Disappearing Track Searches at the LHC and Future Colliders

1703.09675, in preparation

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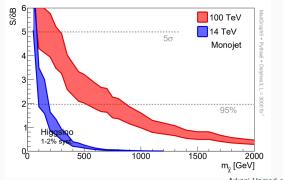
Kavli IPMU, U. Tokyo

Dark Matter and Supersymmetry

- DM is rather a motivation for SUSY
- Suppose Higgsino LSP as an example
 - Same goes for wino and even the minimal DM
- Thermal abundance fixes the mass
 - 1.1 TeV for Higgsino DM
 - It still sounds within the range of colliders

Dark matter @ Colliders

- Basic strategy: $E_T + X$
- Poorer than the thermal mass even for $100 \,\mathrm{TeV}$



Arkani-Hamed et al. 15

• Disappearing tracks from the decay of charged partners

Take a detailed look at SU(2) multiplet

$$\tilde{h} = \begin{pmatrix} \tilde{h}^0 \\ \tilde{h}^+ \end{pmatrix}$$

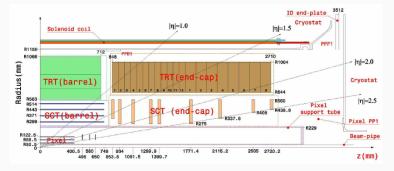
- Charged partners are always there
 - Must be heavier then the neutral one

Mass difference and the lifetime

- No $SU(2) \times U(1)$ breakdown, no mass diff.
 - Mass diff. is much smaller than the thermal mass
- For smaller mass diff., *larger lifetimes* are expected

•
$$c\tau(\chi^+ \to \chi^0 \pi^+) \sim 0.7 \operatorname{cm}\left(\frac{\Delta m}{350 \operatorname{MeV}}\right)^{-3}$$
 for Higgsino

Disappearing tracks in colliders



ATLAS collabolation, arXiv:1110.6191

- A charged partner decays inside the tracker region → *Disappearing* track
 - The *p_T* of π[±] is not large enough to reconstruct its track

Two origin of the mass difference

- Radiative correction $\Delta m \sim \frac{\alpha}{4\pi} v_{\rm EW}$
 - 350 MeV for Higgsino
- Mixing with other particles

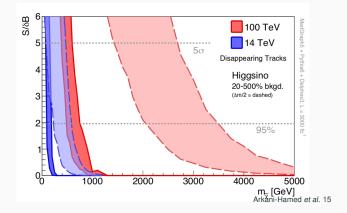
Mixing with other particles

- For Higgsino: bino or wino-like particle
- Roughly, mass diff. \simeq coupling to Higgs
 - The larger the diff. is, the larger the couplings to **SM particles** are
 - If the mass diff. is large enough, direct detection cross section is larger than the ν-floor HF, Nagata, Otono, Shirai 17

Setup for disappearing track search

- Assume only radiative mass difference;
 - $c\tau(\chi^+ \to \chi^0 \pi^+) \sim 0.7 \text{ cm for Higgsino}$
 - (Such "decoupling" DM models are also theoretically motivated like split-type SUSY)

Conventional study

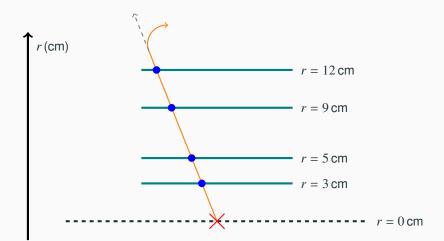


• For Higgsino, even 100 TeV is still not enough!

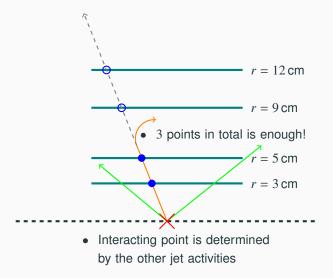
What limits the sensitivity?

- Numerous Higgsinos are there
- Just we can *not* see enough disappearing tracks
- Why? Matter of analysis/tracker

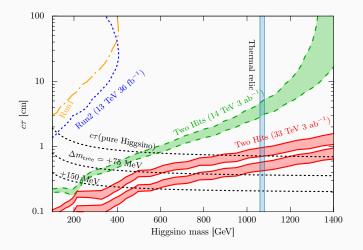
Look inside the tracker



Improvement to 2-point search



Result



Conclusions

- Disappearing track is essential for the DM and SUSY search
- We need to not only increase the energy but also **improve trackers and analysis**
- BG estimation & reduction is very important

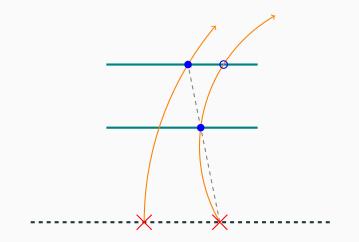
HF, Nagata, Otono, Shirai in prep

Backup

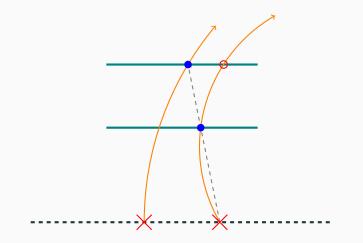
- Misidentification of kink tracks
 - Does not change much for 2-pt strategy
 - Also, the displaced vertex analysis can be used
- Fake disappearing tracks

ATLAS-CONF-2017-017

Fake disappearing tracks



Reduce them by additional points



HF, Nagata, Otono, Shirai in prep