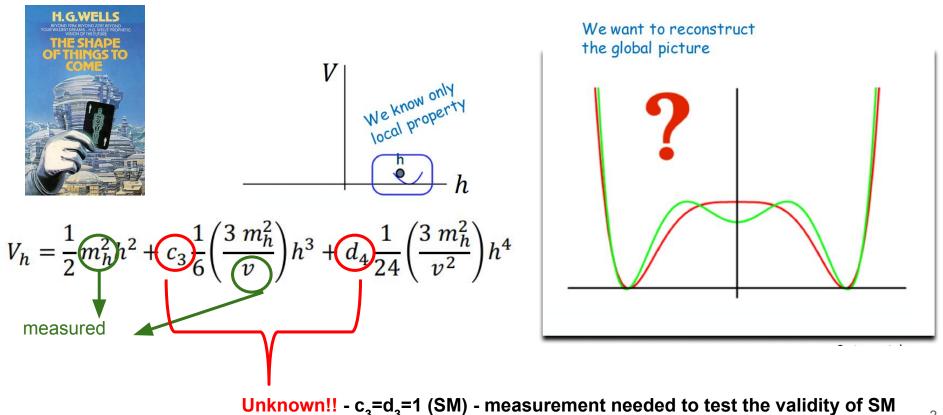
# On the Validity of the Effective Potential and the Precision of Higgs Self Couplings

# Bithika Jain IFT-UNESP &ICTP-SAIFR

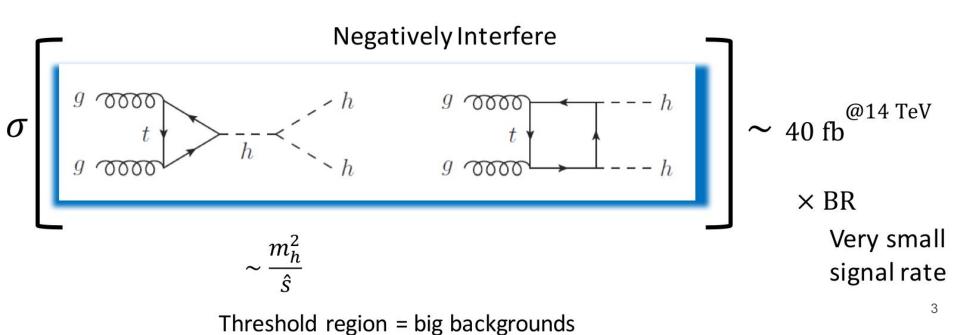
Work done with Minho Son, Seung J Lee Based on arXiv:1709.03232

# The Shape of Things to come



# Can Higgs potential be tested by self coupling measurements?

- → Cubic higgs self coupling
- → HH production via gluon fusion is the best channel



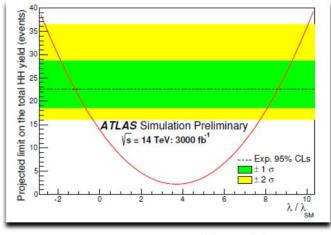
Nothing to discuss about Higgs self coupling using current data

#### Cubic coupling @ HL LHC using 3000/fb

- ✓ Very tough process
  - $b\bar{b}\gamma\gamma$  (0.264%)

Seems to be the best channel so far

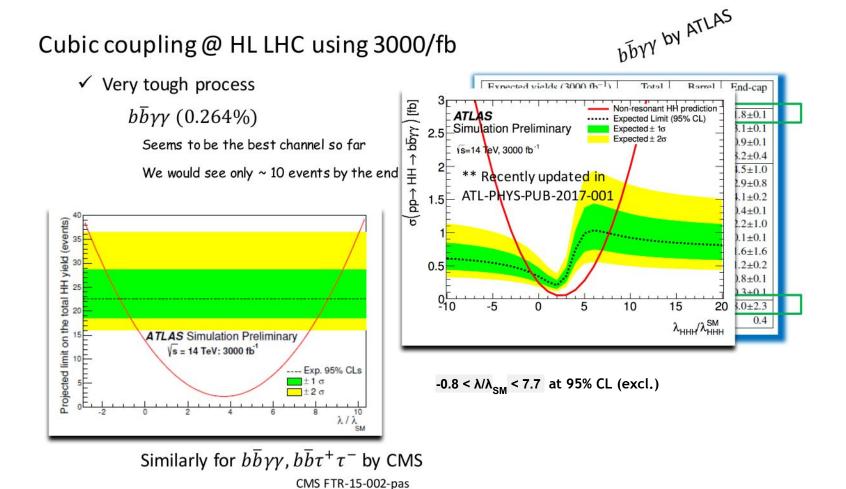
We would see only ~ 10 events by the end of HL LHC



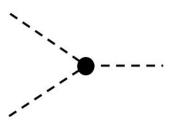
CI	Expected yields (3000 fb <sup>-1</sup> )	<mark>15-РИ</mark> ББүү <sup>К</sup>	3-2014 DN ATLA	4-019 S	
ſ	Expected yields (3000 fb <sup>-1</sup> ) Samples	Total	Barrel	End-cap	
	$H(b\bar{b})H(\gamma\gamma)(\lambda/\lambda_{SM}=1)$	8.4±0.1	6.7±0.1	1.8±0.1	
Т	$H(bb)H(\gamma\gamma)(\lambda/\lambda_{SM}=0)$	13.7±0.2	$10.7 \pm 0.2$	3.1±0.1	
L	$H(b\bar{b})H(\gamma\gamma)(\lambda/\lambda_{SM}=2)$	4.6±0.1	$3.7 \pm 0.1$	0.9±0.1	
L	$H(b\bar{b})H(\gamma\gamma)(\lambda/\lambda_{SM}=10)$	36.2±0.8	27.9±0.7	8.2±0.4	
L	bbγγ	9.7±1.5	5.2±1.1	4.5±1.0	
L	c̄cγγ	7.0±1.2	4.1±0.9	2.9±0.8	
L	b̄bγj	8.4±0.4	4.3±0.2	4.1±0.2	
L	bbjj	$1.3 \pm 0.2$	$0.9 \pm 0.1$	0.4±0.1	
L	јјуу	7.4±1.8	5.2±1.5	2.2±1.0	
L	$t\bar{t} \ge 1$ lepton)	$0.2 \pm 0.1$	$0.1 \pm 0.1$	0.1±0.1	
L	tīy	$3.2 \pm 2.2$	1.6±1.6	1.6±1.6	
	$t\bar{t}H(\gamma\gamma)$	6.1±0.5	4.9±0.4	1.2±0.2	
	$Z(b\bar{b})H(\gamma\gamma)$	2.7±0.1	1.9±0.1	0.8±0.1	
	hhH(yy)	12+01	1.0+0.1	03+01	
	Total Background	47.1±3.5	29.1±2.7	18.0±2.3	
	$S/\sqrt{B(\lambda/\lambda_{SM}=1)}$		1.2		

an exclusion at 95% C.L. of BSM models with  $\lambda_3 / \lambda_{3,SM} \leq -1.3$  and  $\lambda_3 / \lambda_{3,SM} \geq 8.7$ 

Similarly for  $b\bar{b}\gamma\gamma$ ,  $b\bar{b}\tau^+\tau^-$  by CMS CMS FTR-15-002-pas Nothing to discuss about Higgs self coupling using current data



#### cubic coupling @ 100 lev using 30/ab



 $\checkmark b\bar{b}\gamma\gamma$  becomes Golden channel at 100 TeV pp collider

40 imes enhanced xsec due to PDF

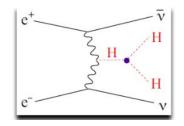
Barr, Dolan, Englert, Lima, M.Spannowsky 15' Contino, Azatov, Panico, SON 15' H. He, J. Ren, W. Yao 16' Physics at 100 TeV

Contino, Panico, Papaefstathiou, Selvaggi, SON in progress

#### $\sim$ 3.4 % is possible with 30 $\mathrm{ab^{-1}}$

Physics at 100 TeV, arXiv:1606.09408

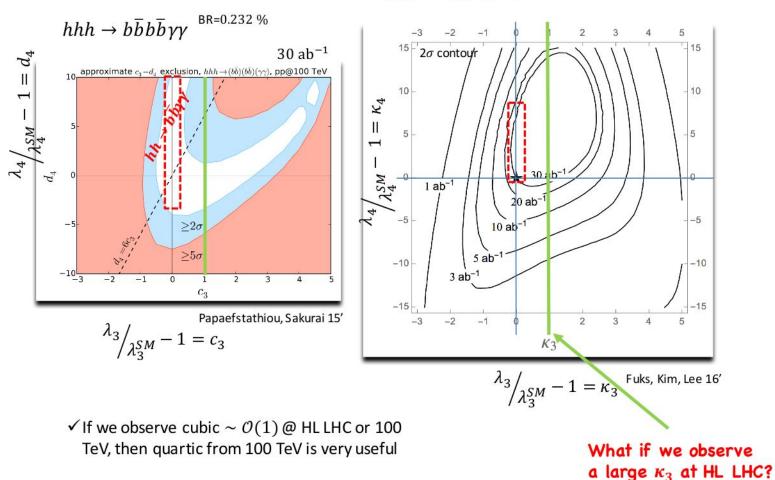
✓ ILC via VBF at 1 TeV 8 ab<sup>-1</sup> ~10%



J. Tian, LC-REP-2013-003 M. Kurata, LC-REP-2014-025 C. Duerig, Ph.D. thesis at DESY, 2016 HHàbbbb, bbWW\* combination

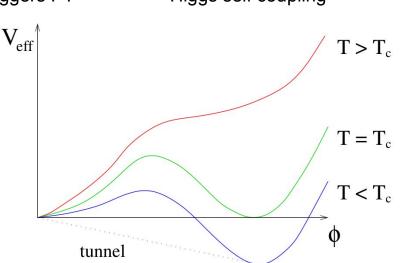
#### Quartic coupling @ 100 TeV

 $hhh \rightarrow b\overline{b}b\overline{b}\tau^+\tau^- {}^{BR\,=\,6.46\,\%}$ 

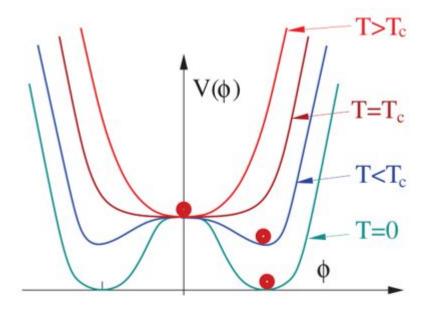


Dynamics of EWPT  
Mass param, 
$$\mu^2$$
  
 $V_{\rm EFF}(\varphi, T) = D(T^2 - T_0^2)\varphi^2 - (ET + e)\varphi^3 + \bar{\lambda}\varphi^4 + \cdots$   
Higgs background  
field  
Higgs sackground  
Higgs self coupling

- $\rightarrow$  In SM, e =0 while in BSM its nonzero
- → E is generated at loop level (SM & BSM)
- → (ET+e) and  $\bar{\lambda}$  decide the nature of EWPT
- → II order PT both E and e vanish
- ➔ I order PT occurs with E>0 and/ro e>0
- → I order PT has interesting consequences
- → Eg: baryogenesis (more about this later)



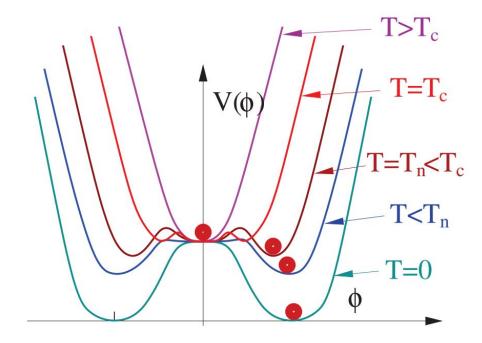
In SM , with  $m_{\mu}$  = 125 GeV , Numerical simulations suggest a crossover



- → Field, $\Phi$  develops a vev when T< Tc
  - Φ rolls smoothly from zero to nonzero value
- → Such transitions are always in thermal equilibrium
- System remembers nothing about the unbroken phase
- → Cosmologically uninteresting

K. Kajantie, et al. PRL 77 (1996) 2887, F. Karsch, et al. NPPS 53 (1997) 623, Y. Aoki et al., PRD 56 (1997) 3860 M. Gurtler et al., PRD 56 (1997) 3888

In BSM, we can have first order phase transitions



- → @Tc all 3 minima become equal
- Φ is still trapped in the origin but secondary minima is energetically favorable
- → At Tn potential barrier is small
- $\rightarrow$   $\Phi$  tunnels to true vacuum
- → Out of equilibrium processes
- → Relevant for cosmology

# Phase transition model classes

I. Thermally Driven - new particles in early universe plasma

- barrier formation via thermal loop effects associated with bosonic zero modes
- IA. barrier from light scalar via thermal cubic terms
- IIB. from heavy particles with large coupling to Higgs field

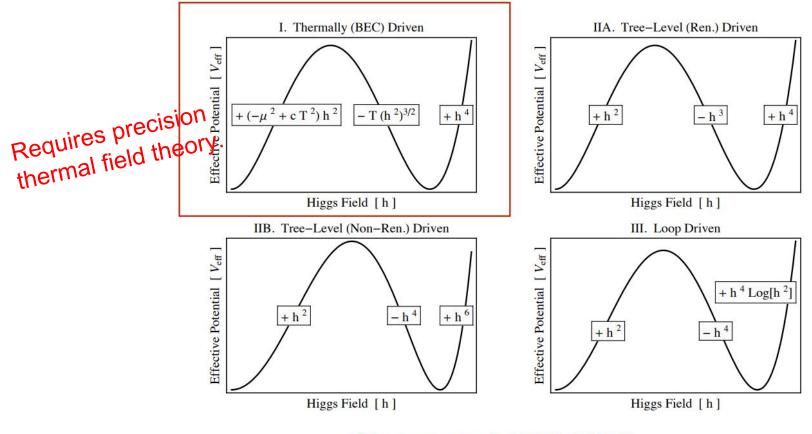
#### II. Tree-Level Driven - BSM physics

*IIA. Renormalizable Operators* - effective h<sup>3</sup> operator from extra scalars which acquire nonzero vevs during EWPT eg:- SM+ real singlet
 *IIB. Non-Renormalizable Operators.* - higher dimensional operators like h<sup>6</sup>

#### III. Loop Driven

- Eq:- quartic correction of form h<sup>4</sup>ln h<sup>2</sup> - which competes with unstable -h<sup>4</sup> term

# Extensions to Higgs



Chung et.al. 1209. 1819

# Baryogenesis

 $\rightarrow$  Evidence from cosmology:

y: 
$$\frac{n_B}{s} = (8.59 \pm 0.11) \times 10^{-11}$$
 (Planck)

- → Sakharov's 3 conditions (1967), for baryogenesis
  - Baryon number violation
  - C and CP violation
  - Out of equilibrium
- → EW baryogenesis is one of the potential solutions
- → Need new physics because in SM:
  - EW phase transition is a crossover, instead of 1 st order
  - CP violation is too small



# Criteria for baryon number violation

Baryon number can be violated by non-perturbative EW processes

Unstable solutions of  $S_{FFF}$  i.e sphalerons interpolate b/w topologically distinct vacua

Sphaleron energy depends on shape of potential away from the minimum

#### Litmus test for its global structure !!!

For successful EWBG, EW sphaleron processes are out of equilibrium in broken phase - washout avoidance condition

$$\frac{\mathbf{v} (\mathbf{T}_{PT})}{\mathbf{T}_{PT}} \ge 1 \times \left(\frac{\mathbf{E}_{sph,0}}{9 \text{ TeV}}\right)^{-1} \in [0.6, 1.4]$$
 Strong first order PT

No unique value  $\Rightarrow$  adds to uncertainty on Higgs self coupling

# Testability

→ LHC is running!



- → What's the sensitivity of HL-LHC, 100 TeV pp colliders, and future e+ ecolliders to the region of parameter space where SFOPT is allowed?
- → Gravitational waves: Bubble collisions

# Overview of effective potential at Finite T

The effective potential

Truncated Full Dressing (TFD): we call this is prescription A →: thermal mass IL is still obtain the prescription A

 $\longrightarrow$ : thermal mass  $\Pi_i$  is still obtained in the high-T approximation

$$V_{eff} = V_{tree} + V_{CW}[m_i^2(h) + \Pi_i] + V_T[m_i^2(h) + \Pi_i]$$

$$V_T = \sum_{i=B,F} (-1)^{F_i} \frac{g_i T^4}{2\pi^2} \int_0^\infty dx \ x^2 \log \left[ 1 \mp \exp\left(-\sqrt{x^2 + \frac{(m_i^2(h) + \Pi_i)}{T^2}}\right) \right]$$
$$m_i^2(h) = m^2 + \operatorname{coupling} \times h^2$$
For  $v_c \gtrsim T_c$  and  $\gtrsim \mathcal{O}(1)$  coupling, integral needs to be exactly evaluated

#### ✓ Validity of High-T approximation/Validity of perturbation

: not rigorously addressed in most literature in the context of BSM physics

Curtin, Meade, Ramani 16' for a recent discussion

We do not intend to clarify this problem in this talk, but just point out a few observations in the process of reproducing others.

In the High-T approximation  

$$V_{eff} = V_{tree} + V_{CW} + V_T + V_{ring}$$

$$V_{cw} = \sum_{i=t,W,Z,h,G,\cdots} (-1)^{F_i} \frac{g_i}{64\pi^2} \left[ m_i^4(h) \left( \log \frac{m_i^2(h)}{m_i^2(v)} - \frac{3}{2} \right) + 2m_i^2(h)m_i^2(v) \right]$$

$$V_T = \sum_{i=B,F} (-1)^{F_i} \frac{g_i T^4}{2\pi^2} J_{B/F} \left( \frac{m_i^2(h)}{T^2} \right)$$

$$V_{ring} = \sum_{i=\text{bosons}} \frac{\bar{g}_i T}{12\pi} \left[ m_i^3(h) - \left( m_i^2(h) + \Pi_i(T) \right)^{\frac{3}{2}} \right]$$

$$x^2 = \frac{m^2}{T^2} \ll 1$$

$$J_{B/F}(x^2) = \int_0^\infty dy \ y^2 \log[1 \mp \exp(-\sqrt{y^2 + x^2})]$$

$$J_{F}(x^2) \approx \frac{7\pi^4}{360} - \frac{\pi^2}{24}x^2 - \frac{1}{32}x^4 \log\left(\frac{x^2}{c_f}\right)$$

The effective potential

Truncated Full Dressing (TFD):  $Ne^{call this}$  prescription A Prescription A Prescription A Prescription A Prescription A Prescription A

$$V_{eff} = V_{tree} + V_{CW}[m_i^2(h) + \Pi_i] + V_T[m_i^2(h) + \Pi_i]$$

$$\begin{aligned} V_T(m_i^2(\phi), T) &= \sum_i (-1)^{F_i} \frac{g_i T^4}{2\pi^2} J_{B/F}\left(\frac{m_i^2(\phi)}{T^2}\right) \\ J_B(\alpha^2) &\equiv J_B(\alpha^2; n) = -\sum_{k=1}^n \frac{1}{k^2} \alpha^2 K_2(\alpha k) , \\ J_F(\alpha^2) &\equiv J_F(\alpha^2; n) = -\sum_{k=1}^n \frac{(-1)^k}{k^2} \alpha^2 K_2(\alpha k) , \end{aligned}$$

✓ Validity of High-T approximation/Validity of perturbation

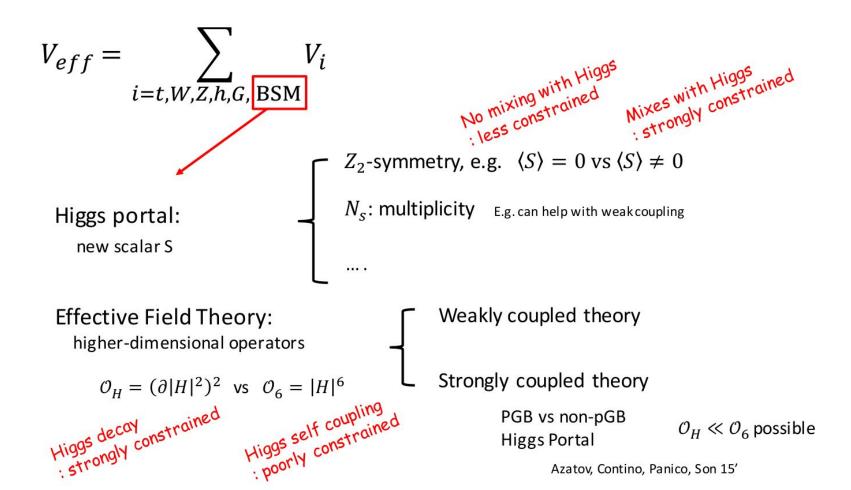
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Curtin, Meade, Ramani 16' for a recent discussion

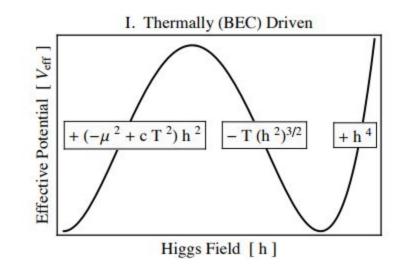
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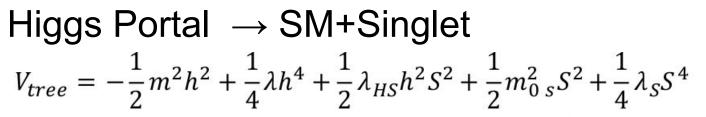
## **Benchmark Scenarios**

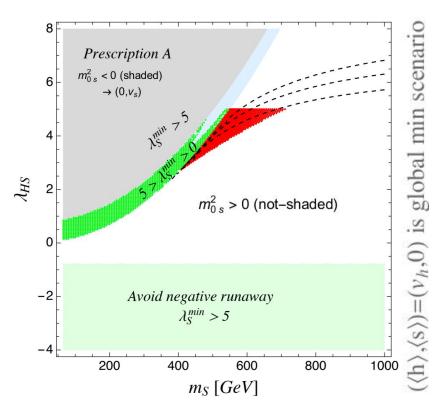
#### Most commonly considered frameworks



# Higgs portal







Based on naïve criterion, existence of degenerate vacua, with  $v_c/T_c > 1$  \* Note cutoff  $\lambda_{HS} > 5$  by hand

 One-step strong 1<sup>st</sup> phase transition (dotted RED)

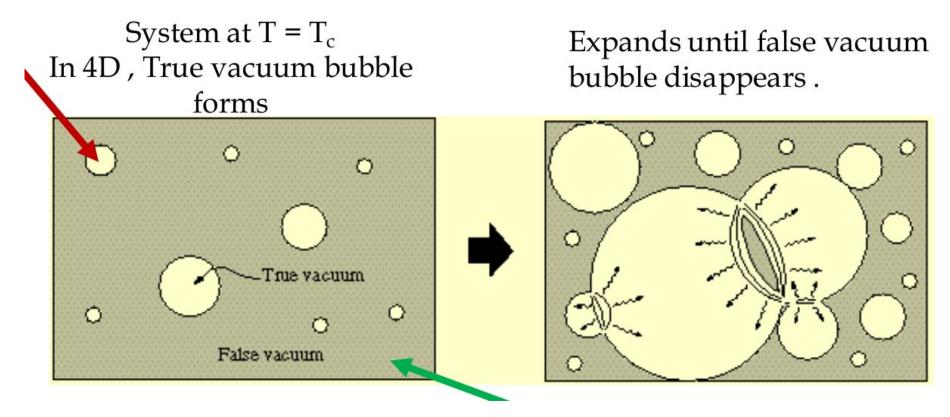
$$V(0,0) \rightarrow V(v,0)$$
 ,  $\langle S \rangle = 0$ 

2. Two-step strong 1<sup>st</sup> phase transition (dotted GREEN)

 $V(0,0) \rightarrow V(0,v_s) \rightarrow V(v,0)$ 

$$V(0, v_s) > V(v, 0) \rightarrow$$
$$\lambda_s > \lambda_s^{\min} \equiv \lambda \frac{m_{0s}^4}{m^4} = \frac{2(m_s^2 - v^2 \lambda_{Hs})^2}{m_h^2 v^2}$$

At low temperatures, or at the interior of a bubble in the cooling universe, there exists a "true vacuum" (broken phase)



At high temperatures, or far away from the bubble, "false vacuum"

# Bubble nucleation

As the Universe continues to cool, bubbles of the broken-minimum phase are nucleated.

Nucleation probability per unit time per unit volume at temperature T is given by

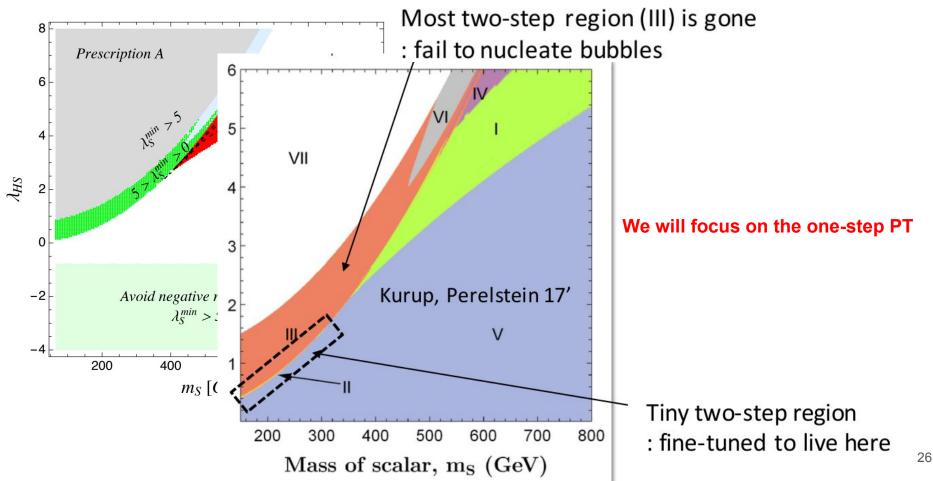
$$P \sim T^4 \exp(-S_3/T)$$

 $S_{3}$  is the bubble action

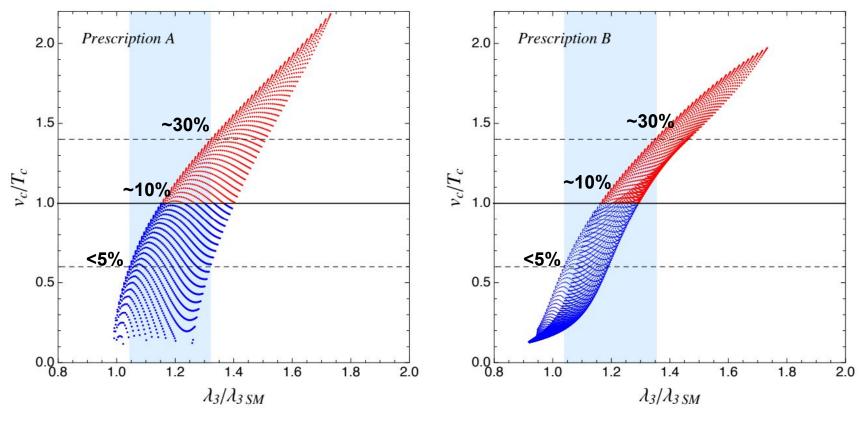
Nucleation temperature  $T_N$  is the temperature at which the nucleation probability per Hubble volume becomes of order one; for electroweak phase transition, this corresponds to  $S_3/T_N \approx 100$ .

### To avoid washout $v(T_N)/T_N > 1$

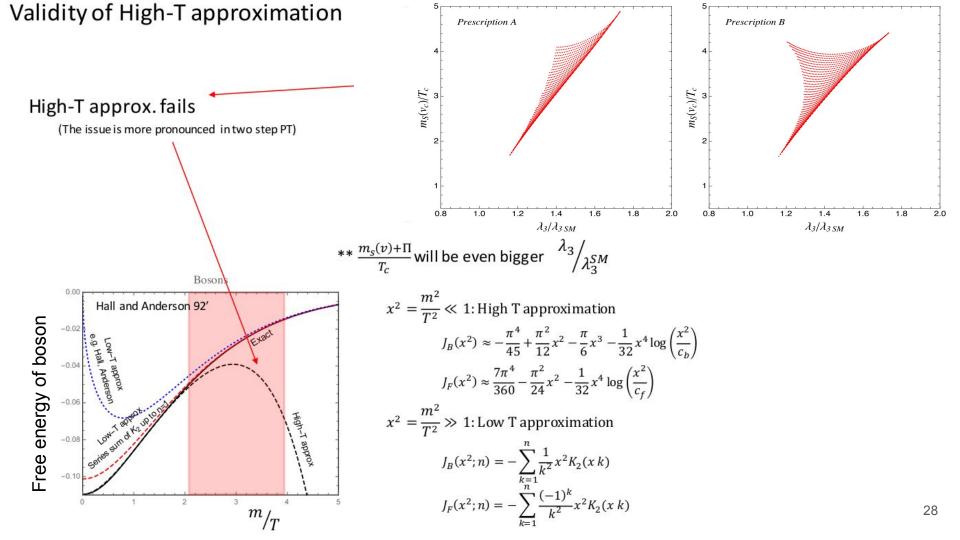
# Two step PT

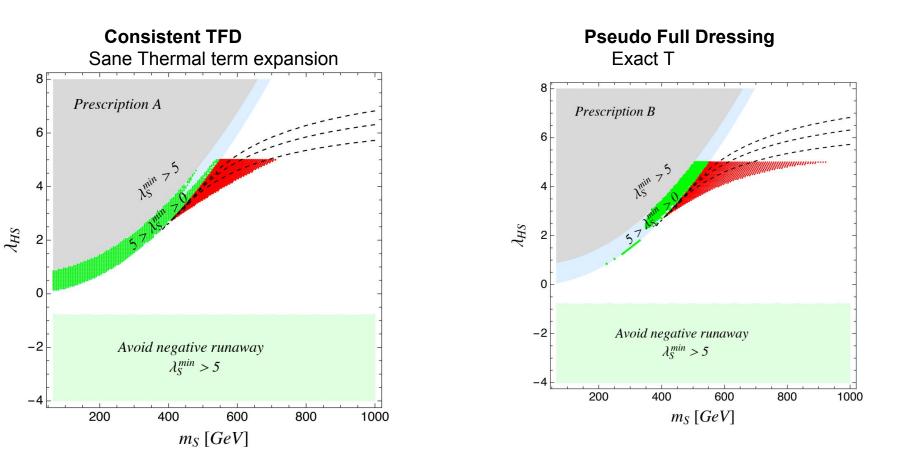


### Precision of the Higgs self couplings



- ✓ Depending on the criteria on  $v_c/T_c$ , the target precision can fluctuate by O(1) amout
- ~ 10% deviation of the Higgs
   Achievable @100 TeV pp collider and ILC

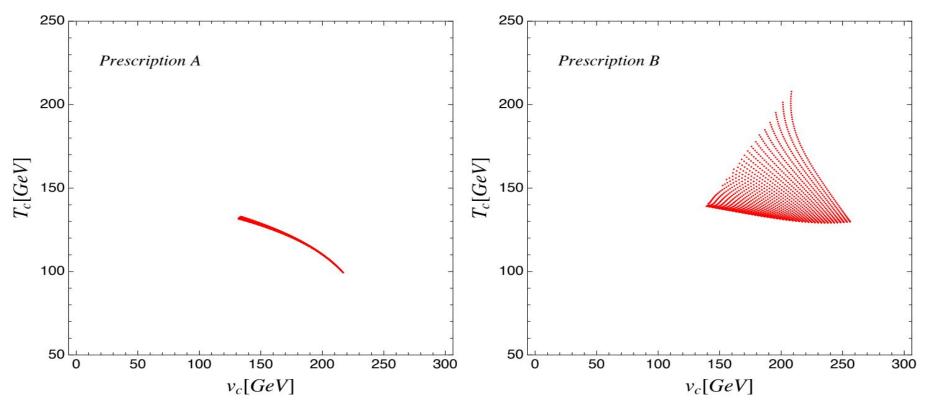




The results look similar in these plots

**Consistent TFD** 

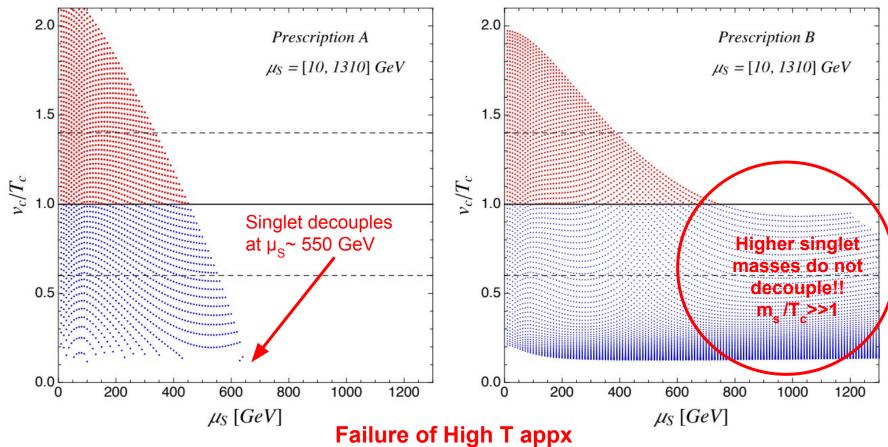
Pseudo Full Dressing



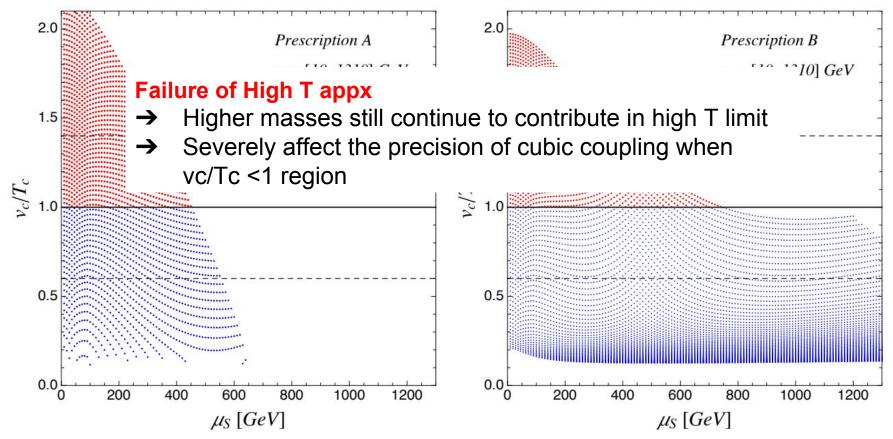
### ✓ More focused

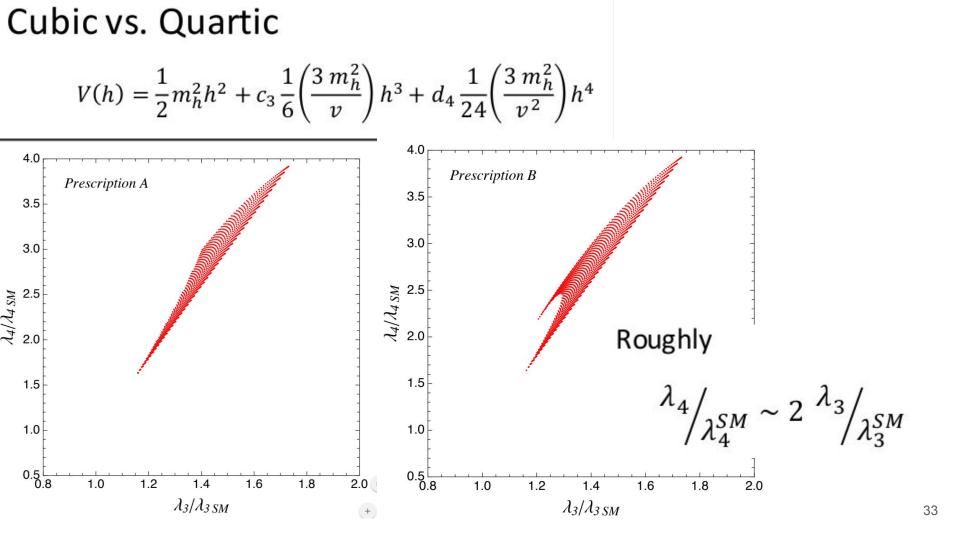
✓  $v_c < v$ =246 GeV is satisfied better

# Decoupling of higher singlet masses for one-step PT

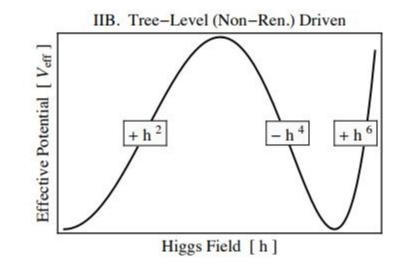


# Decoupling of higher singlet masses for one-step PT





# EFT - higher dimension operators



#### 1<sup>st</sup> order phase transition in EFT approach

E.g. dim-6 operator

$$\begin{split} \mathbf{V}_{EFT} &= m^2 \, |H|^2 + \lambda |H|^4 + \frac{c_6}{v^2} \frac{m_h^2}{2 \, v^2} \, |H|^6 \\ \mathbf{V}_{EFT} &= \frac{1}{2} \, m^2 h^2 + \frac{1}{4} \, \lambda h^4 + \frac{1}{8} \, \frac{c_6}{v^2} \frac{m_h^2}{2 \, v^2} h^6 \end{split}$$

$$0 = \frac{dV_{EFT}}{dh}\Big|_{h=v} = m^2 + \lambda h^2 + \frac{3c_6m_h^2}{8v^4}h^4\Big|_{h=v}$$
$$m_h^2 \equiv \frac{d^2V_{EFT}}{dh^2}\Big|_{h=v} = m^2 + 3\lambda h^2 + \frac{15}{8}c_6m_h^2$$

$$\lambda_{3} \equiv \frac{d^{3}V_{EFT}}{dh^{3}}\Big|_{h=\nu} = \frac{3m_{h}^{2}}{\nu}(1+c_{6})$$
$$\lambda_{4} \equiv \frac{d^{4}V_{EFT}}{dh^{4}}\Big|_{h=\nu} = \frac{3m_{h}^{2}}{\nu^{2}}(1+6c_{6})$$

$$\frac{\lambda_3}{\lambda_3^{SM}} = 1 + c_6$$
$$\frac{\lambda_4}{\lambda_4^{SM}} = 1 + 6 c_6$$

$$m^{2} = -\frac{m_{h}^{2}}{2} \left(1 - \frac{3}{4}c_{6}\right)$$
$$\lambda = \frac{m_{h}^{2}}{2\nu^{2}} \left(1 - \frac{3}{2}c_{6}\right)$$

#### 1<sup>st</sup> order phase transition in EFT approach

E.g. dim-6 operator

Nobel, Perelstein 08' Delaunay, Grojean, Wells, 08' Huang, Joglekar, Li, Wagner 15' Chung, Long, Wang 16' ....

$$V_{EFT} = \frac{1}{2}m^2h^2 + \frac{1}{4}\lambda h^4 + \frac{1}{8}\frac{c_6}{\nu^2}\frac{m_h^2}{2\nu^2}h^6$$

 $(m^2, \lambda, c_6)$  3 diff. local curvatures : 1<sup>st</sup> order PT becomes in principle possible

When keeping only  $T^2$ -term as thermal effect,  $m^2(T) = m^2 + aT^2$ , the analytic solution is possible

$$\frac{dV_{eff}}{dh}\Big|_{h=v_c} T=T_c} = 0 \qquad \text{Should be extreme point}$$

 $V(v_c, T_c) = V(0, T_c)$  Degeneracy of the vaccua

$$v_{c}^{2} = -\frac{4m^{2}(T_{c})}{\lambda} = -\frac{2\lambda v^{4}}{c_{6}m_{h}^{2}}$$

$$v_{c}^{2} < v^{2} & \& \lambda < 0$$

$$c_{6} = \frac{2}{3}\frac{1}{1 - \frac{2}{3}\frac{v_{c}^{2}}{v^{2}}}$$

$$\lambda = \frac{m_{h}^{2}}{2v^{2}}\left(1 - \frac{3}{2}c_{6}\right)$$

$$\frac{2}{3} < c_{6} < 2$$

#### Re-summed higher-dim operators

$$V_{tree} = -\mu^2 |H|^2 + \lambda |H|^4 + \sum_{n=1}^{\infty} \frac{c_{4+2n}}{v^{2n}} \frac{m_h^2}{2v^2} |H|^{4+2n}$$

$$c_{4+2n} = c (v/f)^{2n}$$
 with  $c \sim \mathcal{O}(1)$ 

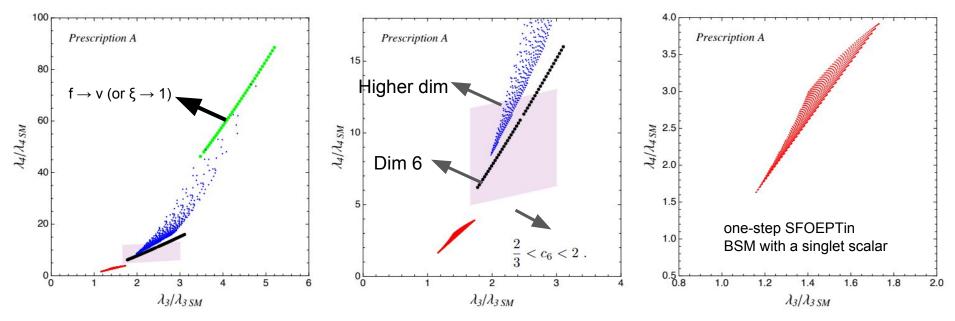
NDA scaling for all coeff

$$\begin{aligned} V_{tree} &= -\frac{\mu^2}{2}h^2 + \frac{\lambda}{4}h^4 + \frac{1}{8}\frac{c}{f^2}\frac{m_h^2}{2v^2}h^6\frac{1}{1 - \frac{h^2}{2f^2}} \\ \lambda_3 &= \frac{d^3V_{tree}(h)}{dh^3}\Big|_{h=v} = \frac{3m_h^2}{v}\left[1 + 16\,c\,\frac{\xi}{(2-\xi)^4}\right] \\ \lambda_4 &= \frac{d^4V_{tree}(h)}{dh^4}\Big|_{h=v} = \frac{3m_h^2}{v^2}\left[1 + 32\,c\,\frac{(6+\xi)\xi}{(2-\xi)^5}\right] \end{aligned}$$

Deviation in quartic is Bigger than cubic by  $2(6+\xi)/(2-\xi)$ 

# BSM Map in Higgs self coupling space

We wish to add more interesting BSM scenarios



→ HL-LHC has sensitivity 0f  $\lambda_3/\lambda_{3,SM}$  – 1 at 68%CL has two intervals, [–1.0, 1.8] U [3.5, 5.1]

 $\rightarrow$  Interval around SM can test O(1) deviations and exclude EFT cases

#### Lessons

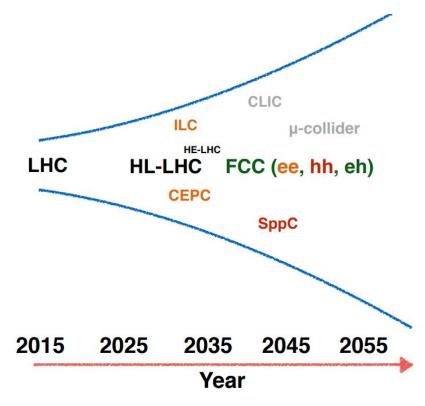
➔ Break-down of the high-temperature approximation is more pronounced in the two-step SFOEPT.

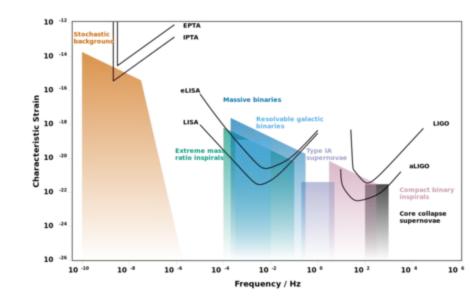
#### → Criteria for SFOEPT?

- v<sub>c</sub>/T<sub>c</sub> > 1 (1.4) requires the measurement of the coupling at ~ 15% (35%) precision achievable at ILC (via VBF process at higher c.o.m energy) and 100 TeV pp collider,
- more stronger criteria,  $v_c/T_c > 0.6$  requires ~ 5% precision of  $\lambda_3$  which is likely plausible only at 100 TeV pp collider
- → Various BSM scenarios can appear in different islands in ( $\lambda_3$ ,  $\lambda_4$ ) space
  - $\lambda_4$  could be important to distinguish different scenarios, and it has chance @ 100 TeV in case that a large deviation of  $\lambda_3$ , is observed
- → What if we observe any hint of Strong 1 st order Phase Transition?
  - Likely strongly coupled dynamics not far away from EW scale ?
- → Very strong EWPT could generate a stochastic background of gravitational waves signal is potentially within reach of future space-based gravitational wave interferometers, such as eLISA

Grojean et al' 99, Beniwal et al ' 17, Huang et. al. '17

#### The Road Ahead





**Future is bright!!** 

## Backup slides

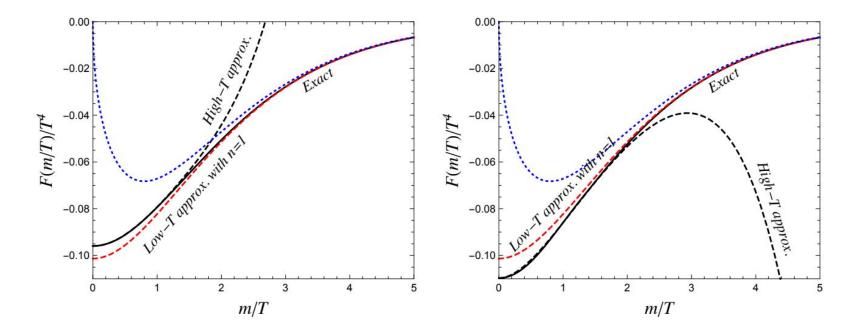
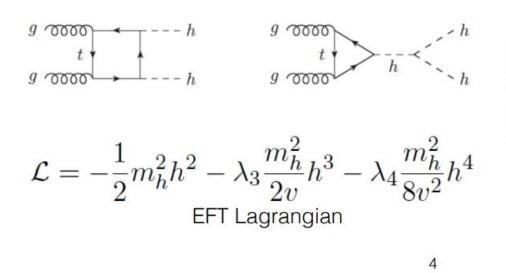


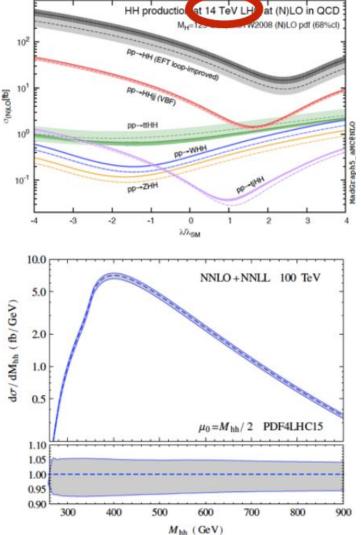
FIG. 1: Free energy for the fermions (left) and bosons (right) as a function of m/T in the high-T approximation (black-dashed), low-T approximation in Eq. (9) (red-dashed) with n = 1, and in the exact form (black-solid). The dotted-blue line is the low-T approximation with the approximated  $K_2$  as in [42].

### Simulation details - SM extension with the scalar singlet

- → For the one-step phase transition the quartic coupling  $\lambda_s$  does not play much role directly in the phenomenology apart from ensuring the stability of the potential at a large field.
- → Fix  $\lambda_s$  =0 and scan over  $\mu_s$  = [10, 1310] GeV (in steps of 10 GeV) and  $\lambda_{HS}$  = [0, 5] (in steps of 0.05).
- → For two-step cascade, the  $\lambda_s$  needs to stay above the minimum  $\lambda_{min s}$  so that (v, 0) remains the global minimum.
- → Scan intervals m<sub>S</sub> = [65, 700] GeV (in steps of 5 GeV) and  $\lambda_{HS}$  = [0, 5] (in steps of 0.05) for a few choices of  $\lambda_{S}$ , parameterized as  $\lambda_{S}$  =  $\lambda_{min S}$  + δS.
- → Assume: singlet mass is heavier than roughly mh/2 to avoid the Higgs decays to the singlet scalar.
- → Impose arbitrary hard cutoffs  $\lambda_{HS}$  < 5 and  $\lambda_{S}$  < 5 (smaller than 4 $\pi$  which is the typical unitarity bound), to avoid the strongly coupled regime.

- Probing triple-Higgs coupling with double Higgs production
  - Consistency of check of EWSB
  - Reconstructing the Higgs potential
  - Sensitivity through yields and kinematics
  - Large enhancement through BSM possible
  - Exhaustive program at the (HL-)LHC

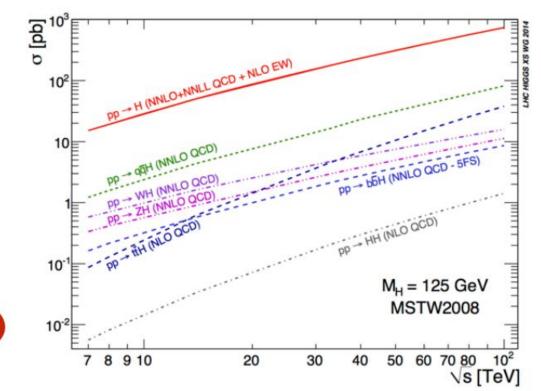




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# Dihiggs production

Process	8 TeV	14 TeV	100 TeV
gF	0.38	1	14.7
VBF	0.38	1	18.6
WH	0.43	1	9.7
ZH	0.47	1	12.5
ttH	0.21	1	61
bbH	0.34	1	15
gF to HH	0.24	1	42



Decay Channel	Branching Ratio	Total Yield ( $3000 \text{ fb}^{-1}$ )
$b\overline{b} + b\overline{b}$	33%	40,000
$b\overline{b} + W^+W^-$	25%	31,000
$b\overline{b} + \tau^+\tau^-$	7.3%	8,900
$ZZ + b\overline{b}$	3.1%	3,800
$W^+W^- + \tau^+\tau^-$	2.7%	3,300
$ZZ + W^+W^-$	1.1%	1,300
$\gamma\gamma + b\overline{b}$	0.26%	320
$\gamma\gamma + \gamma\gamma$	0.0010%	1.2

Table 1: Branching ratios for different HH final states, and their corresponding approximate expected yields in 3000 fb<sup>-1</sup> of data before any event selection is applied, assuming a total production cross section of 40.8 fb and  $m_H = 125$  GeV.

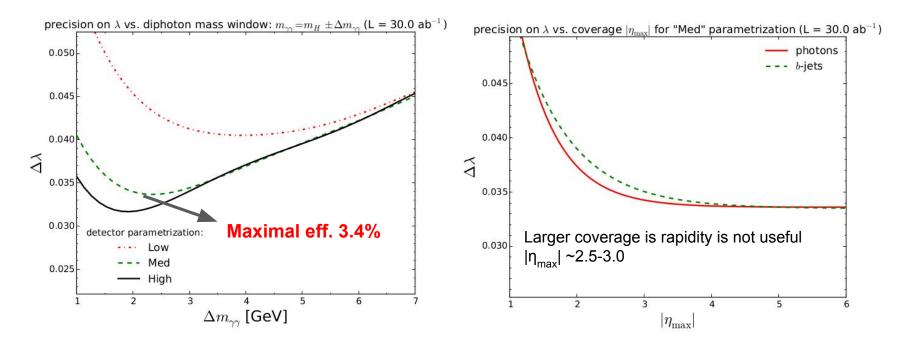
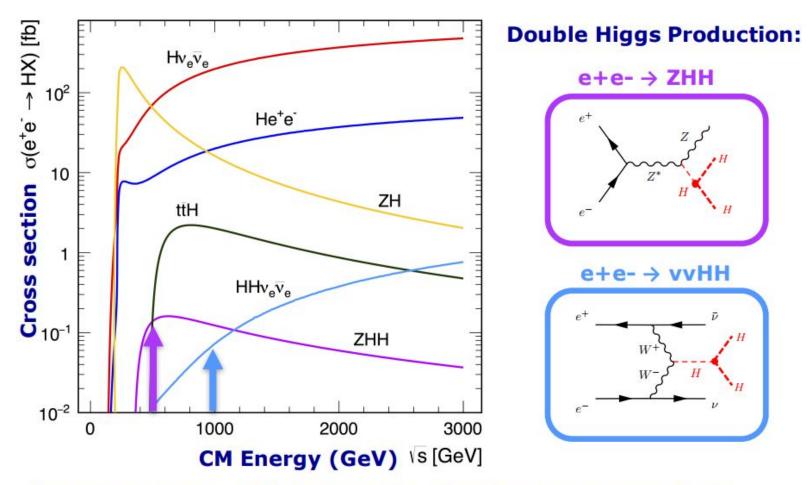


Fig. 65: Estimated precision on the measurement of the Higgs trilinear self-coupling. The left panel shows to result as a function of the cut on the invariant mass os the photon pair  $\Delta m_{\gamma\gamma}$  for the three detector benchmark scenarios, "Low" (dot-dashed red), "Medium" (dashed green) and "High" (solid black). In the right panel the result is shown as a function of the cut on the maximal rapidity of the reconstructed objects  $\eta_{max}$  assuming the "Medium" detector benchmark (the solid red and dashed green curves correspond to a variation of the photon and *b*-jets acceptances respectively). All the results have been obtained for an integrated luminosity of 30 ab<sup>-1</sup>.

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Require at least ~500 GeV for the direct measurement of the triple-Higgs coupling via double Higgs production.

# Baryogenesis ... in a nutshell

