

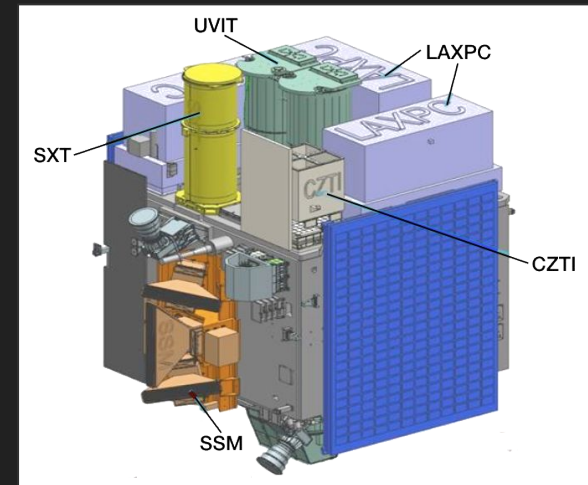
AstroSat/CZTI mass model

Sujay Mate
TIFR Mumbai



The AstroSat

- First multi-wavelength space observatory launched by India.
- Four co-pointing telescopes and one sky monitor.
- Broadband coverage from UV to hard X-rays.
- Spectral and temporal studies of astrophysical sources.



The Cadmium Zinc Telluride Imager

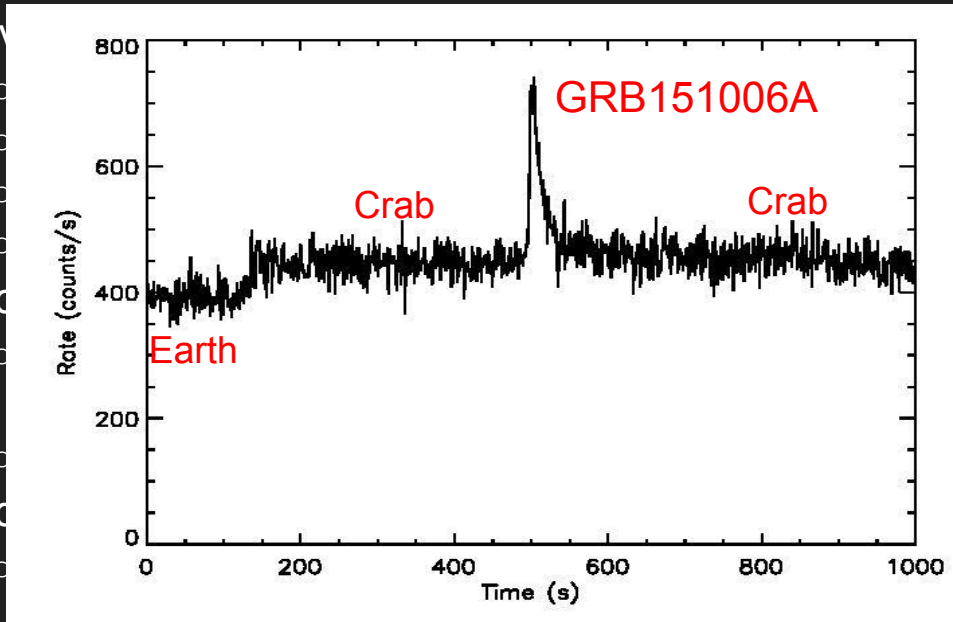
- Overview:
 - Imaging and spectroscopy in 20 - 200 keV (and polarimetry!)
 - Effective area: 488 cm² at E <90 keV
 - Energy resolution: 9 % at 60 keV
 - Angular resolution: 8 arc min
- Scientific goals
 - Spectral and temporal studies of galactic X-ray binaries and pulsars
 - Polarimetry of bright sources
- Additional advantage:
 - Open sky detector > 100 keV → Excellent for detecting bright high-energy transients such as Gamma-ray bursts and Soft gamma repeaters (SGRs)



Slide credits: Mithun NPS

The Cadmium Zinc Telluride Imager

- Overview
- Science
- Applications



metry!)

s and

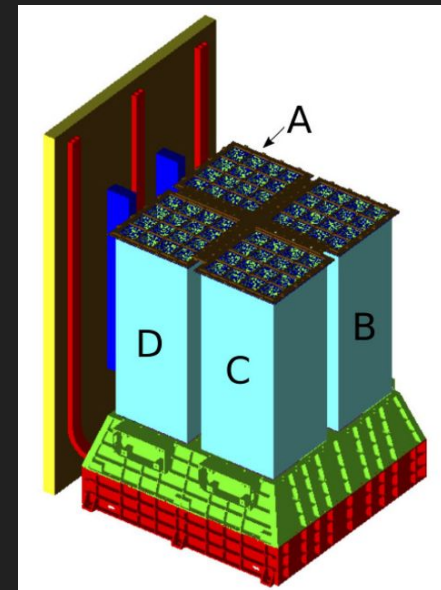
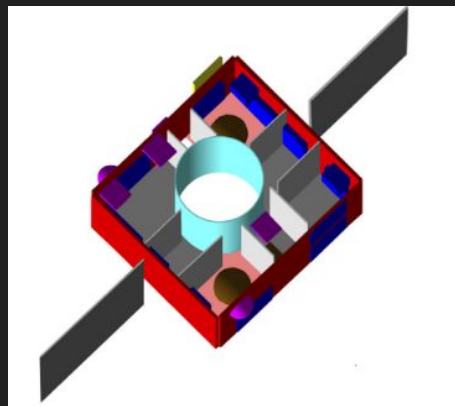
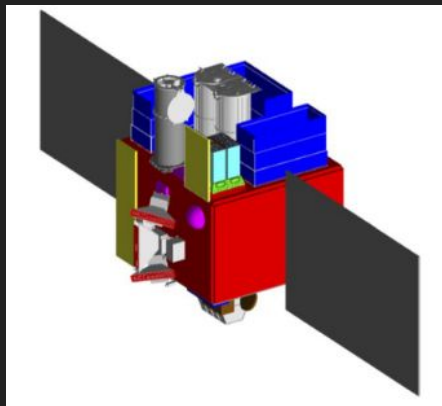


g bright
 high-energy transients such as Gamma-ray bursts and Soft
 gamma repeaters (SGRs)

Slide credits: Mithun NPS

The AstroSat/CZTI mass model

- Full GEANT4 simulation of the AstroSat satellite with CZTI detectors as active volume.
- Simulate off-axis spectral and polarimetric response of CZTI for Gamma-ray Bursts and/or other bright hard X-ray transients.

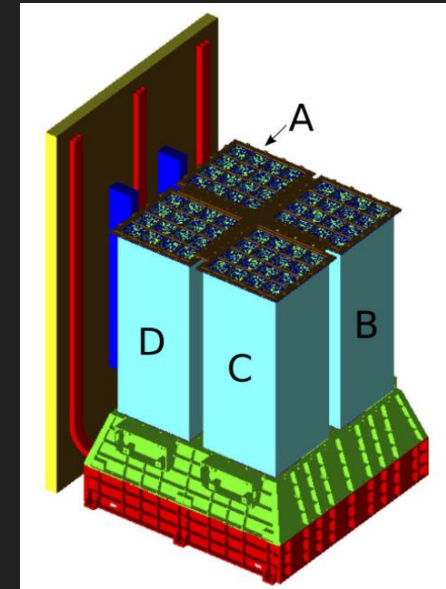


Mate et. al. 2021

The AstroSat/CZTI mass model

- Full GEANT4 simulation of the AstroSat satellite with CZTI detectors as active volume.
- Simulate off-axis spectral and polarimetric response of CZTI for Gamma-ray Bursts and/or other bright hard X-ray transients.

| Instrument | Simulated mass (kg) | Actual mass (kg) |
|-------------------------------|---------------------|------------------|
| CZTI | 41.87 | 50.29 |
| SXT | 44.03 | 57.64 |
| UVIT | 212.58 | 202.06 |
| LAXPC | 350.87 | 389.10 |
| SSM | 71.53 | 71.53 |
| Satellite bus and electronics | 657.22 | 668.47 |
| Total | 1378.10 | 1439.09 |

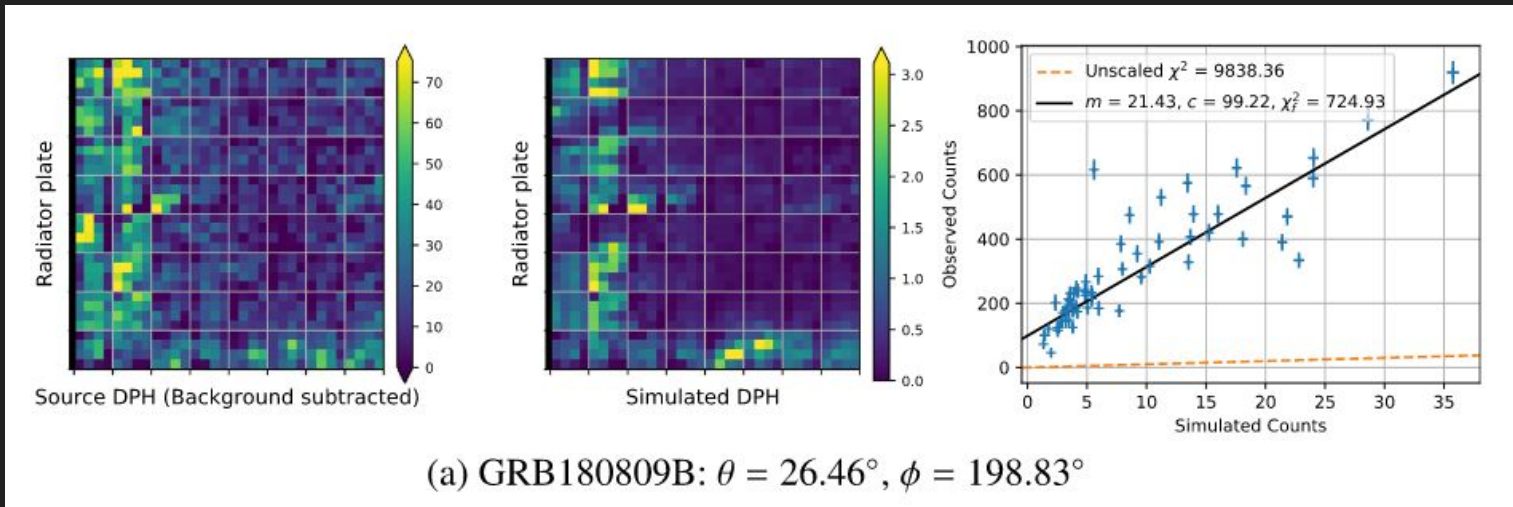


Mate et. al. 2021

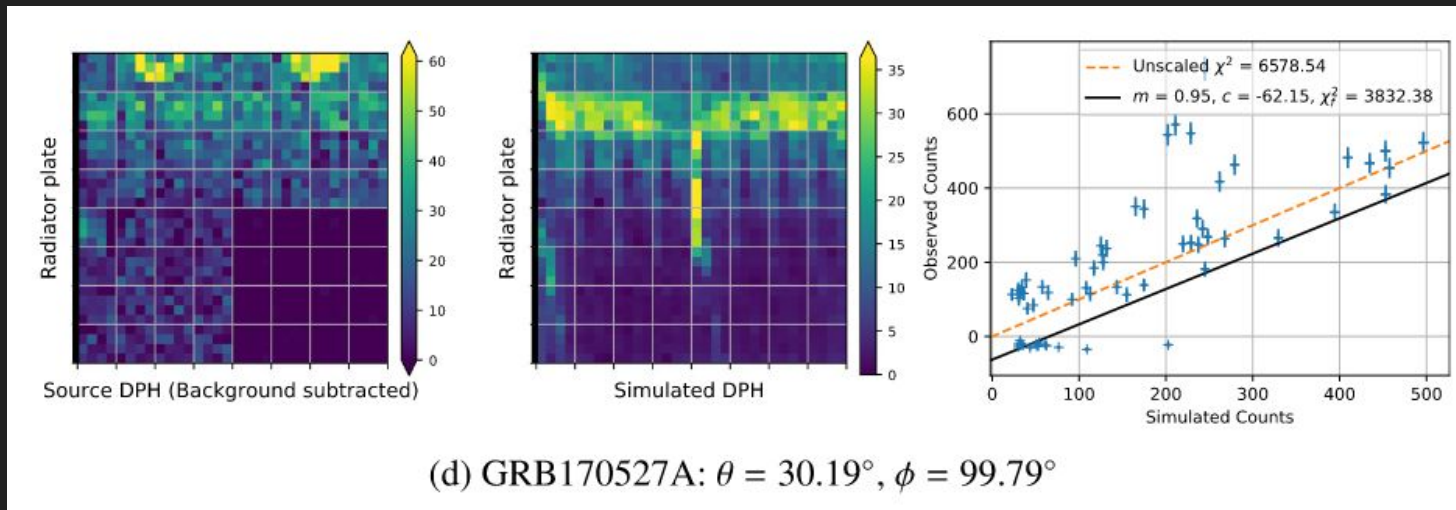
The AstroSat/CZTI mass model: Comparison with data

- No lab calibration with full or partial mass model.
- “Bright” GRBs chosen to verify the “imaging” response.
- Due to many approximation, spectral response depends on direction and energy band.
- In general, good agreement with data for directions above the detector plane

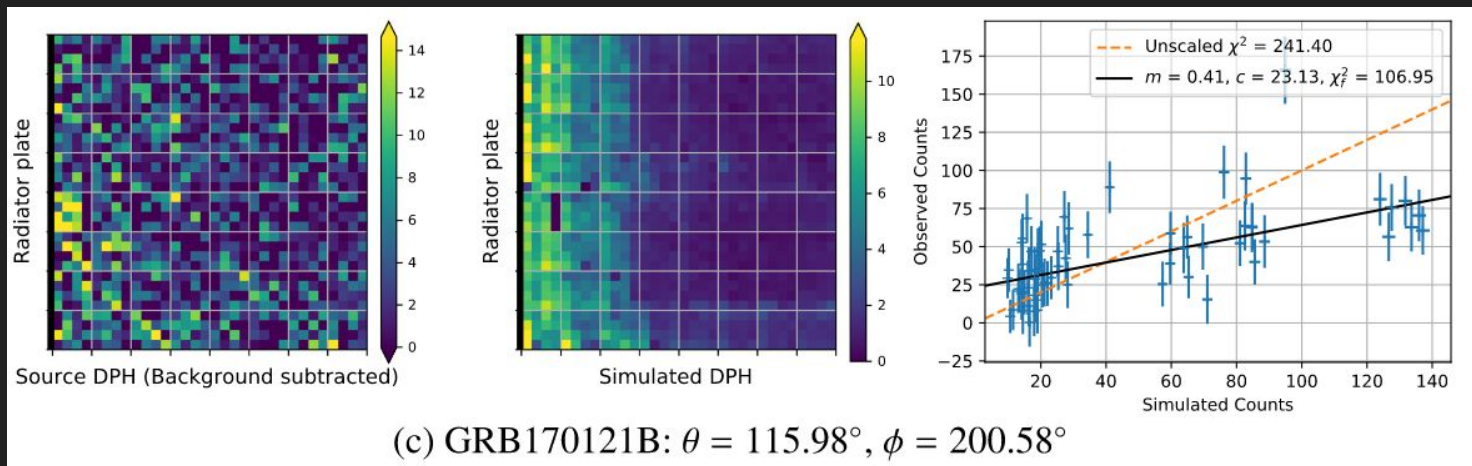
The AstroSat/CZTI mass model: Imaging



The AstroSat/CZTI mass model: Imaging

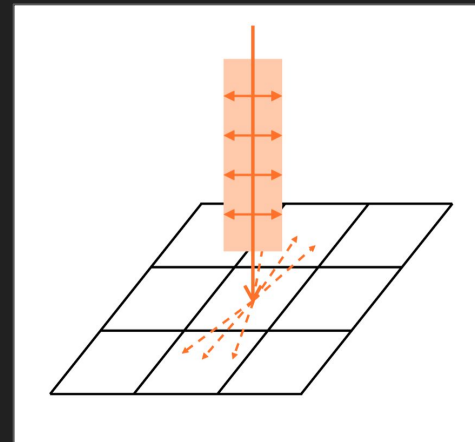


The AstroSat/CZTI mass model: Imaging



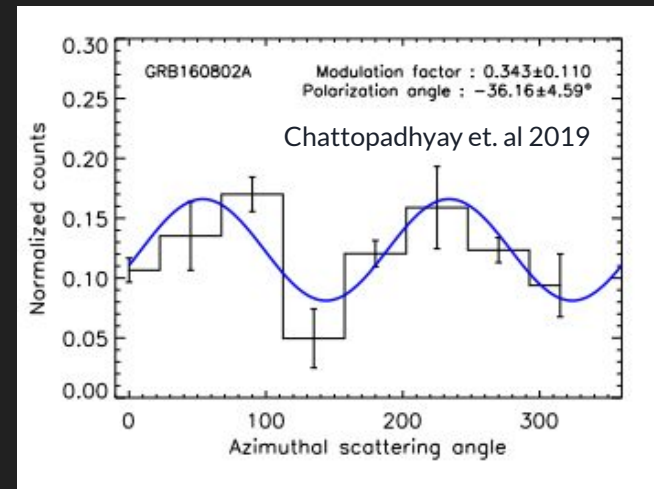
The AstroSat/CZTI mass model: Polarimetry

- Pixelated detector → 8 azimuthal pixel bins surrounding centre scattering pixel.
- Compton scattering dominate in 100 - 400 keV range.
- Polarisation measurement possible by selecting neighboring pixel “Compton” events and fitting the resulting modulation



The AstroSat/CZTI mass model: Polarimetry

- Pixelated detector → 8 azimuthal pixel bins surrounding centre scattering pixel.
- Compton scattering dominate in 100 - 400 keV range.
- Polarisation measurement possible by selecting neighboring pixel “Compton” events and fitting the resulting modulation



The AstroSat/CZTI mass model: Lessons learnt

- Start the modelling at before the launch, most information available easily.
- More systematic approach to writing the GEANT4 code.
- Lab calibrations at different level must for out of FoV angles:
 - Individual detector level
 - Individual instrument level

Thank you