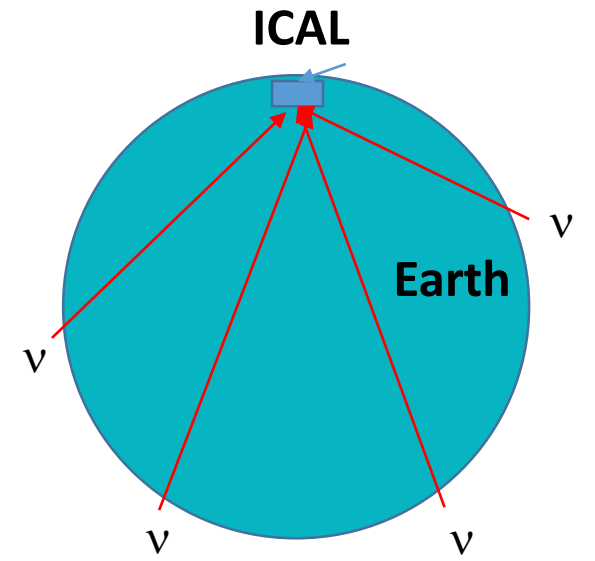


Simulated B-field for 60 kA-turns



# Iron Calorimeter (ICAL) detector

- Atmospheric neutrinos (**free!**): large range of energies ( $E_\nu \sim 1-20$  GeV) and matter traversal distances  $\sim 1 - 13000$  kms
- ICAL will identify  $\nu_\mu$  and  $\bar{\nu}_\mu$  events via *charged current interaction*  
 $\nu_\mu + n \rightarrow \mu^- + p$ ,  $\bar{\nu}_\mu + p \rightarrow \mu^+ + n$  and tracking outgoing  $\mu^- / \mu^+$
- Matter affects  $\nu_\mu$  and  $\bar{\nu}_\mu$  differently and can be used effectively

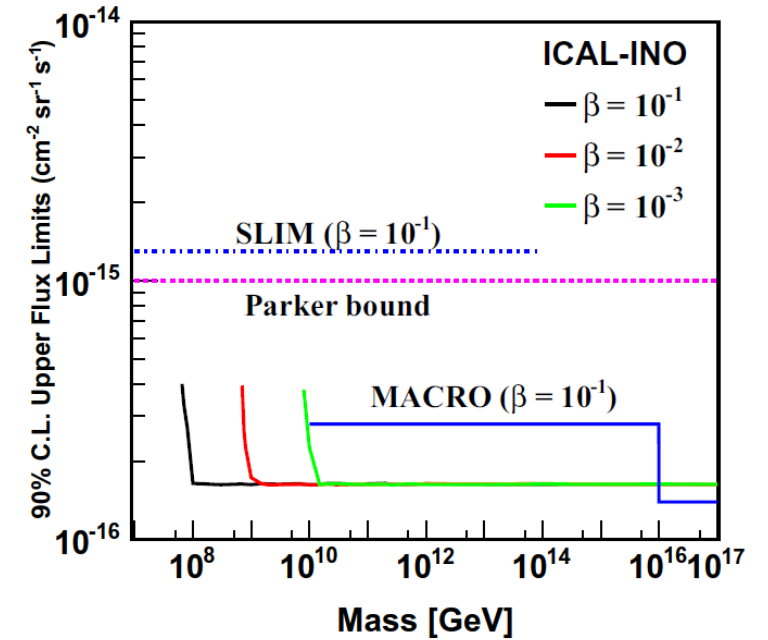
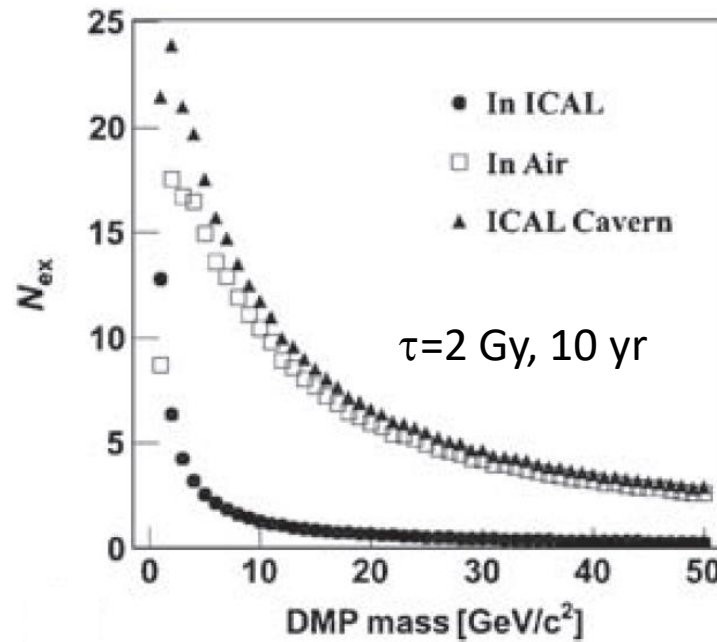
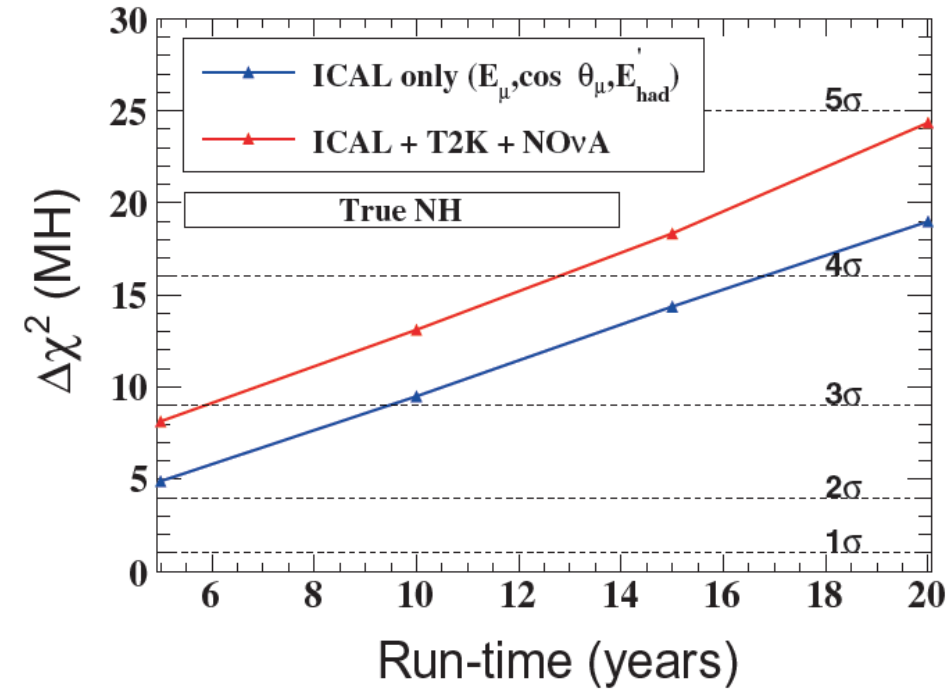


[White paper on Physics with ICAL : Pramana **88**, 79 (2017)]



# Physics with ICAL – some examples

- Neutrino **mass ordering**, mixing parameters, non-standard interactions, neutrino decay, **anomalous KGF events**, **magnetic monopoles** ...



## **The mini-ICAL (85 ton, 4m × 4m × 11 layers of Fe) @ Madurai**

- Performance of Magnet: Measured magnetic field (*using sense coils and Hall probes*) vs 3D FE simulation
- Performance over long period of RPC including DC-DC supply, FE electronics in fringe B-field, EMI, closed loop gas system.....
- Feasibility of Muon Spin Rotation ( $\mu$ SR) for information about B-field complementary to sense loop and Hall probe data
- Measure  $\Phi(\mu^+)$ ,  $\Phi(\mu^-)$  at Madurai (near equator) and compare with simulation (by Athar, Honda)
- Prototype cosmic muon veto detector for mini-ICAL



**Plate machining Job**

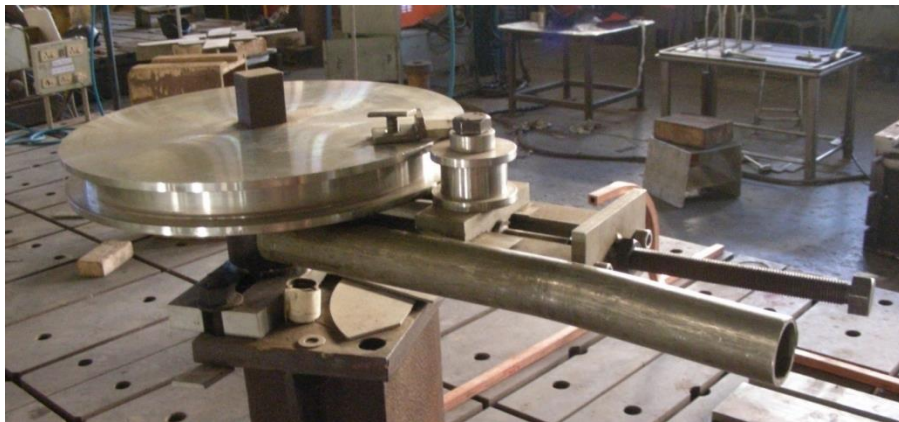


**Spacers and Pins**



**Copper Conductor Spool**

## **Magnet Components (Core & Coil) **BARC** group**



**Conductor bending machine**



**Conductor straightening machine**



**Coil fabrication**



# mini-ICAL assembly



Pillars for magnet, G10 boards for Cu coils



OFHC Cu "U" and "C" sections



Magnet assembly in progress



Assembly of iron plates of mini-ICAL magnet

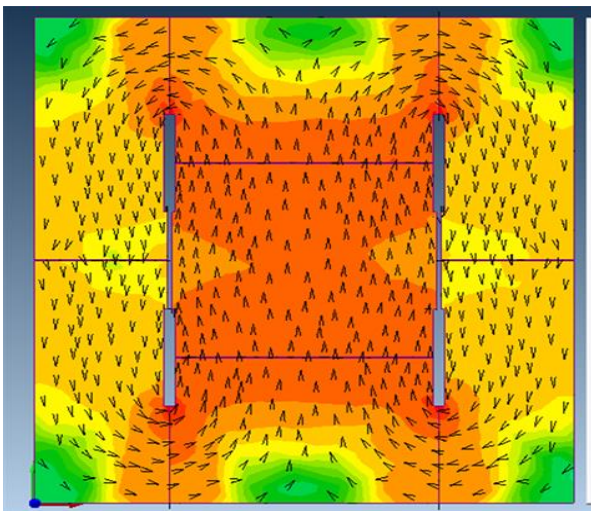


Cu Coil Brazing

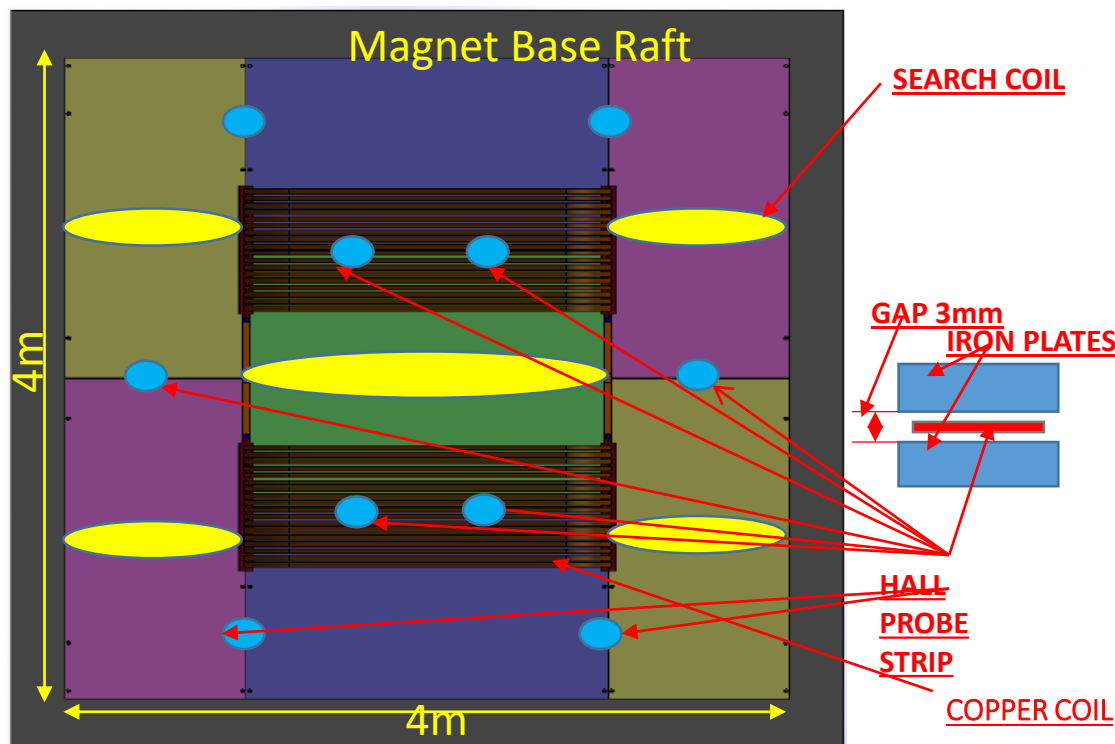


# Low conductivity water cooling system for magnet & power supply

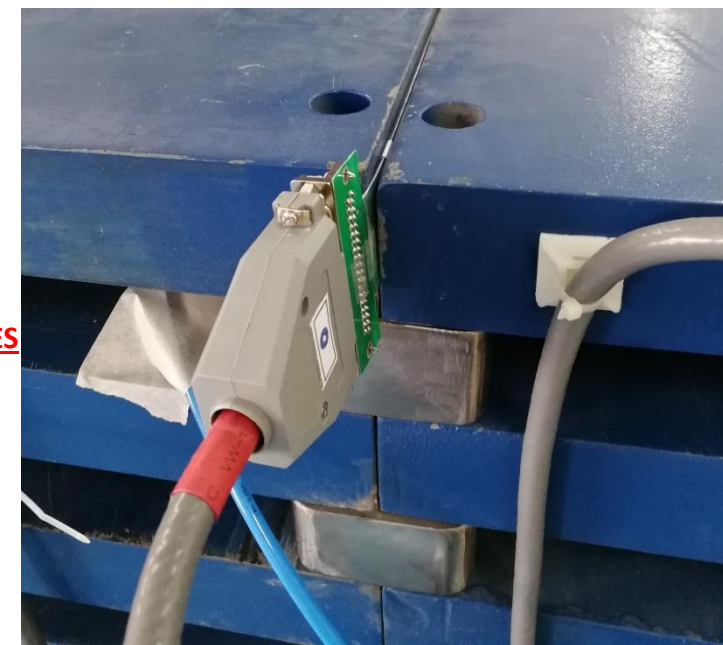




**Field map at 26kAT**



**Magnetic measurement system  
(1<sup>st</sup>, 6<sup>th</sup>, 11<sup>th</sup> layer)**



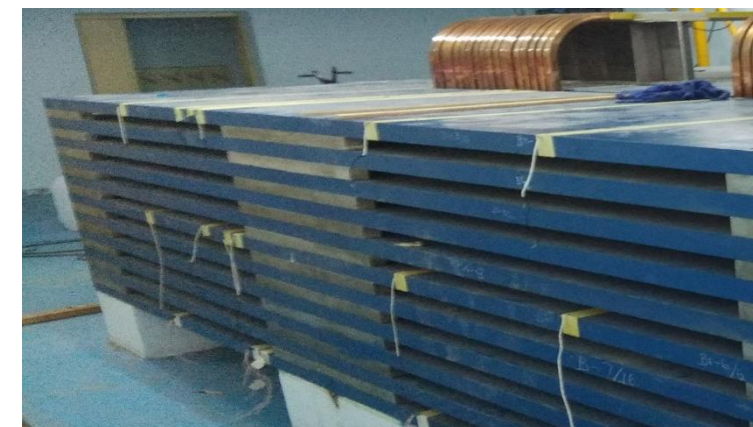
**Hall probe PCB in the gap**



**Magnet power supply  
30V DC, 1200 AMP**



**Hall probe PCB strip**



**Search coils for flux measurement**



# RPC re-assembly



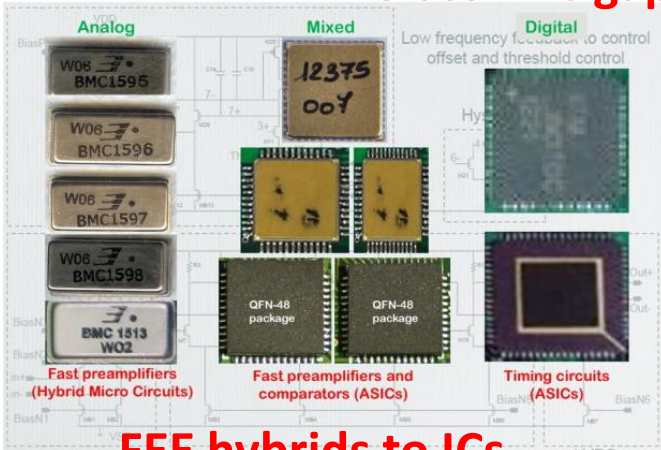




**Glass RPC gap at St. Gobain**



**Closed loop gas system**



**FEE hybrids to ICs**



**DAQ card with FPGA, HPTDC**



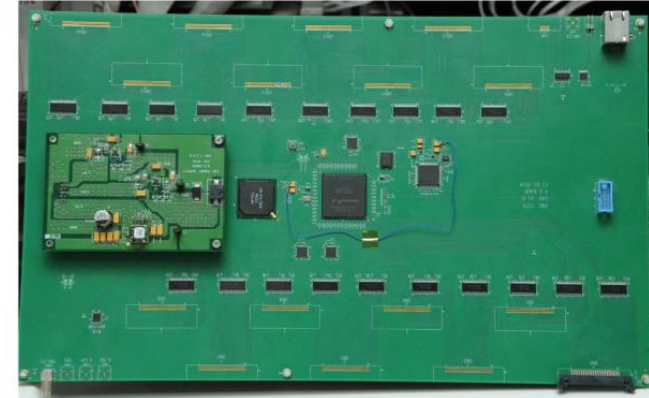
**±5 kV DC-DC HV card**



**8 channel FEE board**



**Trigger and Calibration system**





## Mini-ICAL team members:

**BARC:** Sourabh Pathak, Sandip Patel, S. Ajith, N.S. Dalal, S.P. Prabhakar, T.S. Sreenivasan, D.N. Badodkar (DRHR), S.P. Srivastava, K.N. Karn, P.I. Hadagali, P.K. Biswas, Alok Tripathi, Sachin Dolas, Prabhat Singh, Vinay Sharma, Sanjay Patil, Suresh Jaiswar (CDM) , R. Rengan, K. Srinivas (CED), S. Achrekar, N. Ayyagiri, A. Behere, V.B. Chandratre, D. Das, A. Jain, N. Kamble, T. Kasbekar, H. Kolla, A.Manna, S. Mohanan, S. Moitra, P.M. Nair, S. Padmini, M. Punna, S.M. Raut, S. Prafulla, S. Sikder, M. Sukhwani (ED), P.S. Shetty, B. Sivaramakrishna, Mathew Dominic, Shashank Padwal (TSD), special thanks to Dr. Sekhar Basu

**SINP:** N.K.Mondal

**TIFR:** B.S. Acharya, Vishal Asgolkar, Rajkumar Bharathi, Apoorva Bhatt\*, Santosh Chavan, S. Dasgupta, V.M. Datar, Upendra Gokhale, Darshana Gonji, S.R. Joshi, Suresh Kalmani, Puneet Kaur, A. Lokapure, G. Majumder, Suryanarayan Mondal\*, P. Nagaraj, Neha\*, Pathaleswar, S. Pethuraj\*, K.C. Ravindran, Mandar Saraf, B. Satyanarayana, Ravindra Shinde, Dipankar Sil, Thoi Salam Singh, N. Sivaramakrishnan, Pavan V., L. Umesh, Suresh Upadhyaya, Piyush Verma, E. Yuvaraj

**VECC:** S.K. Thakur, A. Bera, A. Ghosh, Noor Mohamed

### Mini-ICAL Design Safety Review Committee of BARC Safety Council for their suggestions

Essar Steel (steel plates), Green & Green (assembly), St. Gobain (RPC gaps), Ferrite India (Pune), BEC (Bhilai), Entech (B'luru) ....

**Mag. Monopole, DM particle decay:** Nitali Dash\*, VMD

**FE 3D EM simulation of ICAL magnet:** Shiba Behera<sup>+</sup>, M.S. Bhatia, VMD

\* INO Graduate students + HBNI and NPD-BARC

# ICAL GOALS

Neutrino mass ordering using matter effect on  $\nu_\mu$  & anti- $\nu_\mu$

Neutrino mixing parameters, Non-standard interactions, CPT violation, decay, Sterile neutrinos, Earth Tomography

Search physics beyond Standard Model incl. Magnetic monopoles, long-lived particles, dark matter

Long-term detector for atmospheric & astrophysical phenomena, search for unknown, Multimessenger astronomy

# INO GOALS

Low radiation lab for experiments in High energy physics, Biology, Material Science, Geology, Nuclear astrophysics

Large state-of-art experiment(s) in India for developing expertise and future pioneers, detector development, education & research opportunities for students

ICAL@INO

INO