Analyzing signals of compressed spectra in SUSY

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• Compressed spectra in MSSM

- with ^χ⁰₁ LSP.
 with ^G_G LSP.
- Signals & Results
- Conclusions

Compressed spectra

- Compressed spectra has relatively closely spaced sparticles.
- Such a spectrum produces softer jets, leptons and missing transverse energy → may not pass the signal selection criterions leading to weak bounds on such spectra.



- Most studies in this direction focus on simplified models with a squark or gluino compressed with the LSP and rest of the spectrum decoupled.



Case I: Compression in MSSM

Non-universal SUSY breaking at high scale could give rise to a compressed spectrum with masses of gluinos, squarks, sleptons close to the $\tilde{\chi}_1^0$ LSP.





• At 1-loop order, the lightest CP-even Higgs mass is:

$$m_h^2 = m_z^2 \cos^2(2\beta) + \frac{3m_t^4}{4\pi^2 v^2} \ln\left(\frac{M_S^2}{m_t^2} + \frac{X_t^2}{M_S^2} - \frac{X_t^4}{12M_S^2}\right)$$

where $M_{S}=\sqrt{m_{\widetilde{t}_{1}}m_{\widetilde{t}_{2}}},~X_{t}=A_{t}-\mu\coteta$

Require atleast one heavy stop as well as large mixing X_t in the stop sector to fit $122 < m_h < 128$ GeV.

- LEP lower bound on the lightest chargino mass, i.e, $m_{\widetilde{\chi}^1_+} > 103.5~{\rm GeV}.$
- Constraints from branching ratios of rare decays such as $BR(b \rightarrow s\gamma)$ and $BR(B_s \rightarrow \mu\mu)$.
- For parameter scans, we have considered only the upper bound on dark matter relic density, i.e, $\Omega h^2 < 0.138$.
- Constraints from direct detection cross-sections (σ_{SI}) from LUX data.



$$\Delta M = m_{S} - m_{\widetilde{\chi}_{1}^{0}},$$

$$S \in [\widetilde{g}, \widetilde{t}_{2}, \widetilde{b}_{2}, \widetilde{\tau}_{2}, \widetilde{\chi}_{2}^{0}, \widetilde{\chi}_{1}^{\pm}]$$

Heavy spectra and large μ parameter facilitates compression (ΔM) in the spectra.



 $\widetilde{\chi}_1^0$ LSP and cold dark matter candidate, satisfies observed thermal relic density.

Low μ values, (~ 2 TeV) strongly constrained from direct detection cross-section data from LUX due to large bino-higgsino mixing.

Case II: Compression in MSSM + \widetilde{G} LSP

• We focus on a compressed MSSM spectra with a bino-like $\tilde{\chi}_1^0$ NLSP extended with a keV gravitino LSP.



• Presence of light \widetilde{G} relaxes DM constraints on $\widetilde{\chi}_1^0$.

Branching ratios of \widetilde{g}

•
$$\Gamma(\widetilde{g} \to g\widetilde{G}) \propto m_{\widetilde{g}}^5 m_{\widetilde{G}}^{-2}$$

• Competing decay modes: $\widetilde{g} o g \widetilde{G}$, $\widetilde{g} o q ar{q} \widetilde{\chi}_1^0$



- Small compression ($\Delta M \sim 50 \text{ GeV}$) and $m_{\widetilde{G}} \sim 1 \text{ keV}$: BR($\widetilde{g} \rightarrow q \overline{q} \widetilde{\chi}_1^0$) > BR($\widetilde{g} \rightarrow g \widetilde{G}$).
- Large compression ($\Delta M \sim 10$ GeV) and $m_{\widetilde{G}} \sim 1$ keV: BR($\widetilde{g} \rightarrow q \overline{q} \widetilde{\chi}_{1}^{0}$) < BR($\widetilde{g} \rightarrow g \widetilde{G}$).
- For sub-keV \widetilde{G} : BR($\widetilde{g} \to g \widetilde{G}$) dominant.

Branching Ratios of \widetilde{q}_L , \widetilde{q}_R

•
$$\Gamma(\widetilde{q}
ightarrow q\widetilde{G}) \propto m_{\widetilde{q}}^5 m_{\widetilde{G}}^{-2}$$

• Competing decay modes: $\widetilde{q} o q \widetilde{G}$, $\widetilde{q} o q \widetilde{\chi}_1^0$



- Small compression and $m_{\widetilde{G}} > 1 \text{ eV} : \text{BR}(\widetilde{q} \to q \widetilde{\chi}_1^0) > \text{BR}$ $(\widetilde{q} \to q \widetilde{G}).$
- Large compression and $m_{\widetilde{G}} > 1 \text{ eV}$: $\mathsf{BR}(\widetilde{q} \to q \widetilde{\chi}_1^0) < \mathsf{BR}(\widetilde{q} \to q \widetilde{G})$.
- For sub-eV \widetilde{G} : BR $(\widetilde{q} \rightarrow q\widetilde{G})$ dominant.

The bino-like \$\tilde{\chi_1}^0\$ NLSP decays dominantly to \$\gamma\$ and \$\tilde{G}\$ and a small fraction to \$Z + \$\tilde{G}\$. This leads to extremely hard photons and large \$\mathcal{E}_T\$.



 These hard photon associated signals can be very effective to probe a heavy compressed SUSY spectra with a light gravitino as there would be rarely any Standard Model events with such hard photons. We consider the following signals at $\sqrt{s}=13~{\rm TeV}$ for our study :

- Monojet + $\not\!\!\!E_T$
- Multijets ($\geq 2 j$) + $\not \in_T$

Simulation Details

SUSY signal: $\tilde{q}\tilde{g}$, $\tilde{q}\tilde{q}$, $\tilde{q}\tilde{q}^*$, $\tilde{g}\tilde{g} + \leq 2$ partons.

- Spectrum Generator: SPheno
- Madgraph5 → Pythia6 → Delphes-v3 for event generation, showering and detector simulation.
- MLM matching with showerKT performed duly with QCUT = 120 GeV (SUSY), 30-50 GeV (SM).

Background:

- MSSM: $Z + \leq 4j$, $W + \leq 4j$, $QCD (\leq 4j)$, $t\overline{t} + \leq 2j$, $t + \leq 3j$, $ZZ + \leq 2j$, $WZ + \leq 2j$.
- $MSSM + \tilde{G}$: from existing ATLAS study, i.e. ATLAS-CONF-2016-066.

Signal cross-sections computed at NLO (NLO+NLL) for MSSM (MSSM+ \tilde{G}) using Prospino (NLL-Fast). Background cross-section upto NLO (using Madgraph5) for MSSM.

Benchmarks

Parameters	BP1	BP2
A_t	-1535.1	2300.0
μ	3000.0	3000.0
aneta	23.9	20.0
$m_{\widetilde{g}}$	1497.4	1534.7
$m_{\widetilde{q}_L}$	1452.3	1524.5
$m_{\widetilde{q}_R}$	1451.3	1520.8
$m_{\tilde{t}_1}$	1330.6	1507.6
$m_{\tilde{t}_2}$	1509.0	1686.6
$m_{\widetilde{b}_1}$	1407.4	1521.9
$m_{\tilde{b}_2}$	1494.5	1619.5
$m_{\widetilde{\chi}_1^0}$	1323.9	1496.3
$m_{\widetilde{\chi}_{2}^{0}}$	1342.9	1559.0
$m_{\widetilde{\chi}_1^\pm}$	1342.9	1559.1
m_h	122.5	122.4
Ωh^2	0.113	0.105
$\sigma_{\it SI} imes 10^{11}~{ m (pb)}$	4.65	0.13
ΔM_i (GeV)	173.5	38.4

 $\Delta M_i = m_{\mathcal{S}} - m_{\widetilde{\chi}^0_1}$, where $\mathcal{S} \in [\widetilde{q}, \widetilde{g}]$

Signal		Cross-section after cuts (fb)					
Benchmark	Production	Preselection	M _{Eff}	₹⊤	$\not \in_T / \sqrt{H_T}$	∉ _T /M _{Eff}	
Points	cross-section(fb)		> 800 GeV	$> 160 { m ~GeV}$	$> 15 GeV^{1/2}$	> 0.35	
BP1	126.93	59.72	20.74	19.84	9.99	9.93	
BP2	95.58	12.45	6.34	6.24	4.72	4.68	
SM Background	2.0E+08	253042	2833	8.85	1.36	1.35	

Multijets $+ \not \in_T$ cross-section for signal and background (at NLO).

(Preselection: $p_T(j_1) > 130 \text{ GeV}, p_T(j_2) > 80 \text{ GeV}, \Delta \phi(j_{1/2}, \not \!\!\! E_T) > 0.4)$

Signal		Cross-section after cuts (fb)			
Benchmark	Production	Preselection	<i>∉_T</i> >160 GeV	$M_{Eff} > 800 \text{ GeV}$	
Points	cross-section(fb)				
BP1	126.93	12.06	8.22	0.88	
BP2	95.58	7.48	6.20	1.63	
SM background	2×10 ⁸	46254	2602	0.938	

Monojet $+ \not \in_T$ cross-section for signal and background (at NLO).

	Luminosity (in fb^{-1}) for 3σ excess		
Signal	BP1	BP2	
$Multijets \ (\geq 2 \ j) + \not\!\!\! E_{T}$	123	558	
$Monojet + \not\!\!\! E_T$	10926	3204	

- Multijet + met searches still more efficient to look for compressed scenarios than traditional monojet + met channels.
- However both are viable modes of discovery for compressed spectra at the Run 2 of LHC.

Signals and Results (for MSSM extended with a \widetilde{G} LSP)

We consider the following signal:

•
$$\geq 1\gamma + > 2j + \not \in_T$$

Experimental collaborations (ATLAS-CONF-2016-66) consider signal events coming from gluino pair production only, assuming rest of the sparticles decoupled, ruling out $m_{\widetilde{g}} \leq 1.95$ TeV for $m_{\widetilde{\chi}_1^0} \sim 1.8$ TeV.



- However for a compressed spectra, presence of closely spaced sparticles lead to added contributions to the same final state.
- Thus, the limits on sparticles are stronger for a compressed spectra.

Using the ATLAS analysis for $\geq 1\gamma + > 2$ jets $+ \notin_T$ and SM background estimates at 13.3 fb^{-1} , mass bounds significantly increase for a compressed spectra, i.e., $m_{\tilde{g}/\tilde{q}} \geq 2.5$ TeV.

Hard photons are a characteristic feature of both compressed and uncompressed spectra.

Benchmarks

	Compressed spectra		Uncompressed spectra
Parameters	C4	C5	U2
A _t	-3750	-3197	2895
μ	4000	3500	3000
aneta	6	25	15
M _A	1800	2500	2500
$m_{\widetilde{g}}$	2783	2562	2102
$m_{\tilde{q}_L}$	2753	2571	4721
$m_{\widetilde{q}_R}$	2751	2574	4742
$m_{\tilde{t}_1}$	2625	2532	4678
$m_{\tilde{t}_2}$	2863	2718	4765
$m_{\tilde{b}_1}$	2778	2594	4558
$m_{\tilde{b}_2}$	2846	2677	4744
$m_{\widetilde{\chi}_1^0}$	2585	2526	1191
$m_{\widetilde{\chi}_{0}^{0}}^{\chi_{1}^{0}}$	2724	2619	2383
$m_{\widetilde{\chi}_1^{\pm}}$	2724	2619	2382
m _h	124	125	125
ΔM_i	198	48	911

Using the existing ATLAS analysis cuts (using hard cuts on photon p_T and $\not \in_T$) and SM background estimates at 13.3 fb^{-1} for the same final state:

Signal		Cross-section (in fb) after cuts					
Benchmark	Production	$p_T(\gamma_1)$	$N_j > 2$	$\Delta \phi(j_{1/2}, \not \in_T)$	$\Delta \phi(\gamma_1, \not \in_T)$	Ĕτ	M _{Eff}
Points	cross-section(fb)	> 400	$N_{I} = 0$	> 0.4	> 0.4	> 400	> 2000
C4	0.21	0.15	0.12	0.08	0.08	0.08	0.07
C5	0.49	0.34	0.15	0.13	0.13	0.12	0.11
U2	0.20	0.13	0.12	0.10	0.09	0.08	0.08

we compute the required luminosity for some benchmarks:

Signal	Luminosity ${\cal L}$ (in fb $^{-1}$) for		
	$S = 3\sigma$	$\mathcal{S}=5\sigma$	
C4	176	489	
C5	79	219	
U2	139	385	

Using $p_T(\gamma)$, $p_T(j)$ and N_j , a set of new kinematic variables identified which act as a discriminant for a compressed and uncompressed spectra with similar event rates:

$$r'_1 = N_j r_1, r'_2 = N_j r_2, \text{ where } r_1 = \frac{p_T(j_1)}{p_T(\gamma_1)}, r_2 = \frac{p_T(j_2)}{p_T(\gamma_1)}$$

For C4, C5: $r'_1 \sim 0.2 - 0.5$, $r'_2 \sim 0.1 - 0.3$ while for U2: $r'_1 \sim 4$, $r'_2 \sim 2.5$.



Conclusions

- Compressed spectra in MSSM with [~]χ⁰₁ LSP gives rise to multiple jets and ∉_T which fare better over traditional monojet and ∉_T signal.
- We further extend the MSSM spectra with a light G̃. Presence of a light G̃ relaxes DM constraints on the MSSM part of the spectrum.
- With existing data, exclusion limits on colored sparticles improve significantly for a compressed spectra in photonic channels.
- Hard photons are a characteristic feature of both compressed and uncompressed spectra. Simlar event rates may be obtained for compressed and uncompressed spectra.
- Difference in compression reflected in kinematic variables involving hardness of photons, jet and jet multiplicity to distinguish such spectra.
- For sub-keV gravitinos, \widetilde{G} associated decay mode of the sparticles become relevant and alternate channels of interest are multi-jets and missing energy signals requiring $\sim 1000 \, fb^{-1}$ for observing a 3σ excess at LHC.

Thank You

Backup

Parameters	Ranges
M_1, M_2, M_3	(100, 2500) GeV
A_t	(-3000, 3000) GeV
taneta	(2, 50)
$M_L = M_R$	$(M_1, M_1 + 200) \text{ GeV}(\text{if } M_1 < M_2)$
	$(M_2, M_2 + 200) \text{ GeV}(\text{if } M_2 < M_1)$

Table : Ranges of the relevant parameters for the scan. M_1 , M_2 , M_3 are the gaugino mass parameters, varied in the same range but independent of each other. M_L and M_R are the left-handed and the right-handed soft mass parameters of squarks and sleptons.

Differential distributions of kinematic variables

